## PHYSICS 200B : CLASSICAL MECHANICS SOLUTION SET \#3

(1) Create your own pixelated image to iterate under the cat map. You can also find many interesting images over the web. (Nothing pornographic, please! ${ }^{1}$ ) Iterate the pixel coordinates under the cat map. Show how your image gets scrambled after a few iterations of the map, but is nevertheless recurrent. You'll need to write a computer code to do this problem.

Solution: Here is the matlab code I wrote:

```
function cat_map(filename,num_iter)
orig_fig=importdata(filename);
    %import the figure, which ideally should be square (n by n)
len=length(orig_fig);
[p,q]=meshgrid(1:len); %generate the meshgrid
q_prime=q+p;
p_prime=q+2*p;
q_prime=mod(q_prime-1,len)+1;p_prime=mod(p_prime-1,len)+1; %the cat map
%q_prime is the new x coordinate; p_prime is the new y coordinate
linear_indx=sub2ind([len,len],p_prime,q_prime);
linear_indx=linear_indx(:);
old_fig=orig_fig;
old_fig=reshape(old_fig,[len*len,3]);
for i_iter=1:num_iter
    new_fig=old_fig(linear_indx,:);
    old_fig=new_fig;
end
%figure;
new_fig=reshape(new_fig,[len,len,3]);
imagesc(new_fig);
axis equal
axis off
xlim([1,len]);
ylim([1,len]);
```

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Figure 1: The cat map on a piglet.

The argument 'num_iter' is the number of the cat map you would like to perform on your cool figure. Here is mine.
(2) Numerically integrate the system

$$
\begin{aligned}
& \dot{r}=r\left(1-r^{2}\right)+\lambda r \cos \theta \\
& \dot{\theta}=1
\end{aligned}
$$

with $0<\lambda<1$, and show that any initial condition lying between the concentric circles of radii $\sqrt{1 \pm \lambda}$ approaches a closed limit cycle in the long time limit. Choose whatever value of $\lambda$ suits your taste.

Solution: One can use the function ode45 to integrate the differential equation above with certain initial condition. Without loss of generality, $\theta(0)$ is set to be zero. The time evolution of $r$ with various initial condition is shown in Fig. 2


Figure 2:
(3) Consider the equation

$$
\ddot{x}+x=\epsilon x^{5}
$$

with $\epsilon \ll 1$.
(a) Develop a two term straightforward expansion for the solution and discuss its uniformity.
(b) Using the Poincaré-Lindstedt method, find a uniformly valid expansion to first order.
(c) Using the multiple time scale method, find a uniformly valid expansion to first order.

Solution: We can formally expand the solution as:

$$
\begin{equation*}
x(t)=x_{0}(t)+\epsilon x_{1}(t)+\epsilon^{2} x_{2}(t)+\ldots \tag{1}
\end{equation*}
$$

Plugging the expression into differentiate equation above, we obtain the equation for each order by matching the power of $\epsilon$. The zeroth order simply describes harmonic oscillators:

$$
\begin{equation*}
\ddot{x}_{0}+x_{0}=0 \tag{2}
\end{equation*}
$$

The solution is $x_{0}(t)=A \cos (t+\phi)$. The second order equation is:

$$
\begin{equation*}
\ddot{x}_{1}+x_{1}=x_{0}^{5} \tag{3}
\end{equation*}
$$

The time dependence of right hand side (R.H.S) is known from the zeroth order solution:

$$
\begin{equation*}
x_{0}^{5}=\frac{A^{5}}{16}(10 \cos (t+\phi)+5 \cos (3 t+3 \phi)+\cos (5 t+5 \phi)) \tag{4}
\end{equation*}
$$

The first term causes resonance in this order, and invalids the simple expansion.
Now we apply the Poincaré-Lindstedt method. We define $s=\Omega t$, then the Hamiltonian becomes:

$$
\begin{equation*}
\Omega^{2} \frac{\partial^{2} x}{\partial s^{2}}+x=\epsilon x^{5} \tag{5}
\end{equation*}
$$

Besides the expansion of $x$, we also need expand $\Omega^{2}$ as $1+\epsilon a_{1}+\epsilon^{2} a_{2}+\ldots$. Again we separate the equations by matching different orders of $\epsilon$. The zeroth order is unchanged. The first order becomes:

$$
\begin{equation*}
\frac{\partial^{2} x_{1}}{\partial s^{2}}+x_{1}=x_{0}^{5}-a_{1} \frac{\partial^{2} x_{0}}{\partial s^{2}} \tag{6}
\end{equation*}
$$

Plugging in the zeroth order solution, the R.H.S reads:

$$
\begin{equation*}
x_{0}^{5}=\left(\frac{5}{8} A^{5}+a_{1} A\right) \cos (s+\phi)+\frac{5}{16} A^{5} \cos (3 s+3 \phi)+\frac{1}{16} A^{5} \cos (5 s+5 \phi) \tag{7}
\end{equation*}
$$

The resonant term can be removed by setting $a_{1}=-\frac{5}{8} A^{4}$. Then the first order solution is:

$$
\begin{equation*}
x_{1}(s)=-\frac{A^{5}}{384}(15 \cos (3 s+3 \phi)+\cos (5 s+5 \phi)) \tag{8}
\end{equation*}
$$

Therefore, up to first order, we have

$$
\begin{equation*}
x(s)=A \cos (s+\phi)-\epsilon \frac{A^{5}}{384}(15 \cos (3 s+3 \phi)+\cos (5 s+5 \phi)) \tag{9}
\end{equation*}
$$

where $s=\sqrt{1-\frac{5}{8} A^{4} \epsilon} t$.
Next, we apply the multiple time scale method to this problem. We define $\frac{\partial}{\partial t}=\frac{\partial}{\partial T_{0}}+\epsilon \frac{\partial}{\partial T_{1}} \ldots$, along with the usual expansion of $x(t)$. The zeroth order equation is:

$$
\begin{equation*}
\frac{\partial^{2} x_{0}}{\partial T_{0}^{2}}+x_{0}=0 \tag{10}
\end{equation*}
$$

The solution is $x_{0}=A \cos \left(T_{0}+\phi\right)$. For simplicity, we set $\theta$ to be $T_{0}+\phi$. Then we plug it into the first order equation:

$$
\begin{equation*}
\frac{\partial^{2} x_{1}}{\partial \theta^{2}}+x_{1}=A^{5} \cos ^{5} \theta+2 \frac{\partial}{\partial T_{1}} A \sin \theta \tag{11}
\end{equation*}
$$

The R.H.S reads:

$$
\begin{equation*}
\frac{5}{8} A^{5} \cos \theta+\frac{5}{16} A^{5} \cos 3 \theta+\frac{1}{16} A^{5} \cos 5 \theta+2 \frac{\partial}{\partial T_{1}} A \sin \theta \tag{12}
\end{equation*}
$$

In order to remove the resonant term, we demand:

$$
\begin{align*}
& \frac{\partial}{\partial T_{1}} A=0  \tag{13}\\
& \frac{\partial}{\partial T_{1}} \theta=-\frac{5}{16} A^{4}
\end{align*}
$$

Therefore $\theta=T_{0}-\frac{5}{16} A^{4} T_{1}+\theta_{0}=\left(1-\frac{5}{16} \epsilon A^{4}\right) t+\theta_{0}$. The solution of Eq. 17 is :

$$
\begin{equation*}
x_{1}(\theta)=-\frac{A^{5}}{384}(15 \cos 3 \theta+\cos 5 \theta) \tag{14}
\end{equation*}
$$

and the full solution is:

$$
\begin{equation*}
x(\theta)=A \cos \theta-\epsilon \frac{A^{5}}{384}(15 \cos 3 \theta+\cos 5 \theta) \tag{15}
\end{equation*}
$$

with $\theta=\left(1-\frac{5}{16} \epsilon A^{4}\right) t+\theta_{0}$. This is consistent with what we got using he Poincaré-Lindstedt method.
(4) Consider the equation

$$
\ddot{x}+\epsilon \dot{x}^{3}+x=0
$$

with $\epsilon \ll 1$. Using the multiple time scale method, find a uniformly valid expansion to first order.

We apply the multiple time scale method to this problem. We define $\frac{\partial}{\partial t}=\frac{\partial}{\partial T_{0}}+\epsilon \frac{\partial}{\partial T_{1}} \ldots$, along with the usual expansion of $x(t)$. The zeroth order equation is:

$$
\begin{equation*}
\frac{\partial^{2} x_{0}}{\partial T_{0}^{2}}+x_{0}=0 \tag{16}
\end{equation*}
$$

The solution is $x_{0}=A \cos \left(T_{0}+\phi\right)$. For simplicity, we set $\theta$ to be $T_{0}+\phi$. Then we plug it into the first order equation:

$$
\begin{equation*}
\frac{\partial^{2} x_{1}}{\partial \theta^{2}}+x_{1}=A^{3} \sin ^{3} \theta+2 \frac{\partial}{\partial T_{1}} A \sin \theta \tag{17}
\end{equation*}
$$

The R.H.S reads:

$$
\begin{equation*}
\frac{3}{4} A^{3} \sin \theta-\frac{1}{4} A^{3} \sin 3 \theta+2 \frac{\partial}{\partial T_{1}} A \sin \theta \tag{18}
\end{equation*}
$$

In order to remove the resonant term, we demand:

$$
\begin{align*}
& \frac{\partial}{\partial T_{1}} A=-\frac{3}{8} A^{3} \\
& \frac{\partial}{\partial T_{1}} \theta=0 \tag{19}
\end{align*}
$$

Therefore $\frac{1}{A^{2}}=\frac{3}{4} T_{1}+A_{0}=\frac{3}{4} \epsilon t+A_{0}$, where $A_{0}$ is an integrating constant, and $\theta=T_{0}+\theta_{0}=$ $t+\theta_{0}$. The solution of the first order equation is :

$$
\begin{equation*}
x_{1}(\theta)=\frac{1}{32} A^{3} \sin 3 \theta \tag{20}
\end{equation*}
$$

and the full solution is:

$$
\begin{equation*}
x(\theta)=A \cos \theta-\epsilon \frac{A^{3}}{32} \sin 3 \theta \tag{21}
\end{equation*}
$$

with $A=\left(\frac{3}{4} \epsilon t+A_{0}\right)^{-1 / 2}$ and $\theta=t+\theta_{0}$.
(5) Analyze the forced oscillator

$$
\ddot{x}+x=\epsilon\left(\dot{x}-\frac{1}{3} \dot{x}^{3}\right)+\epsilon f_{0} \cos (t+\epsilon \nu t)
$$

using the discussion in $\S 3.3 .1$ and $\S 3.3 .2$ of the notes as a template.
Solution: We still apply the multiple time scale method. The zeroth order solution again is $x_{0}=A \cos \left(T_{0}+\phi\right)$. The first order equation is:

$$
\begin{equation*}
\frac{\partial^{2} x_{1}}{\partial T_{0}^{2}}+x_{1}=\left(\frac{A^{3}}{4}-A\right) \sin \left(T_{0}+\phi\right)-\frac{A^{3}}{12} \sin \left(3 T_{0}+3 \phi\right)+f_{0} \cos \left(T_{0}+\nu T_{1}\right)+2 \frac{\partial}{\partial T_{1}} A \sin \left(T_{0}+\phi\right) \tag{22}
\end{equation*}
$$

In order to eliminate all the terms with frequency 1 , we demand:

$$
\begin{align*}
& \left(\frac{A^{3}}{4}-A\right) \cos \phi-f_{0} \sin \left(\nu T_{1}\right)+2 \frac{\partial A}{\partial T_{1}} \cos \phi-2 A \frac{\partial \phi}{\partial T_{1}} \sin \phi=0  \tag{23}\\
& \left(\frac{A^{3}}{4}-A\right) \sin \phi+f_{0} \cos \left(\nu T_{1}\right)+2 \frac{\partial A}{\partial T_{1}} \sin \phi+2 A \frac{\partial \phi}{\partial T_{1}} \cos \phi=0
\end{align*}
$$

Now we seek the fixed point solution $\left(A\left(T_{1}\right)=A, \phi\left(T_{1}\right)=\nu T_{1}+\phi_{0}\right)$ (Here I've used a more symmetric convention compared with the one in the lecture notes). The above equation can be organized as:

$$
\begin{equation*}
\left(\frac{A^{3}}{4}-A+2 i A \nu\right) e^{i \phi_{0}}+i f_{0}=0 \tag{24}
\end{equation*}
$$

Therefore, $A$ is given by the real solutions of the following equation:

$$
\begin{equation*}
\left(\frac{A^{3}}{4}-A\right)^{2}+4 A^{2} \nu^{2}=f_{0}^{2} \tag{25}
\end{equation*}
$$

Once $A$ is obtained, one can find $\phi_{0}$ easily.
The root structure is determined by the following polynomial:

$$
\begin{equation*}
y^{3}-8 y^{2}+16\left(1+4 \nu^{2}\right) y-16 f_{0}^{2}=0 \tag{26}
\end{equation*}
$$

The two extrema are $y=\frac{4}{3}\left(2 \mp \sqrt{1-12 \nu^{2}}\right)$ with the corresponding values:

$$
\begin{equation*}
g^{ \pm}=-16 f_{0}^{2}+\frac{128}{27}\left(1 \pm \sqrt{1-12 \nu^{2}}+12 \nu^{2}\left(3 \mp \sqrt{1-12 \nu^{2}}\right)\right) \tag{27}
\end{equation*}
$$

If $g^{+}>0$ and $g^{-}<0$, then there are three real solutions. The number of real solutions of A depends on $f_{0}$ and $\nu$ and is illustrated in Fig. 3(a).

Next, we analyze the stability of these solutions. Adding the second equation to the first equation in Eq. 23 with an factor of $i$, we obtain:

$$
\begin{equation*}
\left(\frac{4}{3} A^{3}-A\right) e^{i \phi_{0}}+i f_{0}+2 \frac{\partial A}{\partial T_{1}} e^{i \phi_{0}}+i 2 A \nu e^{i \phi_{0}}+i A \frac{\partial \phi_{0}}{\partial T_{1}} e^{i \phi_{0}}=0 \tag{28}
\end{equation*}
$$

After linearization around the fixed points, the equation can be simplified to

$$
\frac{\partial}{\partial T_{1}}\binom{\delta A}{\delta \phi_{0}}=\left(\begin{array}{cc}
\frac{1}{2}-\frac{3}{8} A^{2} & A \nu  \tag{29}\\
-\frac{\nu}{A}, & \frac{1}{2}-\frac{A^{2}}{8}
\end{array}\right)\binom{\delta A}{\delta \phi_{0}}
$$

The matrix has two eigenvalues:

$$
\begin{equation*}
E=\frac{1}{8}\left(4-2 A^{2} \pm \sqrt{A^{4}-64 \nu^{2}}\right) \tag{30}
\end{equation*}
$$

from which the property of the attractors can be determined straightforward: the curve $A^{4}-64 \nu^{2}=0$ separates spiral region and node region. Within the spiral region, the curve $4-2 A^{2}$ separates stable spiral region and unstable spiral region. Within the node region, if both of the eigenvalues are larger than zero, the attractor is a unstable node; if both of the eigenvalues are less than zero, the attractor is a stable node; otherwise the attractor is a saddle point. The phase diagram is plotted in Fig. 3(b).


Figure 3: (a) The number of solutions that $A$ can take. (b) The phase diagram of the attractors.


[^0]:    ${ }^{1}$ Well, I suppose animal sex is OK, if you must.

