Tracing the Pathway from Drift-Wave Turbulence with Broken Symmetry to the Production of Sheared Axial Mean Flow

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This study traces the emergence of sheared axial flow from collisional drift-wave turbulence with broken symmetry in a linear plasma device—the controlled shear decorrelation experiment. As the density profile steepens, the axial Reynolds stress develops and drives a radially sheared axial flow that is parallel to the magnetic field. Results show that the nondiffusive piece of the Reynolds stress is driven by the density gradient, results from spectral asymmetry of the turbulence, and, thus, is dynamical in origin. Taken together, these findings constitute the first simultaneous demonstration of the causal link between the density gradient, turbulence, and stress with broken spectral symmetry and the mean axial flow.

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Symmetry breaking plays a vital role in pattern formation, in particular, the generation of macroscopic flow by turbulence [1–3]. In stars and planetary atmospheres, convection-driven turbulence drives macroscopic flows of interest via Reynolds stresses. Broken symmetry defines the cross-phase (coherence) of the turbulent stresses and, thus, is fundamental to the flow structures. Examples of flow generation mechanisms exploiting broken symmetry include, but are not limited to, the anisotropic kinetic α effect [4] and Λ effect [5]. The need for such broken symmetry also appears in the magnetic dynamo, where turbulence with broken reflection symmetry (i.e., helicity) is required to produce a mean field dynamo [6].

The modelization of intrinsic macroscopic flow in plasmas also involves symmetry breaking in the turbulence. Intrinsic flow is of great importance in magnetic confinement fusion due to its promise of stabilizing magnetohydrodynamics instabilities [7] and reducing turbulent transport [8,9], particularly in burning plasma devices like ITER, where momentum input is limited. Measurements from Alcator C-Mod [10] indicate that intrinsic toroidal flow in the H mode is driven by edge ∇T . One possible mechanism [10,11] is that free energy stored in radial gradients is converted into shear flows via underlying turbulent stress, analogous to a heat engine process [12]. In this mechanism, broken symmetry in the spectra of driftwave turbulence, $\langle k_{\theta}k_{\tau}\rangle \neq 0$, induces a residual, nondiffusive component (Π_{rz}^{res}) in the Reynolds stress $(\langle \tilde{v}_r \tilde{v}_z \rangle =$ $-\chi_z \partial_r V_z + V_p V_z + \Pi_{rz}^{res}$ [11,13]. The divergence of this residual stress then defines an intrinsic torque that drives the macroscopic shear flow.

While many observations manifest the correlation between macroscopic intrinsic flow and edge profile gradients [9,14], investigations of the microscopic mechanism have been limited. Probe measurements from the edge of TJ-II stellarator [15] and TEXTOR tokamak [16] suggest that the nondiffusive residual stress $\Pi_{rz}^{\rm res}$ scales linearly with edge gradients. Parallel flow driven by turbulent Reynolds stress has been observed in a linear device, plasma assembly for nonlinear turbulence analysis [17]. A recent gyrokinetic simulation has predicted a dipole structure for $\Pi_{r_2}^{res}$, implying an intrinsic torque that is consistent with the measured rotation profile in DIII-D tokamak [18]. Despite of all these advances, our understanding of the underlying physics is still far from complete. Until now, there is no direct evidence linking the turbulence symmetry breaking to the development of residual stress. Moreover, it is still not clear if the residual stress can efficiently convert the thermodynamic free energy into the kinetic energy of the intrinsic flow.

In this Letter, we address the fundamental physics of how intrinsic flow develops in a confined plasma without magnetic shear. The results presented here constitute the first experimental evidence that macroscopic radially sheared parallel flows develop from a finite residual stress which emerges from a dynamical symmetry breaking in the spectral correlator $\langle k_z k_\theta \rangle$ of collisional drift-wave turbulence. In this case, the symmetry breaking [19] is *not* due to magnetic geometry but rather due to a phenomenon similar to modulational instability, in which a small seed axial shear is amplified by the turbulence. Note that modulational instability is also responsible for the production of zonal flows. However, zonal flow generation does not require broken symmetry in the turbulence spectra [20]. Thus, the mechanism for the generation of axial flow is more delicate than that for azimuthal flow.

The experiments have been conducted on the controlled shear decorrelation experiment (CSDX), a cylindrical plasma device with an overall length of 2.8 m and a diameter of 0.2 m [21,22]. The working gas is argon at a gas fill pressure of 1.8 mTorr. This relatively lower neutral pressure is used to avoid the volumetric recombination and detachment phenomena. The argon plasma is produced by a 13.56 MHz 1800 W rf helicon source via an m=1 antenna and is terminated by insulating (glass) end plates at both ends. The uniform magnetic field is in the axial direction (denoted as the \hat{z} direction) and is scanned from 500 to 1000 G in this study. The peak electron temperature is about 4 eV, and the peak ion temperature is about 0.5 eV. More details on this device can be found in previous publications [21,22].

Plasma velocity is measured using a combined Langmuir and Mach probe array. Two Mach probe tips, which are aligned along the axial direction, are used to measure the axial flow. The axial velocity is given by $v_z = Mc_s =$ $0.45c_s \ln [(J_u/J_d)]$, where $c_s = \sqrt{T_e/m_i}$ is the sound speed and $J_{u,d}$ are the ion saturation fluxes collected by Mach probe tips at the up- and downstream sides. The probe geometry is small enough to avoid shadowing effects that lead to a spuriously large inference of parallel velocities [23]. The mean parallel velocities found with this Mach probe are consistent with published measurements made using laser-induced fluorescence diagnostics [24]. The fluctuating $\mathbf{E} \times \mathbf{B}$ velocities are estimated from the floating potential gradients between two adjacent tips $(\nabla \tilde{\phi}_f)$, i.e., $\tilde{v}_r = -\nabla_{\theta} \tilde{\phi}_f / B$ and $\tilde{v}_{\theta} = \nabla_r \tilde{\phi}_f / B$. The distance between two adjacent floating potential tips is about 3 mm. The sampling rate of the probe data is $f_s =$ 500 kHz, which is well above the frequency of the observed dominant fluctuations (f < 30 kHz) [22]. Similar probe configurations have also been used in other studies on the structure of parallel flows [17].

In this study, we obtained different equilibrium profiles and fluctuation intensities by varying the B field. As shown in Fig. 1(a), when the B field is raised, the plasma density profile *steepens*. During the B scan, the variation in electron temperature profiles is negligible. The axial Reynolds stress $\langle \tilde{v}_z \tilde{v}_r \rangle$ is estimated using velocity fluctuations in the frequency range of 2 < f < 30 kHz. Previous studies have identified these as resistive drift-wave fluctuations [21]. $\langle \tilde{v}_z \tilde{v}_r \rangle$ is negligible for r < 3 cm at a lower B field but becomes substantially negative at a higher Bfield [Fig. 1(b)]. The Reynolds force $\mathcal{F}_{z}^{Re} = -\partial_{r} \langle \tilde{v}_{z} \tilde{v}_{r} \rangle$ [Fig. 1(e)] increases significantly in the core and becomes more negative at the edge (3 < r < 6 cm), and the radial shear of axial flow gets stronger as the B field increases [Fig. 1(b)]. The residual stress (described in more detail later) is computed from measured quantities using $\Pi_{r_7}^{res}$ = $\langle \tilde{v}_z \tilde{v}_r \rangle + \langle \tilde{v}_r^2 \rangle \tau_c \partial_r V_z$ [25], where τ_c is the eddy correlation time. Note that the momentum pinch (V_pV_z) vanishes due to the lack of toroidal effects in the CSDX. The magnitude of the resulting Π_{rz}^{res} also increases as the B field is raised [Fig. 1(c)]. At B = 800 G, the axial Reynolds force is much larger than the force on the ions arising from the

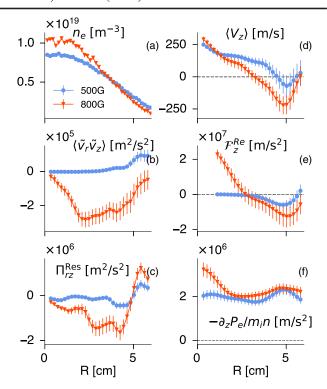


FIG. 1. Equilibrium profiles of (a) plasma density n_e , (b) parallel Reynolds stress $\langle \tilde{v}_r \tilde{v}_z \rangle$, (c) residual stress $\Pi_{rz}^{\rm res} = \langle \tilde{v}_z \tilde{v}_r \rangle + \langle \tilde{v}_r^2 \rangle \tau_c \partial_r V_z$, (d) axial velocity, (e) Reynolds force $\mathcal{F}_z^{\rm Re} = -\partial_r \langle \tilde{v}_z \tilde{v}_r \rangle$, and (f) the force on ions due to axial electron pressure gradient $-(\partial_z P_e/m_i n)$.

parallel electric field. Here, Boltzmann equilibrium is assumed, and the weak electric field is inferred from the measured electron pressure drop along the axial direction, $-\partial_z P_e/m_i n$ [Fig. 1(e)], which is measured by two Langmuir probes at up- and downstream locations ($\Delta z = 1.5$ m).

In order to determine if the observed changes in the turbulent stress are responsible for the observed increase in the sheared axial flow, an axial force balance analysis has also been performed. The axial ion momentum equation is written as

$$\frac{1}{r}\frac{\partial}{\partial r}(r\langle \tilde{v}_z \tilde{v}_r \rangle) = -\frac{1}{m_i \langle n \rangle} \frac{\partial P_e}{\partial z} - \nu_{\rm in} V_z + \frac{1}{r} \frac{\partial}{\partial r} \left(\mu_{ii} r \frac{\partial V_z}{\partial r} \right), \tag{1}$$

where the ion viscosity $\mu_{ii} = \frac{6}{5} \rho_i^2 \nu_{ii} \sim 5-10 \text{ m}^2/\text{s}$ and ionneutral collision frequency $\nu_{\text{in}} = n_{\text{gas}} v_{ti} \sigma_{\text{in}} \sim 3-6 \times 10^3 \text{ s}^{-1}$ are estimated using density and ion temperature profiles [24,26]. μ_{ii} and ν_{in} are likely to have small spatial variations, i.e., $\mu_{ii} \propto n T_i^{-1/2}$ and $\nu_{\text{in}} \propto T_i^{-1/2}$. Here, we assume the neutral pressure is radially uniform, and the neutral temperature is approximated by the ion temperature, which has been measured using LIF techniques [24]. This assumption gives the smallest estimate of neutral gas density depletion in the core and, thus, higher ion-neutral

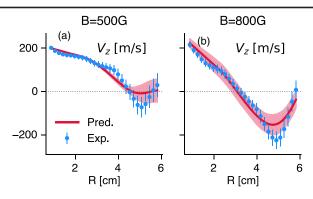


FIG. 2. Radial profiles of the mean axial velocity predicted by force balance analysis (solid lines) and measured Mach probe (circles) at 500 (a) and 800 G (b). The shaded area indicates the uncertainties of the predicted V_z profile.

frictional dissipation. But, in this experiment, the ionneutral drag dissipation is much smaller than the ion-ion collisional dissipation, and the force balance analysis is not sensitive to neutral profiles. A no-slip boundary condition is also imposed due to the observations in Fig. 1(d), i.e., $V_z \rightarrow 0$ at r = 6 cm. Taking the measured profiles of the Reynolds stress and the axial pressure gradient shown in Fig. 1, we can solve Eq. (1) for V_z using a finite difference method. As shown in Fig. 2, the calculated results (red curves) are in agreement with the mean axial ion flow profiles measured by the Mach probe (blue circles). This result shows that the weak shear flow found at 500 G is consistent with the weak axial equilibrium pressure gradient, while the stronger shear flow at 800 G is consistent with the observed turbulent stress. Similar comparisons have been carried out throughout the data set (B = 500-1000 G), and the normalized sum of squared errors between calculated and measured profiles, NSS = $\sum_r ((V_z^{\text{pred}} - V_z^{\text{exp}})^2 / \sigma_{V_z}^2(r))$, does not change much during the B field scan [Fig. 3(c)]. Therefore, an agreement between measured and calculated axial flow profiles is found across a range of B fields.

Using a steady-state shot-by-shot B field scan, we illustrate the link between ∇n , the turbulent flow drive, and the macroscopic intrinsic flow. The Reynolds power $\mathcal{P}_{z}^{\text{Re}} = -\langle V_{z} \rangle \partial_{r} \langle \tilde{v}_{z} \tilde{v}_{r} \rangle$ gives the rate of work performed by the turbulent fluctuations on the mean axial flow [27] at any point in the plasma. The volume-averaged Reynolds power $\mathcal{P}_{z}^{av} = \int -\langle V_{z} \rangle \partial_{r} \langle \tilde{v}_{z} \tilde{v}_{r} \rangle r dr / \int r dr$, where 1 < r < 5 cm, then gives the overall strength of the turbulent flow drive. Below a threshold value of $B \approx 650$ G, the turbulent flow drive is small, and the seed axial flow shear is driven by the axial pressure drop as shown earlier in Fig. 2(a) and varies at best weakly with ∇n . The magnitude of the axial flow shearing rate, $|V'_z| = |\partial_r V_z|$, then increases sharply when the density gradient exceeds a critical value, $\nabla n >$ 1.6×10^{20} m⁻⁴ [Fig. 3(a)] corresponding to $B \approx 650$ G. Further increases in ∇n associated with increased B then

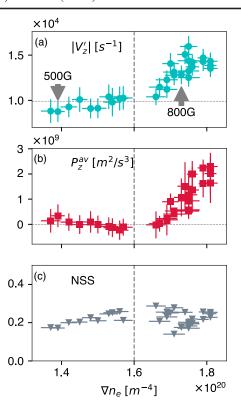


FIG. 3. The magnitude of axial flow shearing rate $|\partial_r V_z|$ (a), the volume-averaged axial Reynolds power $\mathcal{P}_z^{\mathrm{av}}$ (b), and normalized sum of squared errors (NSS) (c) plotted against the density gradient ∇n_e . Data sets are obtained by a shot-by-shot B field scan.

are associated with a strong increase in the flow shear and Reynolds power. These observations show that both the axial shear flow and its turbulent drive increase as ∇n increases.

The critical density gradient behavior shown in Fig. 3 has been reported previously [28], which is in agreement with numerical simulations of the coupled drift-wave turbulence–zonal flow system. Another possible mechanism regarding this transition is that the small plasma radius in the CSDX sets a lower bound for possible $k_{\perp}\rho_s$ values and thus prevents the growth of m=1 fluctuations at lower B fields. Raising the B field lessens this geometry constraint and allows the onset of stronger turbulence.

The above observations suggest that an increase in ∇n should drive a larger residual stress and thus a stronger intrinsic force, which then results in a larger sheared flow. Here, we derive a relation between the residual stress and the density gradient and account for the symmetry breaking in this relation. The fluctuating parallel ion flow evolves according to

$$\frac{d\tilde{v}_z}{dt} = -c_s^2 \nabla_z \left(\frac{e\tilde{\phi}}{T} + \frac{\tilde{P}}{P_0} \right) - \tilde{v}_r \frac{\partial V_z}{\partial r},$$

where c_s denotes the sound speed, \tilde{v}_r is the eddy radial velocity, \tilde{P} is the pressure fluctuation, and $\tilde{\phi}$ is the potential fluctuation. Also, $V_z \nabla_z \tilde{v}_z$ is neglected due to the small parallel wave number of the fluctuations $k_z \ll \omega/V_z$. For drift-wave turbulence with adiabatic electrons, one has $e\tilde{\phi}/T \sim \tilde{n}/n_0$, and $\tilde{P}/P_0 \sim \tilde{n}/n_0$ as temperature fluctuations are negligible in this experiment [21,22]. The axial flow fluctuation can then be written as $\tilde{v}_z \approx$ $-\sigma_{vT}(c_s^2\tau_c/L_z)(\tilde{n}/n_0) - \tilde{v}_r\tau_c(\partial V_z/\partial r)$. Here, τ_c is the eddy correlation time, L_z is the characteristic axial dimension, and σ_{vT} is a coefficient for acoustic coupling. Using a mixing length model for the density fluctuation, $\tilde{n}/n_0 \sim (l_c/n_0)|(\partial n_0/\partial r)|$, where $l_c \sim \tilde{v}_r \tau_c$ denotes the mixing length, one obtains $\tilde{v}_z \approx -\sigma_{vT}(c_s^2 l_c^2/L_z \tilde{v}_r)(1/n_0)$ $|(\partial n_0/\partial r)| - \tilde{v}_r \tau_c (\partial V_z/\partial r)$. After multiplying by \tilde{v}_r and taking an ensemble average, one obtains the expression for the total turbulent stress. This consists of two parts: a turbulent diffusive flux proportional to the velocity shear and a residual term driven by the density gradient,

$$\left\langle \tilde{v}_r \tilde{v}_z \right\rangle = - \chi_z \frac{\partial V_z}{\partial r} - \sigma_{vT} \frac{c_s^2 \langle l_c^2 \rangle}{L_z} \frac{1}{n_0} \left| \frac{\partial n_0}{\partial r} \right|.$$

Here $\chi_z \sim \langle \tilde{v}_r^2 \rangle \tau_c$ is the turbulent viscosity. The coefficient σ_{vT} accounts for the efficiency of the density gradient in driving the residual stress $\Pi_{rz}^{\rm res}$ via symmetry breaking. In particular, σ_{vT} accounts for the spectral correlation $\langle k_z k_\theta \rangle = \sum_{\bf k} k_z k_\theta |\hat{\phi}_{\bf k}|^2 / \sum_{\bf k} |\hat{\phi}_{\bf k}|^2$, which encodes the broken symmetry of the turbulence. Because all other terms can be measured in our experiment, σ_{vT} can be obtained by a least-squares fit.

The residual stress $\Pi_{rz}^{\rm res}$ was synthesized from the measured Reynolds stress and the diffusive stress inferred from experimental measurements; the result was shown earlier in Fig. 1(c). As shown in Fig. 4, at a smaller density gradient, the magnitude of residual stress, $|\Pi_{rz}^{\rm res}|$, is small and almost independent of the normalized density gradient. At larger ∇n , $|\Pi_{rz}^{\rm res}|$ increases in proportion to the normalized density gradient, with a slope of $\sigma_{vT}\approx 0.10$. Here, $|\Pi_{rz}^{\rm res}|$ is volume-averaged in the range of 1 < r < 5 cm. This finding strongly supports the hypothesis that the residual stress is driven by the density gradient when the gradient exceeds a critical value. The emergence of a finite $\sigma_{vT}\approx 0.1$ then indicates a symmetry-breaking mechanism that emerges at higher ∇n .

The development of residual stress requires symmetry breaking in **k** space [13], i.e., $\langle k_z k_\theta \rangle = \sum_{\mathbf{k}} k_z k_\theta |\hat{\phi}_{\mathbf{k}}|^2 / \sum_{\mathbf{k}} |\hat{\phi}_{\mathbf{k}}|^2 \neq 0$. The symmetry breaking can be assessed by investigating the joint probability density function (PDF) of radial and axial velocity fluctuations, $P(\tilde{v}_r, \tilde{v}_z)$. In the CSDX, we have $\tilde{v}_z \sim \nabla_{\parallel} \tilde{P} \sim k_z \tilde{\phi}$ and $\tilde{v}_r \sim k_\theta \tilde{\phi}$, due to the nearly adiabatic electron response and negligible temperature fluctuations. By normalizing the velocity fluctuations using their standard deviations, $P(\tilde{v}_r, \tilde{v}_z)$

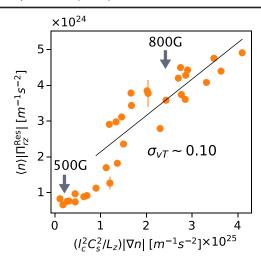


FIG. 4. Comparison between magnitudes of residual stress and normalized density gradient. The coefficient σ_{vT} is estimated to be about 0.10 by a least-squares fit.

represents the correlator $\langle k_z k_\theta \rangle$. As shown in Fig. 5, the anisotropy of $P(\tilde{v}_r, \tilde{v}_z)$ grows with increasing B field strength and ∇n . The highly elongated $P(\tilde{v}_r, \tilde{v}_z)$ at higher ∇n indicates increased asymmetry in $\langle k_z k_\theta \rangle$. Since a larger residual stress occurs at a higher ∇n , we can therefore infer that this symmetry breaking is related to the emergence of finite residual stress.

Conventional models of the origin of symmetry breaking rely on effects of magnetic geometry [13,29-32] and, therefore, are not applicable to zero magnetic shear cases, such as the CSDX. To address this question, a dynamical symmetry-breaking mechanism has been proposed [19]. This mechanism does not require magnetic shear and may also be relevant to intrinsic rotation in tokamaks with flat-q or weak shear. This model is derived from a collisional electron drift-wave system with axial momentum evolution. The mean axial flow shear then introduces a frequency shift proportional to $k_z k_\theta V_z'$ in the drift-wave growth rate. In our experiments, the seed axial flow shear is negative, $V_z' < 0$, because $V_{z}(r)$ is initially driven by the axial pressure drop and decreases from the core to the edge. As a result, modes with $\langle k_z k_\theta \rangle < 0$ grow faster than modes with $\langle k_z k_\theta \rangle > 0$, leading to spectral imbalance, with a predominance of the

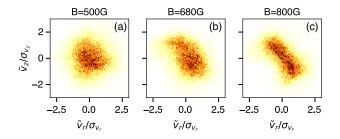


FIG. 5. Joint PDF of radial and axial velocity fluctuations, $P(\tilde{v}_r, \tilde{v}_z)$, at different magnetic fields at $r \approx 3$ cm. Normalization is the standard deviation.

spectral intensity in quadrants II and IV of the $k_{\theta}-k_{z}$ plane. The predicted spectral imbalance $\langle k_{\theta}k_{z}\rangle < 0$ is consistent with the tilted contour of $P(\tilde{v}_{r}, \tilde{v}_{z})$ [Fig. 5(c)]. As demonstrated in Ref. [19], the spectral asymmetry results in a residual stress of the form $-\chi_{z}^{\rm res}V'_{z}$, with $\chi_{z}^{\rm res}<0$, i.e., a negative-definite contribution to the total viscosity [i.e., $\langle \tilde{v}_{r}\tilde{v}_{z}\rangle = -\chi_{z}V'_{z} + \Pi_{rz}^{\rm res} = (-\chi_{z} + |\chi_{z}^{\rm res}|)V'_{z}$]. Then, $|\chi_{z}^{\rm res}| = \chi_{z}$ defines the threshold $\nabla n_{0}/n_{0}$ for the onset of axial flow generation. Using Eq. (36) of Ref. [19] for $|\chi_{z}^{\rm res}|$, it gives $\nabla n_{\rm crit} = (n_{0}\alpha\omega_{*e}^{2}/\langle k_{\theta}k_{z}\rangle\rho_{s}c_{s}) \times (L_{z}/c_{s}^{2}\tau_{c}) \sim 1.5 \times 10^{20}~{\rm m}^{-4}$ in agreement with the experiment. Here, $\alpha = k_{z}^{2}v_{te}^{2}/\omega_{*e}\nu_{ei} \sim 1$ is the adiabaticity factor, the perpendicular turbulence scale length is $k_{\theta}\rho_{s} \sim 1.5$, the eddy correlation time is $\tau_{c} \sim 6 \times 10^{-5}$ s, and $\sigma_{vT} = \langle k_{\theta}k_{z}\rangle/\langle k_{\theta}^{2}\rangle \sim 0.1$.

In summary, in this study, detailed measurements of axial flows and turbulent Reynolds stresses have been performed in cylindrical plasmas without magnetic shear. As the density profile steepens, Reynolds stress develops and, in turn, drives a sheared mean axial flow. Both the axial flow shearing rate and the turbulent Reynolds power increase with the density gradient. The magnitude of residual stress also scales with the density gradient. $P(\tilde{v}_r, \tilde{v}_z)$ becomes highly tilted and anisotropic at higher ∇n , indicating an asymmetry in the spectral correlator $\langle k_{\theta}k_{\tau}\rangle$. This symmetry breaking in **k** space implies a finite residual stress observed at a higher ∇n and is consistent with a model of dynamical symmetry breaking in the turbulence. These findings constitute the first demonstration of the causal link of spectral symmetry breaking in drift-wave turbulence to the development of a nondiffusive, residual stress and ultimately to the onset of intrinsic axial shear flow.

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- V. Shtern and F. Hussain, Annu. Rev. Fluid Mech. 31, 537 (1999).
- [2] P. H. Diamond, Y.-M. Liang, B. A. Carreras, and P. W. Terry, Phys. Rev. Lett. 72, 2565 (1994).
- [3] K.-T. Wu, J. B. Hishamunda, D. T. N. Chen, S. J. DeCamp, Y.-W. Chang, A. Fernández-Nieves, S. Fraden, and Z. Dogic, Science 355, eaal1979 (2017).
- [4] U. Frisch, Z. She, and P. Sulem, Physica (Amsterdam) 28D, 382 (1987).
- [5] G. Rüdiger, Differential Rotation and Stellar Convection: Sun and Solar-Type Stars (Taylor & Francis, London, 1989), Vol. 5.
- [6] H. K. Moffatt, Field Generation in Electrically Conducting Fluids (Cambridge University Press, Cambridge, England, 1978).

- [7] A. M. Garofalo, E. J. Strait, L. C. Johnson, R. J. La Haye, E. A. Lazarus, G. A. Navratil, M. Okabayashi, J. T. Scoville, T. S. Taylor, and A. D. Turnbull, Phys. Rev. Lett. 89, 235001 (2002).
- [8] M. Yoshida, Y. Kamada, H. Takenaga, Y. Sakamoto, H. Urano, N. Oyama, and G. Matsunaga (the JT-60 Team), Phys. Rev. Lett. 100, 105002 (2008).
- [9] J. E. Rice, Plasma Phys. Controlled Fusion 58, 083001 (2016).
- [10] J. E. Rice, J. W. Hughes, P. H. Diamond, Y. Kosuga, Y. A. Podpaly, M. L. Reinke, M. J. Greenwald, O. D. Gürcan, T. S. Hahm, A. E. Hubbard, E. S. Marmar, C. J. McDevitt, and D. G. Whyte, Phys. Rev. Lett. 106, 215001 (2011).
- [11] Y. Kosuga, P. H. Diamond, and O. D. Gurcan, Phys. Plasmas 17, 102313 (2010).
- [12] H. Ozawa, A. Ohmura, R. D. Lorenz, and T. Pujol, Rev. Geophys. 41, 1018 (2003).
- [13] P. Diamond, Y. Kosuga, O. Gürcan, C. McDevitt, T. Hahm, N. Fedorczak, J. Rice, W. Wang, S. Ku, J. Kwon, G. Dif-Pradalier, J. Abiteboul, L. Wang, W. Ko, Y. Shi, K. Ida, W. Solomon, H. Jhang, S. Kim, S. Yi *et al.*, Nucl. Fusion **53**, 104019 (2013).
- [14] K. Ida and J. Rice, Nucl. Fusion **54**, 045001 (2014).
- [15] B. Gonçalves, C. Hidalgo, M. A. Pedrosa, R. O. Orozco, E. Sánchez, and C. Silva, Phys. Rev. Lett. **96**, 145001 (2006).
- [16] Y. Xu, C. Hidalgo, I. Shesterikov, M. Berte, P. Dumortier, M. V. Schoor, M. Vergote, A. Krämer-Flecken, and R. Koslowski (the TEXTOR Team), Nucl. Fusion 53, 072001 (2013).
- [17] S. Inagaki, T. Kobayashi, Y. Kosuga, S. I. Itoh, T. Mitsuzono, Y. Nagashima, H. Arakawa, T. Yamada, Y. Miwa, N. Kasuya, M. Sasaki, M. Lesur, A. Fujisawa, and K. Itoh, Sci. Rep. 6, 22189 (2016).
- [18] W. X. Wang, B. A. Grierson, S. Ethier, J. Chen, E. Startsev, and P. H. Diamond, Phys. Plasmas **24**, 092501 (2017).
- [19] J. C. Li, P. H. Diamond, X. Q. Xu, and G. R. Tynan, Phys. Plasmas 23, 052311 (2016).
- [20] P. H. Diamond, S.-I. Itoh, K. Itoh, and T. S. Hahm, Plasma Phys. Controlled Fusion 47, R35 (2005).
- [21] M. J. Burin, G. R. Tynan, G. Y. Antar, N. A. Crocker, and C. Holland, Phys. Plasmas 12, 052320 (2005).
- [22] S. C. Thakur, C. Brandt, L. Cui, J. J. Gosselin, A. D. Light, and G. R. Tynan, Plasma Sources Sci. Technol. 23, 044006 (2014).
- [23] J. J. Gosselin, S. C. Thakur, S. H. Sears, J. S. McKee, E. E. Scime, and G. R. Tynan, Phys. Plasmas 23, 073519 (2016).
- [24] S. C. Thakur, J. J. Gosselin, J. McKee, E. E. Scime, S. H. Sears, and G. R. Tynan, Phys. Plasmas **23**, 082112 (2016).
- [25] Z. Yan, M. Xu, P. H. Diamond, C. Holland, S. H. Müller, G. R. Tynan, and J. H. Yu, Phys. Rev. Lett. 104, 065002 (2010).
- [26] C. Holland, J. H. Yu, A. James, D. Nishijima, M. Shimada, N. Taheri, and G. R. Tynan, Phys. Rev. Lett. **96**, 195002 (2006).
- [27] G. R. Tynan, I. Cziegler, P. H. Diamond, M. Malkov, A. Hubbard, J. W. Hughes, J. L. Terry, and J. H. Irby, Plasma Phys. Controlled Fusion **58**, 044003 (2016).
- [28] Z. Yan, G. R. Tynan, C. Holland, M. Xu, S. H. Müller, and J. H. Yu, Phys. Plasmas 17, 012302 (2010).
- [29] O. D. Gürcan, P. H. Diamond, T. S. Hahm, and R. Singh, Phys. Plasmas **14**, 042306 (2007).

- [30] X. Garbet, J. Abiteboul, E. Trier, O. Gürcan, Y. Sarazin, A. Smolyakov, S. Allfrey, C. Bourdelle, C. Fenzi, V. Grandgirard, P. Ghendrih, and P. Hennequin, Phys. Plasmas 17, 072505 (2010).
- [31] A. Peeters, C. Angioni, A. Bortolon, Y. Camenen, F. Casson, B. Duval, L. Fiederspiel, W. Hornsby, Y. Idomura,
- T. Hein, N. Kluy, P. Mantica, F. Parra, A. Snodin, G. Szepesi, D. Strintzi, T. Tala, G. Tardini, P. de Vries, and J. Weiland, Nucl. Fusion **51**, 094027 (2011).
- [32] C. Angioni, Y. Camenen, F. Casson, E. Fable, R. McDermott, A. Peeters, and J. Rice, Nucl. Fusion **52**, 114003 (2012).