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## Chapter 2

## Response, Resonance, and the Electron Gas

### 2.1 Response and Resonance

Consider a damped harmonic oscillator subjected to a time-dependent forcing:

$$
\begin{equation*}
\ddot{x}+2 \gamma \dot{x}+\omega_{0}^{2} x=f(t), \tag{2.1}
\end{equation*}
$$

where $\gamma$ is the damping rate $(\gamma>0)$ and $\omega_{0}$ is the natural frequency in the absence of damping ${ }^{1}$. We adopt the following convention for the Fourier transform of a function $H(t)$ :

$$
\begin{align*}
& H(t)=\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi} \hat{H}(\omega) e^{-i \omega t}  \tag{2.2}\\
& \hat{H}(\omega)=\int_{-\infty}^{\infty} d t H(t) e^{+i \omega t} \tag{2.3}
\end{align*}
$$

Note that if $H(t)$ is a real function, then $\hat{H}(-\omega)=\hat{H}^{*}(\omega)$. In Fourier space, then, eqn. (2.1) becomes

$$
\begin{equation*}
\left(\omega_{0}^{2}-2 i \gamma \omega-\omega^{2}\right) \hat{\boldsymbol{x}}(\omega)=\hat{f}(\omega), \tag{2.4}
\end{equation*}
$$

with the solution

$$
\begin{equation*}
\hat{\boldsymbol{x}}(\omega)=\frac{\hat{f}(\omega)}{\omega_{0}^{2}-2 i \gamma \omega-\omega^{2}} \equiv \hat{\chi}(\omega) \hat{f}(\omega) \tag{2.5}
\end{equation*}
$$

where $\hat{\chi}(\omega)$ is the susceptibility function:

$$
\begin{equation*}
\hat{\chi}(\omega)=\frac{1}{\omega_{0}^{2}-2 i \gamma \omega-\omega^{2}}=\frac{-1}{\left(\omega-\omega_{+}\right)\left(\omega-\omega_{-}\right)}, \tag{2.6}
\end{equation*}
$$

with

$$
\begin{equation*}
\omega_{ \pm}=-i \gamma \pm \sqrt{\omega_{0}^{2}-\gamma^{2}} \tag{2.7}
\end{equation*}
$$

[^0]The complete solution to (2.1) is then

$$
\begin{equation*}
x(t)=\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi} \frac{\hat{f}(\omega) e^{-i \omega t}}{\omega_{0}^{2}-2 i \gamma \omega-\omega^{2}}+x_{\mathrm{h}}(t) \tag{2.8}
\end{equation*}
$$

where $x_{\mathrm{h}}(t)$ is the homogeneous solution,

$$
\begin{equation*}
x_{\mathrm{h}}(t)=A_{+} e^{-i \omega_{+} t}+A_{-} e^{-i \omega_{-} t} \tag{2.9}
\end{equation*}
$$

Since $\operatorname{Im}\left(\omega_{ \pm}\right)<0, x_{\mathrm{h}}(t)$ is a transient which decays in time. The coefficients $A_{ \pm}$may be chosen to satisfy initial conditions on $x(0)$ and $\dot{x}(0)$, but the system 'loses its memory' of these initial conditions after a finite time, and in steady state all that is left is the inhomogeneous piece, which is completely determined by the forcing.
In the time domain, we can write

$$
\begin{align*}
& x(t)=\int_{-\infty}^{\infty} d t^{\prime} \chi\left(t-t^{\prime}\right) f\left(t^{\prime}\right)  \tag{2.10}\\
& \chi(s) \equiv \int_{-\infty}^{\infty} \frac{d \omega}{2 \pi} \hat{\chi}(\omega) e^{-i \omega s} \tag{2.11}
\end{align*}
$$

which brings us to a very important and sensible result:

Claim: The response is causal, i.e. $\chi\left(t-t^{\prime}\right)=0$ when $t<t^{\prime}$, provided that $\hat{\chi}(\omega)$ is analytic in the upper half plane of the variable $\omega$.

Proof. Consider eqn. (2.11). Of $\hat{\chi}(\omega)$ is analytic in the upper half plane, then closing in the UHP we obtain $\chi(s<0)=0$.

For our example (2.6), we close in the LHP for $s>0$ and obtain

$$
\begin{align*}
\chi(s>0) & =(-2 \pi i) \sum_{\omega \in \operatorname{LHP}} \operatorname{Res}\left\{\frac{1}{2 \pi} \hat{\chi}(\omega) e^{-i \omega s}\right\} \\
& =\frac{i e^{-i \omega_{+} s}}{\omega_{+}-\omega_{-}}+\frac{i e^{-i \omega_{-} s}}{\omega_{-}-\omega_{+}}, \tag{2.12}
\end{align*}
$$

i.e.

$$
\chi(s)= \begin{cases}\frac{e^{-\gamma s}}{\sqrt{\omega_{0}^{2}-\gamma^{2}}} \sin \left(\sqrt{\omega_{0}^{2}-\gamma^{2}}\right) \Theta(s) & \text { if } \omega_{0}^{2}>\gamma^{2}  \tag{2.13}\\ \frac{e^{-\gamma s}}{\sqrt{\gamma^{2}-\omega_{0}^{2}}} \sinh \left(\sqrt{\gamma^{2}-\omega_{0}^{2}}\right) \Theta(s) & \text { if } \omega_{0}^{2}<\gamma^{2},\end{cases}
$$

where $\Theta(s)$ is the step function: $\Theta(s \geq 0)=1, \Theta(s<0)=0$. Causality simply means that events occuring after the time $t$ cannot influence the state of the system at $t$. Note that, in general, $\chi(t)$ describes the time-dependent response to a $\delta$-function impulse at $t=0$.

### 2.1.1 Energy Dissipation

How much work is done by the force $f(t)$ ? Since the power applied is $P(t)=f(t) \dot{x}(t)$, we have

$$
\begin{align*}
P(t) & =\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi}(-i \omega) \hat{\chi}(\omega) \hat{f}(\omega) e^{-i \omega t} \int_{-\infty}^{\infty} \frac{d \nu}{2 \pi} \hat{f}^{*}(\nu) e^{+i \nu t}  \tag{2.14}\\
\Delta E & =\int_{-\infty}^{\infty} d t P(t)=\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi}(-i \omega) \hat{\chi}(\omega)|\hat{f}(\omega)|^{2} \tag{2.15}
\end{align*}
$$

Separating $\hat{\chi}(\omega)$ into real and imaginary parts,

$$
\begin{equation*}
\hat{\chi}(\omega)=\hat{\chi}^{\prime}(\omega)+i \hat{\chi}^{\prime \prime}(\omega) \tag{2.16}
\end{equation*}
$$

we find for our example

$$
\begin{align*}
\hat{\chi}^{\prime}(\omega) & =\frac{\omega_{0}^{2}-\omega^{2}}{\left(\omega_{0}^{2}-\omega^{2}\right)^{2}+4 \gamma^{2} \omega^{2}}=+\hat{\chi}^{\prime}(-\omega)  \tag{2.17}\\
\hat{\chi}^{\prime \prime}(\omega) & =\frac{2 \gamma \omega}{\left(\omega_{0}^{2}-\omega^{2}\right)^{2}+4 \gamma^{2} \omega^{2}}=-\hat{\chi}^{\prime \prime}(-\omega) \tag{2.18}
\end{align*}
$$

The energy dissipated may now be written

$$
\begin{equation*}
\Delta E=\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi} \omega \hat{\chi}^{\prime \prime}(\omega)|\hat{f}(\omega)|^{2} \tag{2.19}
\end{equation*}
$$

The even function $\hat{\chi}^{\prime}(\omega)$ is called the reactive part of the susceptibility; the odd function $\hat{\chi}^{\prime \prime}(\omega)$ is the dissipative part. When experimentalists measure a lineshape, they usually are referring to features in $\omega \hat{\chi}^{\prime \prime}(\omega)$, which describes the absorption rate as a function of driving frequency.

### 2.2 Kramers-Kronig Relations

Let $\chi(z)$ be a complex function of the complex variable $z$ which is analytic in the upper half plane. Then the following integral must vanish,

$$
\begin{equation*}
\oint_{\mathcal{C}} \frac{d z}{2 \pi i} \frac{\chi(z)}{z-\zeta}=0 \tag{2.20}
\end{equation*}
$$

whenever $\operatorname{Im}(\zeta) \leq 0$, where $\mathcal{C}$ is the contour depicted in fig. 2.1.
Now let $\omega \in \mathbb{R}$ be real, and define the complex function $\chi(\omega)$ of the real variable $\omega$ by

$$
\begin{equation*}
\chi(\omega) \equiv \lim _{\epsilon \rightarrow 0^{+}} \chi(\omega+i \epsilon) \tag{2.21}
\end{equation*}
$$

Assuming $\chi(z)$ vanishes sufficiently rapidly that Jordan's lemma may be invoked (i.e. that the integral of $\chi(z)$ along the arc of $\mathcal{C}$ vanishes), we have

$$
\begin{align*}
0 & =\int_{-\infty}^{\infty} \frac{d \nu}{2 \pi i} \frac{\chi(\nu)}{\nu-\omega+i \epsilon} \\
& =\int_{-\infty}^{\infty} \frac{d \nu}{2 \pi i}\left[\chi^{\prime}(\nu)+i \chi^{\prime \prime}(\nu)\right]\left[\frac{\mathcal{P}}{\nu-\omega}-i \pi \delta(\nu-\omega)\right] \tag{2.22}
\end{align*}
$$



Figure 2.1: The complex integration contour $\mathcal{C}$.
where $\mathcal{P}$ stands for 'principal part'. Taking the real and imaginary parts of this equation reveals the Kramers-Kronig relations:

$$
\begin{align*}
\chi^{\prime}(\omega) & =\mathcal{P} \int_{-\infty}^{\infty} \frac{d \nu}{\pi} \frac{\chi^{\prime \prime}(\nu)}{\nu-\omega}  \tag{2.23}\\
\chi^{\prime \prime}(\omega) & =-\mathcal{P} \int_{-\infty}^{\infty} \frac{d \nu}{\pi} \frac{\chi^{\prime}(\nu)}{\nu-\omega} \tag{2.24}
\end{align*}
$$

The Kramers-Kronig relations are valid for any function $\chi(z)$ which is analytic in the upper half plane. If $\chi(z)$ is analytic everywhere off the $\operatorname{Im}(z)=0$ axis, we may write

$$
\begin{equation*}
\chi(z)=\int_{-\infty}^{\infty} \frac{d \nu}{\pi} \frac{\chi^{\prime \prime}(\nu)}{\nu-z} \tag{2.25}
\end{equation*}
$$

This immediately yields the result

$$
\begin{equation*}
\lim _{\epsilon \rightarrow 0^{+}}[\chi(\omega+i \epsilon)-\chi(\omega-i \epsilon)]=2 i \chi^{\prime \prime}(\omega) \tag{2.26}
\end{equation*}
$$

As an example, consider the function

$$
\begin{equation*}
\chi^{\prime \prime}(\omega)=\frac{\omega}{\omega^{2}+\gamma^{2}} \tag{2.27}
\end{equation*}
$$

Then, choosing $\gamma>0$,

$$
\chi(z)=\int_{-\infty}^{\infty} \frac{d \omega}{\pi} \frac{1}{\omega-z} \cdot \frac{\omega}{\omega^{2}+\gamma^{2}}= \begin{cases}i /(z+i \gamma) & \text { if } \operatorname{Im}(z)>0  \tag{2.28}\\ -i /(z-i \gamma) & \text { if } \operatorname{Im}(z)<0\end{cases}
$$

Note that $\chi(z)$ is separately analytic in the UHP and the LHP, but that there is a branch cut along the $\operatorname{Re}(z)$ axis, where $\chi(\omega \pm i \epsilon)= \pm i /(\omega \pm i \gamma)$.
EXERCISE: Show that eqn. (2.26) is satisfied for $\chi(\omega)=\omega /\left(\omega^{2}+\gamma^{2}\right)$.

If we analytically continue $\chi(z)$ from the UHP into the LHP, we find a pole and no branch cut:

$$
\begin{equation*}
\tilde{\chi}(z)=\frac{i}{z+i \gamma} \tag{2.29}
\end{equation*}
$$

The pole lies in the LHP at $z=-i \gamma$.

### 2.3 Quantum Mechanical Response Functions

Now consider a general quantum mechanical system with a Hamiltonian $\mathcal{H}_{0}$ subjected to a time-dependent perturbation, $\mathcal{H}_{1}(t)$, where

$$
\begin{equation*}
\mathcal{H}_{1}(t)=-\sum_{i} Q_{i} \phi_{i}(t) \tag{2.30}
\end{equation*}
$$

Here, the $\left\{Q_{i}\right\}$ are a set of Hermitian operators, and the $\left\{\phi_{i}(t)\right\}$ are fields or potentials. Some examples:

$$
\mathcal{H}_{1}(t)= \begin{cases}-\boldsymbol{M} \cdot \boldsymbol{B}(t) & \text { magnetic moment - magnetic field } \\ \int d^{3} r \varrho(\boldsymbol{r}) \phi(\boldsymbol{r}, t) & \text { density - scalar potential } \\ -\frac{1}{c} \int d^{3} r \boldsymbol{j}(\boldsymbol{r}) \cdot \boldsymbol{A}(\boldsymbol{r}, t) & \text { electromagnetic current - vector potential }\end{cases}
$$

We now ask, what is $\left\langle Q_{i}(t)\right\rangle$ ? We assume that the lowest order response is linear, i.e.

$$
\begin{equation*}
\left\langle Q_{i}(t)\right\rangle=\int_{-\infty}^{\infty} d t^{\prime} \chi_{i j}\left(t-t^{\prime}\right) \phi_{j}\left(t^{\prime}\right)+\mathcal{O}\left(\phi_{k} \phi_{l}\right) \tag{2.31}
\end{equation*}
$$

Note that we assume that the $\mathcal{O}\left(\phi^{0}\right)$ term vanishes, which can be assured with a judicious choice of the $\left\{Q_{i}\right\}^{2}$. We also assume that the responses are all causal, i.e. $\chi_{i j}\left(t-t^{\prime}\right)=0$ for $t<t^{\prime}$. To compute $\chi_{i j}\left(t-t^{\prime}\right)$, we will use first order perturbation theory to obtain $\left\langle Q_{i}(t)\right\rangle$ and then functionally differentiate with respect to $\phi_{j}\left(t^{\prime}\right)$ :

$$
\begin{equation*}
\chi_{i j}\left(t-t^{\prime}\right)=\frac{\delta\left\langle Q_{i}(t)\right\rangle}{\delta \phi_{j}\left(t^{\prime}\right)} \tag{2.32}
\end{equation*}
$$

The first step is to establish the result,

$$
\begin{equation*}
|\Psi(t)\rangle=\mathcal{T} \exp \left\{-\frac{i}{\hbar} \int_{t_{0}}^{t} d t^{\prime}\left[\mathcal{H}_{0}+\mathcal{H}_{1}\left(t^{\prime}\right)\right]\right\}\left|\Psi\left(t_{0}\right)\right\rangle \tag{2.33}
\end{equation*}
$$

where $\mathcal{T}$ is the time ordering operator, which places earlier times to the right. This is easily derived starting with the Schrödinger equation,

$$
\begin{equation*}
i \hbar \frac{d}{d t}|\Psi(t)\rangle=\mathcal{H}(t)|\Psi(t)\rangle \tag{2.34}
\end{equation*}
$$

where $\mathcal{H}(t)=\mathcal{H}_{0}+\mathcal{H}_{1}(t)$. Integrating this equation from $t$ to $t+d t$ gives

$$
\begin{align*}
|\Psi(t+d t)\rangle & =\left(1-\frac{i}{\hbar} \mathcal{H}(t) d t\right)|\Psi(t)\rangle  \tag{2.35}\\
\left|\Psi\left(t_{0}+N d t\right)\right\rangle & =\left(1-\frac{i}{\hbar} \mathcal{H}\left(t_{0}+(N-1) d t\right)\right) \cdots\left(1-\frac{i}{\hbar} \mathcal{H}\left(t_{0}\right)\right)\left|\Psi\left(t_{0}\right)\right\rangle \tag{2.36}
\end{align*}
$$

[^1]hence
\[

$$
\begin{align*}
& \left|\Psi\left(t_{2}\right)\right\rangle=U\left(t_{2}, t_{1}\right)\left|\Psi\left(t_{1}\right)\right\rangle  \tag{2.37}\\
& U\left(t_{2}, t_{1}\right)=\mathcal{T} \exp \left\{-\frac{i}{\hbar} \int_{t_{1}}^{t_{2}} d t \mathcal{H}(t)\right\} \tag{2.38}
\end{align*}
$$
\]

$U\left(t_{2}, t_{1}\right)$ is a unitary operator (i.e. $U^{\dagger}=U^{-1}$ ), known as the time evolution operator between times $t_{1}$ and $t_{2}$.
EXERCISE: Show that, for $t_{1}<t_{2}<t_{3}$ that $U\left(t_{3}, t_{1}\right)=U\left(t_{3}, t_{2}\right) U\left(t_{2}, t_{1}\right)$.
If $t_{1}<t<t_{2}$, then differentiating $U\left(t_{2}, t_{1}\right)$ with respect to $\phi_{i}(t)$ yields

$$
\begin{equation*}
\frac{\delta U\left(t_{2}, t_{1}\right)}{\delta \phi_{j}(t)}=\frac{i}{\hbar} U\left(t_{2}, t\right) Q_{j} U\left(t, t_{1}\right) \tag{2.39}
\end{equation*}
$$

since $\partial \mathcal{H}(t) / \partial \phi_{j}(t)=-Q_{j}$. We may therefore write (assuming $t_{0}<t, t^{\prime}$ )

$$
\begin{align*}
\left.\frac{\delta|\Psi(t)\rangle}{\delta \phi_{j}\left(t^{\prime}\right)}\right|_{\left\{\phi_{i}=0\right\}} & =\frac{i}{\hbar} e^{-i \mathcal{H}_{0}\left(t-t^{\prime}\right) / \hbar} Q_{j} e^{-i \mathcal{H}_{0}\left(t^{\prime}-t_{0}\right) / \hbar}\left|\Psi\left(t_{0}\right)\right\rangle \Theta\left(t-t^{\prime}\right) \\
& =\frac{i}{\hbar} e^{-i \mathcal{H}_{0} t / \hbar} Q_{j}\left(t^{\prime}\right) e^{+i \mathcal{H}_{0} t_{0} / \hbar}\left|\Psi\left(t_{0}\right)\right\rangle \Theta\left(t-t^{\prime}\right) \tag{2.40}
\end{align*}
$$

where

$$
\begin{equation*}
Q_{j}(t) \equiv e^{i \mathcal{H}_{0} t / \hbar} Q_{j} e^{-i \mathcal{H}} \mathcal{H}_{0} t / \hbar \tag{2.41}
\end{equation*}
$$

is the operator $Q_{j}$ in the time-dependent interaction representation. Finally, we have

$$
\begin{align*}
\chi_{i j}\left(t-t^{\prime}\right)= & \frac{\delta}{\delta \phi_{j}\left(t^{\prime}\right)}\langle\Psi(t)| Q_{i}|\Psi(t)\rangle \\
= & \frac{\delta\langle\Psi(t)|}{\delta \phi_{j}\left(t^{\prime}\right)} Q_{i}|\Psi(t)\rangle+\langle\Psi(t)| Q_{i} \frac{\delta|\Psi(t)\rangle}{\delta \phi_{j}\left(t^{\prime}\right)} \\
= & \left\{-\frac{i}{\hbar}\left\langle\Psi\left(t_{0}\right)\right| e^{-i \mathcal{H}_{0} t_{0} / \hbar} Q_{j}\left(t^{\prime}\right) e^{+i \mathcal{H}_{0} t / \hbar} Q_{i}|\Psi(t)\rangle\right. \\
& \left.+\frac{i}{\hbar}\langle\Psi(t)| Q_{i} e^{-i \mathcal{H}_{0} t / \hbar} Q_{j}\left(t^{\prime}\right) e^{+i \mathcal{H}_{0} t_{0} / \hbar}\left|\Psi\left(t_{0}\right)\right\rangle\right\} \Theta\left(t-t^{\prime}\right) \\
= & \frac{i}{\hbar}\left\langle\left[Q_{i}(t), Q_{j}\left(t^{\prime}\right)\right]\right\rangle \Theta\left(t-t^{\prime}\right) \tag{2.42}
\end{align*}
$$

were averages are with respect to the wavefunction $|\Psi\rangle \equiv \exp \left(-i \mathcal{H}_{0} t_{0} / \hbar\right)\left|\Psi\left(t_{0}\right)\right\rangle$, with $t_{0} \rightarrow-\infty$, or, at finite temperature, with respect to a Boltzmann-weighted distribution of such states. To reiterate,

$$
\begin{equation*}
\chi_{i j}\left(t-t^{\prime}\right)=\frac{i}{\hbar}\left\langle\left[Q_{i}(t), Q_{j}\left(t^{\prime}\right)\right]\right\rangle \Theta\left(t-t^{\prime}\right) \tag{2.43}
\end{equation*}
$$

This is sometimes known as the retarded response function.

### 2.3.1 Spectral Representation

We now derive an expression for the response functions in terms of the spectral properties of the Hamiltonian $\mathcal{H}_{0}$. We stress that $\mathcal{H}_{0}$ may describe a fully interacting system. Write $\mathcal{H}_{0}|n\rangle=\hbar \omega_{n}|n\rangle$, in which case

$$
\begin{align*}
& \hat{\chi}_{i j}(\omega)= \frac{i}{\hbar} \int_{0}^{\infty} d t e^{i \omega t}\left\langle\left[Q_{i}(t), Q_{j}(0)\right]\right\rangle \\
&=\frac{i}{\hbar} \int_{0}^{\infty} d t e^{i \omega t} \frac{1}{Z} \sum_{m, n} e^{-\beta \hbar \omega_{m}}\left\{\langle m| Q_{i}|n\rangle\langle n| Q_{j}|m\rangle e^{+i\left(\omega_{m}-\omega_{n}\right) t}\right. \\
&\left.\quad-\langle m| Q_{j}|n\rangle\langle n| Q_{i}|m\rangle e^{+i\left(\omega_{n}-\omega_{m}\right) t}\right\}, \tag{2.44}
\end{align*}
$$

where $\beta=1 / k_{\mathrm{B}} T$ and $Z$ is the partition function. Regularizing the integrals at $t \rightarrow \infty$ with $\exp (-\epsilon t)$ with $\epsilon=0^{+}$, we use

$$
\begin{equation*}
\int_{0}^{\infty} d t e^{i(\omega-\Omega+i \epsilon) t}=\frac{i}{\omega-\Omega+i \epsilon} \tag{2.45}
\end{equation*}
$$

to obtain the spectral representation of the (retarded) response function ${ }^{3}$,

$$
\begin{equation*}
\hat{\chi}_{i j}(\omega+i \epsilon)=\frac{1}{\hbar Z} \sum_{m, n} e^{-\beta \hbar \omega_{m}}\left\{\frac{\langle m| Q_{j}|n\rangle\langle n| Q_{i}|m\rangle}{\omega-\omega_{m}+\omega_{n}+i \epsilon}-\frac{\langle m| Q_{i}|n\rangle\langle n| Q_{j}|m\rangle}{\omega+\omega_{m}-\omega_{n}+i \epsilon}\right\} \tag{2.46}
\end{equation*}
$$

We will refer to this as $\hat{\chi}_{i j}(\omega)$; formally $\hat{\chi}_{i j}(\omega)$ has poles or a branch cut (for continuous spectra) along the $\operatorname{Re}(\omega)$ axis. Diagrammatic perturbation theory does not give us $\hat{\chi}_{i j}(\omega)$, but rather the time-ordered response function,

$$
\begin{align*}
\chi_{i j}^{\mathrm{T}}\left(t-t^{\prime}\right) & \equiv \frac{i}{\hbar}\left\langle\mathcal{T} Q_{i}(t) Q_{j}\left(t^{\prime}\right)\right\rangle \\
& =\frac{i}{\hbar}\left\langle Q_{i}(t) Q_{j}\left(t^{\prime}\right)\right\rangle \Theta\left(t-t^{\prime}\right)+\frac{i}{\hbar}\left\langle Q_{j}\left(t^{\prime}\right) Q_{i}(t)\right\rangle \Theta\left(t^{\prime}-t\right) . \tag{2.47}
\end{align*}
$$

The spectral representation of $\hat{\chi}_{i j}^{\mathrm{T}}(\omega)$ is

$$
\begin{equation*}
\hat{\chi}_{i j}^{\mathrm{T}}(\omega+i \epsilon)=\frac{1}{\hbar Z} \sum_{m, n} e^{-\beta \hbar \omega_{m}}\left\{\frac{\langle m| Q_{j}|n\rangle\langle n| Q_{i}|m\rangle}{\omega-\omega_{m}+\omega_{n}-i \epsilon}-\frac{\langle m| Q_{i}|n\rangle\langle n| Q_{j}|m\rangle}{\omega+\omega_{m}-\omega_{n}+i \epsilon}\right\} \tag{2.48}
\end{equation*}
$$

The difference between $\hat{\chi}_{i j}(\omega)$ and $\hat{\chi}_{i j}^{\mathrm{T}}(\omega)$ is thus only in the sign of the infinitesimal $\pm i \epsilon$ term in one of the denominators.
Let us now define the real and imaginary parts of the product of expectations values encountered above:

$$
\begin{equation*}
\langle m| Q_{i}|n\rangle\langle n| Q_{j}|m\rangle \equiv A_{m n}(i j)+i B_{m n}(i j) . \tag{2.49}
\end{equation*}
$$

That is ${ }^{4}$,

$$
\begin{align*}
& A_{m n}(i j)=\frac{1}{2}\langle m| Q_{i}|n\rangle\langle n| Q_{j}|m\rangle+\frac{1}{2}\langle m| Q_{j}|n\rangle\langle n| Q_{i}|m\rangle  \tag{2.50}\\
& B_{m n}(i j)=\frac{1}{2 i}\langle m| Q_{i}|n\rangle\langle n| Q_{j}|m\rangle-\frac{1}{2 i}\langle m| Q_{j}|n\rangle\langle n| Q_{i}|m\rangle . \tag{2.51}
\end{align*}
$$

[^2]Note that $A_{m n}(i j)$ is separately symmetric under interchange of either $m$ and $n$, or of $i$ and $j$, whereas $B_{m n}(i j)$ is separately antisymmetric under these operations:

$$
\begin{align*}
& A_{m n}(i j)=+A_{n m}(i j)=A_{n m}(j i)=+A_{m n}(j i)  \tag{2.52}\\
& B_{m n}(i j)=-B_{n m}(i j)=B_{n m}(j i)=-B_{m n}(j i) \tag{2.53}
\end{align*}
$$

We define the spectral densities

$$
\left\{\begin{array}{l}
\varrho_{i j}^{A}(\omega)  \tag{2.54}\\
\varrho_{i j}^{B}(\omega)
\end{array}\right\} \equiv \frac{1}{\hbar Z} \sum_{m, n} e^{-\beta \hbar \omega_{m}}\left\{\begin{array}{l}
A_{m n}(i j) \\
B_{m n}(i j)
\end{array}\right\} \delta\left(\omega-\omega_{n}+\omega_{m}\right),
$$

which satisfy

$$
\begin{array}{lll}
\varrho_{i j}^{A}(\omega)=+\varrho_{j i}^{A}(\omega) & , & \varrho_{i j}^{A}(-\omega)=+e^{-\beta \hbar \omega} \varrho_{i j}^{A}(\omega) \\
\varrho_{i j}^{B}(\omega)=-\varrho_{j i}^{B}(\omega) & , & \varrho_{i j}^{B}(-\omega)=-e^{-\beta \hbar \omega} \varrho_{i j}^{B}(\omega) \tag{2.56}
\end{array}
$$

In terms of these spectral densities,

$$
\begin{align*}
& \hat{\chi}_{i j}^{\prime}(\omega)=\mathcal{P} \int_{-\infty}^{\infty} d \nu \frac{2 \nu}{\nu^{2}-\omega^{2}} \varrho_{i j}^{A}(\nu)-\pi\left(1-e^{-\beta \hbar \omega}\right) \varrho_{i j}^{B}(\omega)=+\hat{\chi}_{i j}^{\prime}(-\omega)  \tag{2.57}\\
& \hat{\chi}_{i j}^{\prime \prime}(\omega)=\mathcal{P} \int_{-\infty}^{\infty} d \nu \frac{2 \omega}{\nu^{2}-\omega^{2}} \varrho_{i j}^{B}(\nu)+\pi\left(1-e^{-\beta \hbar \omega}\right) \varrho_{i j}^{A}(\omega)=-\hat{\chi}_{i j}^{\prime \prime}(-\omega) . \tag{2.58}
\end{align*}
$$

For the time ordered response functions, we find

$$
\begin{align*}
\hat{\chi}_{i j}^{\prime \mathrm{T}}(\omega) & =\mathcal{P} \int_{-\infty}^{\infty} d \nu \frac{2 \nu}{\nu^{2}-\omega^{2}} \varrho_{i j}^{A}(\nu)-\pi\left(1+e^{-\beta \hbar \omega}\right) \varrho_{i j}^{B}(\omega)  \tag{2.59}\\
\hat{\chi}_{i j}^{\prime \prime \mathrm{T}}(\omega) & =\mathcal{P} \int_{-\infty}^{\infty} d \nu \frac{2 \omega}{\nu^{2}-\omega^{2}} \varrho_{i j}^{B}(\nu)+\pi\left(1+e^{-\beta \hbar \omega}\right) \varrho_{i j}^{A}(\omega) . \tag{2.60}
\end{align*}
$$

Hence, knowledge of either the retarded or the time-ordered response functions is sufficient to determine the full behavior of the other:

$$
\begin{align*}
& {\left[\hat{\chi}_{i j}^{\prime}(\omega)+\hat{\chi}_{j i}^{\prime}(\omega)\right] }=\left[\hat{\chi}_{i j}^{\prime \mathrm{T}}(\omega)+\hat{\chi}_{j i}^{\prime \mathrm{T}}(\omega)\right]  \tag{2.61}\\
& {\left[\hat{\chi}_{i j}^{\prime}(\omega)-\hat{\chi}_{j i}^{\prime}(\omega)\right] }=\left[\hat{\chi}_{i j}^{\prime \mathrm{T}}(\omega)-\hat{\chi}_{j i}^{\prime \mathrm{T}}(\omega)\right] \times \tanh \left(\frac{1}{2} \beta \hbar \omega\right)  \tag{2.62}\\
& {\left[\hat{\chi}_{i j}^{\prime \prime}(\omega)+\hat{\chi}_{j i}^{\prime \prime}(\omega)\right]=\left[\hat{\chi}_{i j}^{\prime \mathrm{T}}(\omega)+\hat{\chi}_{j i}^{\prime \prime \mathrm{T}}(\omega)\right] \times \tanh \left(\frac{1}{2} \beta \hbar \omega\right) }  \tag{2.63}\\
& {\left[\hat{\chi}_{i j}^{\prime \prime}(\omega)-\hat{\chi}_{j i}^{\prime \prime}(\omega)\right] }=\left[\hat{\chi}_{i j}^{\prime \prime \mathrm{T}}(\omega)-\hat{\chi}_{j i}^{\prime \prime \mathrm{T}}(\omega)\right] . \tag{2.64}
\end{align*}
$$

### 2.3.2 Energy Dissipation

The work done on the system must be positive! he rate at which work is done by the external fields is the power dissipated,

$$
\begin{align*}
P & =\frac{d}{d t}\langle\Psi(t)| \mathcal{H}(t)|\Psi(t)\rangle \\
& =\langle\Psi(t)| \frac{\partial \mathcal{H}_{1}(t)}{\partial t}|\Psi(t)\rangle=-\sum_{i}\left\langle Q_{i}(t)\right\rangle \dot{\phi}_{i}(t) \tag{2.65}
\end{align*}
$$

where we have invoked the Feynman-Hellman theorem. The total energy dissipated is thus a functional of the external fields $\left\{\phi_{i}(t)\right\}$ :

$$
\begin{align*}
W=\int_{-\infty}^{\infty} d t P(t) & =-\int_{-\infty}^{\infty} d t \int_{-\infty}^{\infty} d t^{\prime} \chi_{i j}\left(t-t^{\prime}\right) \dot{\phi}_{i}(t) \phi_{j}\left(t^{\prime}\right) \\
& =\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi}(-i \omega) \hat{\phi}_{i}^{*}(\omega) \hat{\chi}_{i j}(\omega) \hat{\phi}_{j}(\omega) . \tag{2.66}
\end{align*}
$$

Since the $\left\{Q_{i}\right\}$ are Hermitian observables, the $\left\{\phi_{i}(t)\right\}$ must be real fields, in which case $\hat{\phi}_{i}^{*}(\omega)=\hat{\phi}_{i}(-\omega)$, whence

$$
\begin{align*}
W & =\int_{-\infty}^{\infty} \frac{d \omega}{4 \pi}(-i \omega)\left[\hat{\chi}_{i j}(\omega)-\hat{\chi}_{j i}(-\omega)\right] \hat{\phi}_{i}^{*}(\omega) \hat{\phi}_{j}(\omega) \\
& =\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi} \mathcal{M}_{i j}(\omega) \hat{\phi}_{i}^{*}(\omega) \hat{\phi}_{j}(\omega) \tag{2.67}
\end{align*}
$$

where

$$
\begin{align*}
\mathcal{M}_{i j}(\omega) & \equiv \frac{1}{2}(-i \omega)\left[\hat{\chi}_{i j}(\omega)-\hat{\chi}_{j i}(-\omega)\right] \\
& =\pi \omega\left(1-e^{-\beta \hbar \omega}\right)\left(\varrho_{i j}^{A}(\omega)+i \varrho_{i j}^{B}(\omega)\right) . \tag{2.68}
\end{align*}
$$

Note that as a matrix $M(\omega)=M^{\dagger}(\omega)$, so that $M(\omega)$ has real eigenvalues.

### 2.3.3 Correlation Functions

We define the correlation function

$$
\begin{equation*}
S_{i j}(t) \equiv\left\langle Q_{i}(t) Q_{j}\left(t^{\prime}\right)\right\rangle \tag{2.69}
\end{equation*}
$$

which has the spectral representation

$$
\begin{align*}
\hat{S}_{i j}(\omega) & =2 \pi \hbar\left[\varrho_{i j}^{A}(\omega)+i \varrho_{i j}^{B}(\omega)\right] \\
& =\frac{2 \pi}{Z} \sum_{m, n} e^{-\beta \hbar \omega_{m}}\langle m| Q_{i}|n\rangle\langle n| Q_{j}|n\rangle \delta\left(\omega-\omega_{n}+\omega_{m}\right) \tag{2.70}
\end{align*}
$$

Note that

$$
\begin{equation*}
\hat{S}_{i j}(-\omega)=e^{-\beta \hbar \omega} \hat{S}_{i j}^{*}(\omega) \quad, \quad \hat{S}_{j i}(\omega)=\hat{S}_{i j}^{*}(\omega) \tag{2.71}
\end{equation*}
$$

and that

$$
\begin{equation*}
\hat{\chi}_{i j}(\omega)-\hat{\chi}_{j i}(-\omega)=\frac{i}{\hbar}\left(1-e^{-\beta \hbar \omega}\right) \hat{S}_{i j}(\omega) \tag{2.72}
\end{equation*}
$$

This result is known as the fluctuation-dissipation theorem, as it relates the equilibrium fluctuations $S_{i j}(\omega)$ to the dissipative quantity $\hat{\chi}_{i j}(\omega)-\hat{\chi}_{j i}(-\omega)$.

## Time Reversal Symmetry

If the operators $Q_{i}$ have a definite symmetry under time reversal, say

$$
\begin{equation*}
\mathcal{T} Q_{i} \mathcal{T}^{-1}=\eta_{i} Q_{i} \tag{2.73}
\end{equation*}
$$

then the correlation function satisfies

$$
\begin{equation*}
\hat{S}_{i j}(\omega)=\eta_{i} \eta_{j} \hat{S}_{j i}(\omega) \tag{2.74}
\end{equation*}
$$

### 2.3.4 Continuous Systems

The indices $i$ and $j$ could contain spatial information as well. Typically we will separate out spatial degrees of freedom, and write

$$
\begin{equation*}
S_{i j}\left(\boldsymbol{r}-\boldsymbol{r}^{\prime}, t-t^{\prime}\right)=\left\langle Q_{i}(\boldsymbol{r}, t) Q_{j}\left(\boldsymbol{r}^{\prime}, t^{\prime}\right)\right\rangle \tag{2.75}
\end{equation*}
$$

where we have assumed space and time translation invariance. The Fourier transform is defined as

$$
\begin{align*}
\hat{S}(\boldsymbol{k}, \omega) & =\int d^{3} r \int_{-\infty}^{\infty} d t e^{-i \boldsymbol{k} \cdot \boldsymbol{r}} S(\boldsymbol{r}, t)  \tag{2.76}\\
& =\frac{1}{V} \int_{-\infty}^{\infty} d t e^{+i \omega t}\langle\hat{Q}(\boldsymbol{k}, t) \hat{Q}(-\boldsymbol{k}, 0)\rangle \tag{2.77}
\end{align*}
$$

### 2.4 Density-Density Correlations

In many systems, external probes couple to the number density $n(\boldsymbol{r})=\sum_{i=1}^{N} \delta\left(\boldsymbol{r}-\boldsymbol{r}_{i}\right)$, and we may write the perturbing Hamiltonian as

$$
\begin{equation*}
\hat{H}_{1}(t)=-\int d^{3} r n(\boldsymbol{r}) U(\boldsymbol{r}, t) \tag{2.78}
\end{equation*}
$$

The response $\delta n \equiv n-\langle n\rangle_{0}$ is given by

$$
\begin{align*}
\langle\delta n(\boldsymbol{r}, t)\rangle & =\int d^{3} r^{\prime} \int d t^{\prime} \chi\left(\boldsymbol{r}-\boldsymbol{r}^{\prime}, t-t^{\prime}\right) U\left(\boldsymbol{r}^{\prime}, t^{\prime}\right)  \tag{2.79}\\
\langle\delta \hat{n}(\boldsymbol{q}, \omega)\rangle & =\chi(\boldsymbol{q}, \omega) \hat{U}(\boldsymbol{q}, \omega)
\end{align*}
$$

where

$$
\begin{align*}
\chi(\boldsymbol{q}, \omega) & =\frac{1}{\hbar V Z} \sum_{m, n} e^{-\beta \hbar \omega_{m}}\left\{\frac{\left.\left|\langle m| \hat{n}_{\boldsymbol{q}}\right| n\right\rangle\left.\right|^{2}}{\omega-\omega_{m}+\omega_{n}+i \epsilon}-\frac{\left.\left|\langle m| \hat{n}_{\boldsymbol{q}}\right| n\right\rangle\left.\right|^{2}}{\omega+\omega_{m}-\omega_{n}+i \epsilon}\right\} \\
& =\frac{1}{\hbar} \int_{-\infty}^{\infty} d \nu S(\boldsymbol{q}, \nu)\left\{\frac{1}{\omega+\nu+i \epsilon}-\frac{1}{\omega-\nu+i \epsilon}\right\} \tag{2.80}
\end{align*}
$$

and

$$
\begin{equation*}
\left.S(\boldsymbol{q}, \omega)=\frac{2 \pi}{V Z} \sum_{m, n} e^{-\beta \hbar \omega_{m}}\left|\langle m| \hat{n}_{\boldsymbol{q}}\right| n\right\rangle\left.\right|^{2} \delta\left(\omega-\omega_{n}+\omega_{m}\right) \tag{2.81}
\end{equation*}
$$

Note that

$$
\begin{equation*}
\hat{n}_{\boldsymbol{q}}=\sum_{i=1}^{N} e^{-i \boldsymbol{q} \cdot \boldsymbol{r}_{i}} \tag{2.82}
\end{equation*}
$$

and that $\hat{n}_{\boldsymbol{q}}^{\dagger}=\hat{n}_{-\boldsymbol{q}} . S(\boldsymbol{q}, \omega)$ is known as the dynamic structure factor. In a scattering experiment, where an incident probe (e.g. a neutron) interacts with the system via a potential $U(\boldsymbol{r}-\boldsymbol{R})$, where $\boldsymbol{R}$ is the probe particle position, Fermi's Golden Rule says that the rate at which the incident particle deposits momentum $\hbar \boldsymbol{q}$ and energy $\hbar \omega$ into the system is

$$
\begin{align*}
\mathcal{I}(\boldsymbol{q}, \omega) & \left.=\frac{2 \pi}{\hbar Z} \sum_{m, n} e^{-\beta \hbar \omega_{m}}\left|\langle m ; \boldsymbol{p}| \hat{H}_{1}\right| n ; \boldsymbol{p}-\hbar \boldsymbol{q}\right\rangle\left.\right|^{2} \delta\left(\omega-\omega_{n}+\omega_{m}\right)  \tag{2.83}\\
& =\frac{1}{\hbar}|\hat{U}(\boldsymbol{q})|^{2} S(\boldsymbol{q}, \omega)
\end{align*}
$$

The quantity $|\hat{U}(\boldsymbol{q})|^{2}$ is called the form factor. In neutron scattering, the "on-shell" condition requires that the incident energy $\varepsilon$ and momentum $\boldsymbol{p}$ are related via the ballistic dispersion $\varepsilon=\boldsymbol{p}^{2} / 2 m_{\mathrm{n}}$. Similarly, the final energy and momentum are related, hence

$$
\begin{equation*}
\varepsilon-\hbar \omega=\frac{\boldsymbol{p}^{2}}{2 m_{\mathrm{n}}}-\hbar \omega=\frac{(\boldsymbol{p}-\hbar \boldsymbol{q})^{2}}{2 m_{\mathrm{n}}} \quad \Longrightarrow \quad \hbar \omega=\frac{\hbar \boldsymbol{q} \cdot \boldsymbol{p}}{m_{\mathrm{n}}}-\frac{\hbar^{2} \boldsymbol{q}^{2}}{2 m_{\mathrm{n}}} \tag{2.84}
\end{equation*}
$$

Hence for fixed momentum transfer $\hbar \boldsymbol{q}$, the frequency $\omega$ can be adjusted by varying the incident momentum $\boldsymbol{p}$.
Another case of interest is the response of a system to a foreign object moving with trajectory $\boldsymbol{R}(t)=\boldsymbol{V} t$. In this case, $U(\boldsymbol{r}, t)=U(\boldsymbol{r}-\boldsymbol{R}(t))$, and

$$
\begin{align*}
\hat{U}(\boldsymbol{q}, \omega) & =\int d^{3} r \int d t e^{-i \boldsymbol{q} \cdot \boldsymbol{r}} e^{i \omega t} U(\boldsymbol{r}-\boldsymbol{V} t)  \tag{2.85}\\
& =2 \pi \delta(\omega-\boldsymbol{q} \cdot \boldsymbol{V}) \hat{U}(\boldsymbol{q})
\end{align*}
$$

so that

$$
\begin{equation*}
\langle\delta n(\boldsymbol{q}, \omega)\rangle=2 \pi \delta(\omega-\boldsymbol{q} \cdot \boldsymbol{V}) \chi(\boldsymbol{q}, \omega) . \tag{2.86}
\end{equation*}
$$

### 2.4.1 Sum Rules

From eqn. (2.81) we find

$$
\begin{align*}
\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi} \omega S(\boldsymbol{q}, \omega) & \left.=\frac{1}{V Z} \sum_{m, n} e^{-\beta \hbar \omega_{m}}\left|\langle m| \hat{n}_{\boldsymbol{q}}\right| n\right\rangle\left.\right|^{2}\left(\omega_{n}-\omega_{m}\right) \\
& =\frac{1}{\hbar V Z} \sum_{m, n} e^{-\beta \hbar \omega_{m}}\langle m| \hat{n}_{\boldsymbol{q}}|n\rangle\langle n|\left[\hat{H}, \hat{n}_{\boldsymbol{q}}^{\dagger}\right]|m\rangle \\
& =\frac{1}{\hbar V}\left\langle\hat{n}_{\boldsymbol{q}}\left[\hat{H}, \hat{n}_{\boldsymbol{q}}^{\dagger}\right]\right\rangle=\frac{1}{2 \hbar V}\left\langle\left[\hat{n}_{\boldsymbol{q}},\left[\hat{H}, \hat{n}_{\boldsymbol{q}}^{\dagger}\right]\right]\right\rangle \tag{2.87}
\end{align*}
$$

where the last equality is guaranteed by $\boldsymbol{q} \rightarrow-\boldsymbol{q}$ symmetry. Now if the potential is velocity independent, i.e. if

$$
\begin{equation*}
\hat{H}=-\frac{\hbar^{2}}{2 m} \sum_{i=1}^{N} \boldsymbol{\nabla}_{i}^{2}+V\left(\boldsymbol{r}_{1}, \ldots, \boldsymbol{r}_{N}\right) \tag{2.88}
\end{equation*}
$$

then with $\hat{n}_{\boldsymbol{q}}^{\dagger}=\sum_{i=1}^{N} e^{i \boldsymbol{q} \cdot \boldsymbol{r}_{i}}$ we obtain

$$
\begin{align*}
{\left[\hat{H}, \hat{n}_{\boldsymbol{q}}^{\dagger}\right] } & =-\frac{\hbar^{2}}{2 m} \sum_{i=1}^{N}\left[\boldsymbol{\nabla}_{i}^{2}, e^{i \boldsymbol{q} \cdot \boldsymbol{r}_{i}}\right]  \tag{2.89}\\
& =\frac{\hbar^{2}}{2 i m} \boldsymbol{q} \cdot \sum_{i=1}^{N}\left(\boldsymbol{\nabla}_{i} e^{i \boldsymbol{q} \cdot \boldsymbol{r}_{i}}+e^{i \boldsymbol{q} \cdot \boldsymbol{r}_{i}} \boldsymbol{\nabla}_{i}\right) \\
{\left[\hat{n}_{\boldsymbol{q}},\left[\hat{H}, \hat{n}_{\boldsymbol{q}}^{\dagger}\right]\right] } & =\frac{\hbar^{2}}{2 i m} \boldsymbol{q} \cdot \sum_{i=1}^{N} \sum_{j=1}^{N}\left[e^{-i \boldsymbol{q} \cdot \boldsymbol{r}_{j}}, \boldsymbol{\nabla}_{i} e^{i \boldsymbol{q} \cdot \boldsymbol{r}_{i}}+e^{i \boldsymbol{q} \cdot \boldsymbol{r}_{i}} \nabla_{i}\right] \\
& =N \hbar^{2} \boldsymbol{q}^{2} / m \tag{2.90}
\end{align*}
$$

We have derived the $f$-sum rule:

$$
\begin{equation*}
\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi} \omega S(\boldsymbol{q}, \omega)=\frac{N \hbar \boldsymbol{q}^{2}}{2 m V} \tag{2.91}
\end{equation*}
$$

Note that this integral, which is the first moment of the structure factor, is independent of the potential!

$$
\begin{equation*}
\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi} \omega^{n} S(\boldsymbol{q}, \omega)=\frac{1}{\hbar V}\langle\hat{n}_{\boldsymbol{q}}[\overbrace{\hat{H},[\hat{H}, \cdots[\hat{H}}^{n \text { times }}, \hat{n}_{\boldsymbol{q}}^{\dagger}] \cdots]]\rangle . \tag{2.92}
\end{equation*}
$$

Moments with $n>1$ in general do depend on the potential. The $n=0$ moment gives

$$
\begin{align*}
S(\boldsymbol{q}) & \equiv \int_{-\infty}^{\infty} \frac{d \omega}{2 \pi} \omega^{n} S(\boldsymbol{q}, \omega)=\frac{1}{\hbar V}\left\langle\hat{n}_{\boldsymbol{q}} \hat{n}_{\boldsymbol{q}}^{\dagger}\right\rangle \\
& =\frac{1}{\hbar} \int d^{3} r\langle n(\boldsymbol{r}) n(0)\rangle e^{-i \boldsymbol{q} \cdot \boldsymbol{r}}, \tag{2.93}
\end{align*}
$$

which is the Fourier transform of the density-density correlation function.

## Compressibility Sum Rule

The isothermal compressibility is given by

$$
\begin{equation*}
\kappa_{T}=-\left.\frac{1}{V} \frac{\partial V}{\partial n}\right|_{T}=\left.\frac{1}{n^{2}} \frac{\partial n}{\partial \mu}\right|_{T} \tag{2.94}
\end{equation*}
$$

Since a constant potential $U(\boldsymbol{r}, t)$ is equivalent to a chemical potential shift, we have

$$
\begin{equation*}
\langle\delta n\rangle=\chi(0,0) \delta \mu \quad \Longrightarrow \quad \kappa_{T}=\frac{1}{\hbar n^{2}} \lim _{\boldsymbol{q} \rightarrow 0} \int_{-\infty}^{\infty} \frac{d \omega}{\pi} \frac{S(\boldsymbol{q}, \omega)}{\omega} \tag{2.95}
\end{equation*}
$$

This is known as the compressibility sum rule.

### 2.5 Structure Factor for the Electron Gas

The dynamic structure factor $S(\boldsymbol{q}, \omega)$ tells us about the spectrum of density fluctuations. The density operator $\hat{n}_{\boldsymbol{q}}^{\dagger}=\sum_{i} e^{i \boldsymbol{q} \cdot \boldsymbol{r}_{i}}$ increases the wavevector by $\boldsymbol{q}$. At $T=0$, in order for $\langle n| \hat{n}_{\boldsymbol{q}}^{\dagger}|\mathrm{G}\rangle$ to be nonzero (where $|\mathrm{G}\rangle$ is the ground state, i.e. the filled Fermi sphere), the state $n$ must correspond to a particle-hole excitation. For a given $\boldsymbol{q}$, the maximum excitation frequency is obtained by taking an electron just inside the Fermi sphere, with wavevector $\boldsymbol{k}=k_{\mathrm{F}} \hat{\boldsymbol{q}}$ and transferring it to a state outside the Fermi sphere with wavevector $\boldsymbol{k}+\boldsymbol{q}$. For $|\boldsymbol{q}|<2 k_{\mathrm{F}}$, the minimum excitation frequency is zero - one can always form particle-hole excitations with states adjacent to the Fermi sphere. For $|\boldsymbol{q}|>2 k_{\mathrm{F}}$, the minimum excitation frequency is obtained by taking an electron just inside the Fermi sphere with wavevector $\boldsymbol{k}=-k_{\mathrm{F}} \hat{\boldsymbol{q}}$ to an unfilled state outside the Fermi sphere with wavevector $\boldsymbol{k}+\boldsymbol{q}$. These cases are depicted graphically in Fig. 2.2.

We therefore have

$$
\begin{align*}
& \omega_{\max }(q)=\frac{\hbar q^{2}}{2 m}+\frac{\hbar k_{\mathrm{F}} q}{m}  \tag{2.96}\\
& \omega_{\min }(q)= \begin{cases}0 & \text { if } q \leq 2 k_{\mathrm{F}} \\
\frac{\hbar q^{2}}{2 m}-\frac{\hbar k_{\mathrm{F}} q}{m} & \text { if } q>2 k_{\mathrm{F}} .\end{cases} \tag{2.97}
\end{align*}
$$

This is depicted in the left panel of Fig. 2.3. Outside of the region bounded by $\omega_{\min }(q)$ and $\omega_{\max }(q)$, there are no single pair excitations. It is of course easy to create multiple pair excitations with arbitrary energy and momentum, as depicted in the right panel of the figure. However, these multipair states do not couple to the ground state |G $\rangle$ through a single application of the density operator $\hat{n}_{\boldsymbol{q}}^{\dagger}$, hence they have zero oscillator strength: $\langle n| \hat{n}_{\boldsymbol{q}}^{\dagger}|\mathrm{G}\rangle=0$ for any multipair state $|n\rangle$.

### 2.5.1 Explicit $T=0$ Calculation

We start with

$$
\begin{align*}
S(\boldsymbol{r}, t) & =\langle n(\boldsymbol{r}, t) n(0,0)  \tag{2.98}\\
& =\int \frac{d^{3} k}{(2 \pi)^{3}} \int \frac{d^{3} k^{\prime}}{(2 \pi)^{3}} e^{i \boldsymbol{k} \cdot \boldsymbol{r}} \sum_{i, j}\left\langle e^{-i \boldsymbol{k} \cdot \boldsymbol{r}_{i}(t)} e^{i \boldsymbol{k}^{\prime} \cdot \boldsymbol{r}_{j}}\right\rangle . \tag{2.99}
\end{align*}
$$



Figure 2.2: Minimum and maximum frequency particle-hole excitations in the free electron gas at $T=0$. (a) To construct a maximum frequency excitation for a given $\boldsymbol{q}$, create a hole just inside the Fermi sphere at $\boldsymbol{k}=k_{\mathrm{F}} \hat{\boldsymbol{q}}$ and an electron at $\boldsymbol{k}^{\prime}=\boldsymbol{k}+\boldsymbol{q}$. (b) For $|\boldsymbol{q}|<2 k_{\mathrm{F}}$ the minumum excitation frequency is zero. (c) For $|\boldsymbol{q}|>2 k_{\mathrm{F}}$, the minimum excitation frequency is obtained by placing a hole at $\boldsymbol{k}=-k_{\mathrm{F}} \hat{\boldsymbol{q}}$ and an electron at $\boldsymbol{k}^{\prime}=\boldsymbol{k}+\boldsymbol{q}$.


Figure 2.3: Left: Minimum and maximum excitation frequency $\omega$ in units of $\varepsilon_{\mathrm{F}} / \hbar$ versus wavevector $q$ in units of $k_{\mathrm{F}}$. Outside the hatched areas, there are no single pair excitations. Right: With multiple pair excitations, every part of $(\boldsymbol{q}, \omega)$ space is accessible. However, these states to not couple to the ground state $|\mathrm{G}\rangle$ through a single application of the density operator $\hat{n}_{\boldsymbol{q}}^{\dagger}$.

The time evolution of the operator $\boldsymbol{r}_{i}(t)$ is given by $\boldsymbol{r}_{i}(t)=\boldsymbol{r}_{i}+\boldsymbol{p}_{i} t / m$, where $\boldsymbol{p}_{i}=-i \hbar \boldsymbol{\nabla}_{i}$. Using the result

$$
\begin{equation*}
e^{A+B}=e^{A} e^{B} e^{-\frac{1}{2}[A, B]} \tag{2.100}
\end{equation*}
$$

which is valid when $[A,[A, B]]=[B,[A, B]]=0$, we have

$$
\begin{equation*}
e^{-i \boldsymbol{k} \cdot \boldsymbol{r}_{i}(t)}=e^{i \hbar \boldsymbol{k}^{2} t / 2 m} e^{-i \boldsymbol{k} \cdot \boldsymbol{r}} e^{-i \boldsymbol{k} \cdot \boldsymbol{p}_{i} t / m} \tag{2.101}
\end{equation*}
$$

hence

$$
\begin{equation*}
S(\boldsymbol{r}, t)=\int \frac{d^{3} k}{(2 \pi)^{3}} \int \frac{d^{3} k^{\prime}}{(2 \pi)^{3}} e^{i \hbar \boldsymbol{k}^{2} t / 2 m} e^{i \boldsymbol{k} \cdot \boldsymbol{r}} \sum_{i, j}\left\langle e^{-i \boldsymbol{k} \cdot \boldsymbol{r}_{i}} e^{i \boldsymbol{k} \cdot \boldsymbol{p}_{i} t / m} e^{i \boldsymbol{k}^{\prime} \cdot \boldsymbol{r}_{j}}\right\rangle \tag{2.102}
\end{equation*}
$$

We now break the sum up into diagonal $(i=j)$ and off-diagonal $(i \neq j)$ terms.
For the diagonal terms, with $i=j$, we have

$$
\begin{align*}
\left\langle e^{-i \boldsymbol{k} \cdot \boldsymbol{r}_{i}} e^{i \boldsymbol{k} \cdot \boldsymbol{p}_{i} t / m} e^{i \boldsymbol{k}^{\prime} \cdot \boldsymbol{r}_{i}}\right\rangle & =e^{-i \hbar \boldsymbol{k} \cdot \boldsymbol{k}^{\prime} t / m}\left\langle e^{i(\boldsymbol{k}-\boldsymbol{k}) \cdot \boldsymbol{r}_{i}} e^{i \boldsymbol{k} \cdot \boldsymbol{p}_{i} t / m}\right\rangle  \tag{2.103}\\
& =e^{-i \hbar \boldsymbol{k} \cdot \boldsymbol{k}^{\prime} t / m} \frac{(2 \pi)^{3}}{N V} \delta\left(\boldsymbol{k}-\boldsymbol{k}^{\prime}\right) \sum_{\boldsymbol{q}} \Theta\left(k_{\mathrm{F}}-q\right) e^{-i \hbar \boldsymbol{k} \cdot \boldsymbol{q} t / m},
\end{align*}
$$

since the ground state $|\mathrm{G}\rangle$ is a Slater determinant formed of single particle wavefunctions $\psi_{\boldsymbol{k}}(\boldsymbol{r})=\exp (i \boldsymbol{q} \cdot \boldsymbol{r}) / \sqrt{V}$ with $q<k_{\mathrm{F}}$.

For $i \neq j$, we must include exchange effects. We then have

$$
\begin{align*}
\left\langle e^{-i \boldsymbol{k} \cdot \boldsymbol{r}_{i}} e^{i \boldsymbol{k} \cdot \boldsymbol{p}_{i} t / m} e^{i \boldsymbol{k}^{\prime} \cdot \boldsymbol{r}_{j}}\right\rangle= & \frac{1}{N(N-1)} \sum_{\boldsymbol{q}} \sum_{\boldsymbol{q}^{\prime}} \Theta\left(k_{\mathrm{F}}-q\right) \Theta\left(k_{\mathrm{F}}-q^{\prime}\right) \\
& \times \int \frac{d^{3} r_{i}}{V} \int \frac{d^{3} r_{j}}{V} e^{-i \hbar \boldsymbol{k} \cdot \boldsymbol{q} t / m}\left\{e^{-i \boldsymbol{k} \cdot \boldsymbol{r}_{i}} e^{i \boldsymbol{k}^{\prime} \boldsymbol{r}_{j}}\right. \\
& \left.-e^{i\left(\boldsymbol{q}-\boldsymbol{q}^{\prime}-\boldsymbol{k}\right) \cdot \boldsymbol{r}_{i}} e^{i\left(\boldsymbol{q}^{\prime}-\boldsymbol{q}+\boldsymbol{k}^{\prime}\right) \cdot \boldsymbol{r}_{j}}\right\} \\
= & \frac{(2 \pi)^{6}}{N(N-1) V^{2}} \sum_{\boldsymbol{q}} \sum_{\boldsymbol{q}^{\prime}} \Theta\left(k_{\mathrm{F}}-q\right) \Theta\left(k_{\mathrm{F}}-q^{\prime}\right) \\
& \times e^{-i \hbar \boldsymbol{k} \cdot \boldsymbol{q} t / m}\left\{\delta(\boldsymbol{k}) \delta\left(\boldsymbol{k}^{\prime}\right)-\delta\left(\boldsymbol{k}-\boldsymbol{k}^{\prime}\right) \delta\left(\boldsymbol{k}+\boldsymbol{q}^{\prime}-\boldsymbol{q}\right)\right\} . \tag{2.104}
\end{align*}
$$

Summing over the $i=j$ terms gives

$$
\begin{equation*}
S_{\mathrm{diag}}(\boldsymbol{r}, t)=\int \frac{d^{3} k}{(2 \pi)^{3}} e^{i \boldsymbol{k} \cdot \boldsymbol{r}} e^{-i \hbar \boldsymbol{k}^{2} t / 2 m} \int \frac{d^{3} q}{(2 \pi)^{3}} \Theta\left(k_{\mathrm{F}}-q\right) e^{-i \hbar \boldsymbol{k} \cdot \boldsymbol{q} t / m} \tag{2.105}
\end{equation*}
$$

while the off-diagonal terms yield

$$
\begin{align*}
& S_{\text {off-diag }}=\int \frac{d^{3} k}{(2 \pi)^{3}} e^{i \boldsymbol{k} \cdot \boldsymbol{r}} \int \frac{d^{3} q}{(2 \pi)^{3}} \int \frac{d^{3} q^{\prime}}{(2 \pi)^{3}} \Theta\left(k_{\mathrm{F}}-q\right) \Theta\left(k_{\mathrm{F}}-q^{\prime}\right) \\
& \times(2 \pi)^{3}\left\{\delta(\boldsymbol{k})-e^{+i \hbar \boldsymbol{k}^{2} t / 2 m} e^{-i \hbar \boldsymbol{k} \cdot \boldsymbol{q} t / m} \delta\left(\boldsymbol{q}-\boldsymbol{q}^{\prime}-\boldsymbol{k}\right)\right\} \\
&=n^{2}-\int \frac{d^{3} k}{(2 \pi)^{3}} e^{i \boldsymbol{k} \cdot \boldsymbol{r}} e^{+i \hbar \boldsymbol{k}^{2} t / 2 m} \int \frac{d^{3} q}{(2 \pi)^{3}} \Theta\left(k_{\mathrm{F}}-q\right) \Theta\left(k_{\mathrm{F}}-|\boldsymbol{k}-\boldsymbol{q}|\right) e^{-i \hbar \boldsymbol{k} \cdot \boldsymbol{q} t / m} \tag{2.106}
\end{align*}
$$

and hence

$$
\begin{align*}
S(\boldsymbol{k}, \omega)= & n^{2}(2 \pi)^{4} \delta(\boldsymbol{k}) \delta(\omega)+\int \frac{d^{3} q}{(2 \pi)^{3}} \Theta\left(k_{\mathrm{F}}-q\right)\left\{2 \pi \delta\left(\omega-\frac{\hbar \boldsymbol{k}^{2}}{2 m}-\frac{\hbar \boldsymbol{k} \cdot \boldsymbol{q}}{m}\right)\right. \\
& \left.-\Theta\left(k_{\mathrm{F}}-|\boldsymbol{k}-\boldsymbol{q}|\right) 2 \pi \delta\left(\omega+\frac{\hbar \boldsymbol{k}^{2}}{2 m}-\frac{\hbar \boldsymbol{k} \cdot \boldsymbol{q}}{m}\right)\right\}  \tag{2.107}\\
= & (2 \pi)^{4} n^{2} \delta(\boldsymbol{k}) \delta(\omega)+\int \frac{d^{3} q}{(2 \pi)^{3}} \Theta\left(k_{\mathrm{F}}-q\right) \Theta\left(|\boldsymbol{k}+\boldsymbol{q}|-k_{\mathrm{F}}\right) \cdot 2 \pi \delta\left(\omega-\frac{\hbar \boldsymbol{k}^{2}}{2 m}-\frac{\hbar \boldsymbol{k} \cdot \boldsymbol{q}}{m}\right) .
\end{align*}
$$

For nonzero $\boldsymbol{k}$ and $\omega$,

$$
\begin{align*}
S(\boldsymbol{k}, \omega) & =\frac{1}{2 \pi} \int_{0}^{k_{\mathrm{F}}} d q q^{2} \int_{-1}^{1} d x \Theta\left(\sqrt{k^{2}+q^{2}+2 k q x}-k_{\mathrm{F}}\right) \delta\left(\omega-\frac{\hbar k^{2}}{2 m}-\frac{\hbar k q}{m} x\right) \\
& =\frac{m}{2 \pi \hbar k} \int_{0}^{k_{\mathrm{F}}} d q q \Theta\left(\sqrt{q^{2}+\frac{2 m \omega}{\hbar}}-k_{\mathrm{F}}\right) \int_{-1}^{1} d x \delta\left(x+\frac{k}{2 q}-\frac{m \omega}{\hbar k q}\right) \\
& =\frac{m}{4 \pi \hbar k} \int_{0}^{k_{\mathrm{F}}^{2}} d u \Theta\left(u+\frac{2 m \omega}{\hbar}-k_{\mathrm{F}}^{2}\right) \Theta\left(u-\left|\frac{k}{2}-\frac{m \omega}{\hbar k}\right|^{2}\right) \tag{2.108}
\end{align*}
$$

The constraints on $u$ are

$$
\begin{equation*}
k_{\mathrm{F}}^{2} \geq u \geq \max \left(k_{\mathrm{F}}^{2}-\frac{2 m \omega}{\hbar},\left|\frac{k}{2}-\frac{m \omega}{\hbar k}\right|^{2}\right) . \tag{2.109}
\end{equation*}
$$

Clearly $\omega>0$ is required. There are two cases to consider.
The first case is

$$
\begin{equation*}
k_{\mathrm{F}}^{2}-\frac{2 m \omega}{\hbar} \geq\left|\frac{k}{2}-\frac{m \omega}{\hbar k}\right|^{2} \quad \Longrightarrow \quad 0 \leq \omega \leq \frac{\hbar k_{\mathrm{F}} k}{m}-\frac{\hbar k^{2}}{2 m} \tag{2.110}
\end{equation*}
$$

which in turn requires $k \leq 2 k_{\mathrm{F}}$. In this case, we have

$$
\begin{align*}
S(\boldsymbol{k}, \omega) & =\frac{m}{4 \pi \hbar k}\left\{k_{\mathrm{F}}^{2}-\left(k_{\mathrm{F}}^{2}-\frac{2 m \omega}{\hbar}\right)\right\} \\
& =\frac{m^{2} \omega}{2 \pi \hbar^{2} k} \tag{2.111}
\end{align*}
$$

The second case

$$
\begin{equation*}
k_{\mathrm{F}}^{2}-\frac{2 m \omega}{\hbar} \leq\left|\frac{k}{2}-\frac{m \omega}{\hbar k}\right|^{2} \quad \Longrightarrow \quad \omega \geq \frac{\hbar k_{\mathrm{F}} k}{m}-\frac{\hbar k^{2}}{2 m} \tag{2.112}
\end{equation*}
$$

However, we also have that

$$
\begin{equation*}
\left|\frac{k}{2}-\frac{m \omega}{\hbar k}\right|^{2} \leq k_{\mathrm{F}}^{2} \tag{2.113}
\end{equation*}
$$

hence $\omega$ is restricted to the range

$$
\begin{equation*}
\frac{\hbar k}{2 m}\left|k-2 k_{\mathrm{F}}\right| \leq \omega \leq \frac{\hbar k}{2 m}\left|k+2 k_{\mathrm{F}}\right| \tag{2.114}
\end{equation*}
$$

The integral in (2.108) then gives

$$
\begin{equation*}
S(\boldsymbol{k}, \omega)=\frac{m}{4 \pi \hbar k}\left\{k_{\mathrm{F}}^{2}-\left|\frac{k}{2}-\frac{m \omega}{\hbar k}\right|^{2}\right\} . \tag{2.115}
\end{equation*}
$$

Putting it all together,

$$
S(\boldsymbol{k}, \omega)= \begin{cases}\frac{m k_{\mathrm{F}}}{\pi^{2} \hbar^{2}} \cdot \frac{\pi \omega}{2 v_{\mathrm{F}} k} & \text { if } 0<\omega \leq v_{\mathrm{F}} k-\frac{\hbar k^{2}}{2 m}  \tag{2.116}\\ \frac{m k_{\mathrm{F}}}{\pi^{2} \hbar^{2}} \cdot \frac{\pi k_{\mathrm{F}}}{4 k}\left[1-\left(\frac{\omega}{v_{\mathrm{F}} k}-\frac{k}{2 k_{\mathrm{F}}}\right)^{2}\right] & \text { if }\left|v_{\mathrm{F}} k-\frac{\hbar k^{2}}{2 m}\right| \leq \omega \leq v_{\mathrm{F}} k+\frac{\hbar k^{2}}{2 m} \\ 0 & \text { if } \omega \geq v_{\mathrm{F}} k+\frac{\hbar k^{2}}{2 m}\end{cases}
$$

See the various plots in Fig. 2.4
Integrating over all frequency gives the static structure factor,

$$
\begin{equation*}
S(\boldsymbol{k})=\frac{1}{V}\left\langle n_{\boldsymbol{k}}^{\dagger} n_{\boldsymbol{k}}\right\rangle=\int_{-\infty}^{\infty} \frac{d \omega}{2 \pi} S(\boldsymbol{k}, \omega) \tag{2.117}
\end{equation*}
$$

The result is

$$
S(\boldsymbol{k})= \begin{cases}\left(\frac{3 k}{4 k_{\mathrm{F}}}-\frac{k^{3}}{16 k_{\mathrm{F}}^{3}}\right) n & \text { if } 0<k \leq 2 k_{\mathrm{F}}  \tag{2.118}\\ n & \text { if } k \geq 2 k_{\mathrm{F}} \\ V n^{2} & \text { if } k=0\end{cases}
$$

where $n=k_{\mathrm{F}}^{3} / 6 \pi^{2}$ is the density (per spin polarization).


Figure 2.4: The dynamic structure factor $S(k, \omega)$ for the electron gas at various values of $k / k_{\mathrm{F}}$.

### 2.6 Screening and Dielectric Response

### 2.6.1 Definition of the Charge Response Functions

Consider a many-electron system in the presence of a time-varying external charge density $\rho_{\text {ext }}(\boldsymbol{r}, t)$. The perturbing Hamiltonian is then

$$
\begin{align*}
\hat{H}_{1} & =-e \int d^{3} r \int d^{3} r^{\prime} \frac{n(\boldsymbol{r}) \rho_{\mathrm{ext}}(\boldsymbol{r}, t)}{\left|\boldsymbol{r}-\boldsymbol{r}^{\prime}\right|} \\
& =-e \int \frac{d^{3} k}{(2 \pi)^{3}} \frac{4 \pi}{\boldsymbol{k}^{2}} \hat{n}(\boldsymbol{k}) \hat{\rho}_{\mathrm{ext}}(-\boldsymbol{k}, t) \tag{2.119}
\end{align*}
$$

The induced charge is $-e \delta n$, where $\delta n$ is the induced number density:

$$
\begin{equation*}
\delta \hat{n}(\boldsymbol{q}, \omega)=\frac{4 \pi e}{\boldsymbol{q}^{2}} \chi(\boldsymbol{q}, \omega) \hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, \omega) . \tag{2.120}
\end{equation*}
$$

We can use this to determine the dielectric function $\epsilon(\boldsymbol{q}, \omega)$ :

$$
\begin{align*}
& \boldsymbol{\nabla} \cdot \boldsymbol{D}=4 \pi \rho_{\mathrm{ext}}  \tag{2.121}\\
& \nabla \cdot \boldsymbol{E}=4 \pi\left(\rho_{\mathrm{ext}}-e\langle\delta n\rangle\right)
\end{align*}
$$

In Fourier space,

$$
\begin{align*}
i \boldsymbol{q} \cdot \boldsymbol{D}(\boldsymbol{q}, \omega) & =4 \pi \hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, \omega)  \tag{2.122}\\
i \boldsymbol{q} \cdot \boldsymbol{E}(\boldsymbol{q}, \omega) & =4 \pi \hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, \omega)-4 \pi e\langle\delta \hat{n}(\boldsymbol{q}, \omega)\rangle
\end{align*}
$$

so that from $\boldsymbol{D}(\boldsymbol{q}, \omega)=\epsilon(\boldsymbol{q}, \omega) \boldsymbol{E}(\boldsymbol{q}, \omega)$ follows

$$
\begin{align*}
\frac{1}{\epsilon(\boldsymbol{q}, \omega)}=\frac{i \boldsymbol{q} \cdot \boldsymbol{E}(\boldsymbol{q}, \omega)}{i \boldsymbol{q} \cdot \boldsymbol{D}(\boldsymbol{q}, \omega)} & =1-\frac{\delta \hat{n}(\boldsymbol{q}, \omega)}{Z \hat{n}_{\mathrm{ext}}(\boldsymbol{q}, \omega)} \\
& =1-\frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \chi(\boldsymbol{q}, \omega) . \tag{2.123}
\end{align*}
$$

A system is said to exhibit perfect screening if

$$
\begin{equation*}
\epsilon(\boldsymbol{q} \rightarrow 0, \omega=0)=\infty \quad \Longrightarrow \quad \lim _{\boldsymbol{q} \rightarrow 0} \frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \chi(\boldsymbol{q}, 0)=1 \tag{2.124}
\end{equation*}
$$

Here, $\chi(\boldsymbol{q}, \omega)$ is the usual density-density response function,

$$
\begin{equation*}
\left.\chi(\boldsymbol{q}, \omega)=\frac{1}{\hbar V} \sum_{n} \frac{2 \omega_{n 0}}{\omega_{n 0}^{2}-(\omega+i \epsilon)^{2}}\left|\langle n| \hat{n}_{\boldsymbol{q}}\right| 0\right\rangle\left.\right|^{2}, \tag{2.125}
\end{equation*}
$$

where we content ourselves to work at $T=0$, and where $\omega_{n 0} \equiv \omega_{n}-\omega_{0}$ is the excitation frequency for the state $|n\rangle$.

From $\boldsymbol{j}_{\text {charge }}=\sigma \boldsymbol{E}$ and the continuity equation

$$
\begin{equation*}
i \boldsymbol{q} \cdot\left\langle\hat{\boldsymbol{j}}_{\text {charge }}(\boldsymbol{q}, \omega)\right\rangle=-i e \omega\langle\hat{n}(\boldsymbol{q}, \omega)\rangle=i \sigma(\boldsymbol{q}, \omega) \boldsymbol{q} \cdot \boldsymbol{E}(\boldsymbol{q}, \omega), \tag{2.126}
\end{equation*}
$$

we find

$$
\begin{equation*}
\left(4 \pi \hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, \omega)-4 \pi e\langle\delta \hat{n}(\boldsymbol{q}, \omega)\rangle\right) \sigma(\boldsymbol{q}, \omega)=-i \omega e\langle\delta \hat{n}(\boldsymbol{q}, \omega)\rangle \tag{2.127}
\end{equation*}
$$

or

$$
\begin{equation*}
\frac{4 \pi i}{\omega} \sigma(\boldsymbol{q}, \omega)=\frac{\langle\delta \hat{n}(\boldsymbol{q}, \omega)\rangle}{e^{-1} \hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, \omega)-\langle\delta \hat{n}(\boldsymbol{q}, \omega)\rangle}=\frac{1-\epsilon^{-1}(\boldsymbol{q}, \omega)}{\epsilon^{-1}(\boldsymbol{q} \omega)}=\epsilon(\boldsymbol{q}, \omega)-1 . \tag{2.128}
\end{equation*}
$$

Thus, we arrive at

$$
\begin{equation*}
\frac{1}{\epsilon(\boldsymbol{q}, \omega)}=1-\frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \chi(\boldsymbol{q}, \omega) \quad, \quad \epsilon(\boldsymbol{q}, \omega)=1+\frac{4 \pi i}{\omega} \sigma(\boldsymbol{q}, \omega) \tag{2.129}
\end{equation*}
$$

Taken together, these two equations allow us to relate the conductivity and the charge response function,

$$
\begin{equation*}
\sigma(\boldsymbol{q}, \omega)=-\frac{i \omega}{\boldsymbol{q}^{2}} \frac{e^{2} \chi(\boldsymbol{q}, \omega)}{1-\frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \chi(\boldsymbol{q}, \omega)} \tag{2.130}
\end{equation*}
$$

### 2.6.2 Static Screening: Thomas-Fermi Approximation

Imagine a time-independent, slowly varying electrical potential $\phi(\boldsymbol{r})$. We may define the 'local chemical potential' $\widetilde{\mu}(\boldsymbol{r})$ as

$$
\begin{equation*}
\mu \equiv \widetilde{\mu}(\boldsymbol{r})-e \phi(\boldsymbol{r}) \tag{2.131}
\end{equation*}
$$

where $\mu$ is the bulk chemical potential. The local chemical potential is related to the local density by local thermodynamics. At $T=0$,

$$
\begin{align*}
\widetilde{\mu}(\boldsymbol{r}) & \equiv \frac{\hbar^{2}}{2 m} k_{\mathrm{F}}^{2}(\boldsymbol{r})=\frac{\hbar^{2}}{2 m}\left(3 \pi^{2} n+3 \pi^{2} \delta n(\boldsymbol{r})\right)^{2 / 3} \\
& =\frac{\hbar^{2}}{2 m}\left(3 \pi^{2} n\right)^{2 / 3}\left\{1+\frac{2}{3} \frac{\delta n(\boldsymbol{r})}{n}+\ldots\right\}, \tag{2.132}
\end{align*}
$$

hence, to lowest order,

$$
\begin{equation*}
\delta n(\boldsymbol{r})=\frac{3 e n}{2 \mu} \phi(\boldsymbol{r}) \tag{2.133}
\end{equation*}
$$

This makes sense - a positive potential induces an increase in the local electron number density. In Fourier space,

$$
\begin{equation*}
\langle\delta \hat{n}(\boldsymbol{q}, \omega=0)\rangle=\frac{3 e n}{2 \mu} \hat{\phi}(\boldsymbol{q}, \omega=0) . \tag{2.134}
\end{equation*}
$$

Poisson's equation is $-\nabla^{2} \phi=4 \pi \rho_{\mathrm{tot}}$, i.e.

$$
\begin{align*}
i \boldsymbol{q} \cdot \boldsymbol{E}(\boldsymbol{q}, 0) & =\boldsymbol{q}^{2} \hat{\phi}(\boldsymbol{q}, 0) \\
& =4 \pi \hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, 0)-4 \pi e\langle\delta \hat{n}(\boldsymbol{q}, 0)\rangle  \tag{2.135}\\
& =4 \pi \hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, 0)-\frac{6 \pi n e^{2}}{\mu} \hat{\phi}(\boldsymbol{q}, 0)
\end{align*}
$$

and defining the Thomas-Fermi wavevector $q_{\text {TF }}$ by

$$
\begin{equation*}
q_{\mathrm{TF}}^{2} \equiv \frac{6 \pi n e^{2}}{\mu} \tag{2.136}
\end{equation*}
$$

we have

$$
\begin{equation*}
\hat{\phi}(\boldsymbol{q}, 0)=\frac{4 \pi \hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, 0)}{\boldsymbol{q}^{2}+q_{\mathrm{TF}}^{2}} \tag{2.137}
\end{equation*}
$$

hence

$$
\begin{equation*}
e\langle\delta \hat{n}(\boldsymbol{q}, 0)\rangle=\frac{q_{\mathrm{TF}}^{2}}{\boldsymbol{q}^{2}+q_{\mathrm{TF}}^{2}} \cdot \hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, 0) \quad \Longrightarrow \quad \epsilon(\boldsymbol{q}, 0)=1+\frac{q_{\mathrm{TF}}^{2}}{\boldsymbol{q}^{2}} \tag{2.138}
\end{equation*}
$$

Note that $\epsilon(\boldsymbol{q} \rightarrow 0, \omega=0)=\infty$, so there is perfect screening.
For a general electronic density of states $g(\varepsilon)$, we have $\delta n(\boldsymbol{r})=e \phi(\boldsymbol{r}) g\left(\varepsilon_{\mathrm{F}}\right)$, where $g\left(\varepsilon_{\mathrm{F}}\right)$ is the DOS at the Fermi energy. Invoking Poisson's equation then yields $q_{\mathrm{TF}}=\sqrt{4 \pi e^{2} g\left(\varepsilon_{\mathrm{F}}\right)}$.

The Thomas-Fermi wavelength is $\lambda_{\mathrm{TF}}=q_{\mathrm{TF}}^{-1}$, and for free electrons may be written as

$$
\begin{equation*}
\lambda_{\mathrm{TF}}=\left(\frac{\pi}{12}\right)^{1 / 6} \sqrt{r_{s}} a_{\mathrm{B}} \simeq 0.800 \sqrt{r_{s}} a_{\mathrm{B}} \tag{2.139}
\end{equation*}
$$

where $r_{s}$ is the dimensionless free electron sphere radius, in units of the Bohr radius $a_{\mathrm{B}}=\hbar^{2} / m e^{2}=0.529 \AA$, defined by $\frac{4}{3} \pi\left(r_{s} a_{\mathrm{B}}\right)^{3} n=1$, hence $r_{s} \propto n^{-1 / 3}$. Small $r_{s}$ corresponds to high density. Since Thomas-Fermi theory is a statistical theory, it can only be valid if there are many particles within a sphere of radius $\lambda_{\mathrm{TF}}$, i.e. $\frac{4}{3} \pi \lambda_{\mathrm{TF}}^{3} n>1$, or $r_{s} \lesssim(\pi / 12)^{1 / 3} \simeq 0.640$. TF theory is applicable only in the high density limit.

In the presence of a $\delta$-function external charge density $\rho_{\text {ext }}(\boldsymbol{r})=Z e \delta(\boldsymbol{r})$, we have $\hat{\rho}_{\text {ext }}(\boldsymbol{q}, 0)=Z e$ and

$$
\begin{equation*}
\langle\delta \hat{n}(\boldsymbol{q}, 0)\rangle=\frac{Z q_{\mathrm{TF}}^{2}}{\boldsymbol{q}^{2}+q_{\mathrm{TF}}^{2}} \Longrightarrow \quad\langle\delta n(\boldsymbol{r})\rangle=\frac{Z e^{-r / \lambda_{\mathrm{TF}}}}{4 \pi r} \tag{2.140}
\end{equation*}
$$



Figure 2.5: Perturbation expansion for RPA susceptibility bubble. Each bare bubble contributes a factor $\chi^{0}(\boldsymbol{q}, \omega)$ and each wavy interaction line $\hat{v}(\boldsymbol{q})$. The infinite series can be summed, yielding eqn. 2.151.

Note the decay on the scale of $\lambda_{\mathrm{TF}}$. Note also the perfect screening:

$$
\begin{equation*}
e\langle\delta \hat{n}(\boldsymbol{q} \rightarrow 0, \omega=0)\rangle=\hat{\rho}_{\mathrm{ext}}(\boldsymbol{q} \rightarrow 0, \omega=0)=Z e \tag{2.141}
\end{equation*}
$$

### 2.6.3 High Frequency Behavior of $\epsilon(\boldsymbol{q}, \omega)$

We have

$$
\begin{equation*}
\epsilon^{-1}(\boldsymbol{q}, \omega)=1-\frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \chi(\boldsymbol{q}, \omega) \tag{2.142}
\end{equation*}
$$

and, at $T=0$,

$$
\begin{equation*}
\left.\chi(\boldsymbol{q}, \omega)=\frac{1}{\hbar V} \sum_{j}\left|\langle j| \hat{n}_{\boldsymbol{q}}^{\dagger}\right| 0\right\rangle\left.\right|^{2}\left\{\frac{1}{\omega+\omega_{j 0}+i \epsilon}-\frac{1}{\omega-\omega_{j 0}+i \epsilon}\right\} \tag{2.143}
\end{equation*}
$$

where the number density operator is

$$
\hat{n}_{\boldsymbol{q}}^{\dagger}= \begin{cases}\sum_{i} e^{i \boldsymbol{q} \cdot \boldsymbol{r}_{i}} & \left(1^{\text {st }} \text { quantized }\right)  \tag{2.144}\\ \sum_{\boldsymbol{k}} \psi_{\boldsymbol{k}+\boldsymbol{q}}^{\dagger} \psi_{\boldsymbol{k}} & \left(2^{\text {nd }} \text { quantized: }\left\{\psi_{\boldsymbol{k}}, \psi_{\boldsymbol{k}^{\prime}}^{\dagger}\right\}=\delta_{\boldsymbol{k} \boldsymbol{k}^{\prime}}\right)\end{cases}
$$

Taking the limit $\omega \rightarrow \infty$, we find

$$
\begin{equation*}
\left.\chi(\boldsymbol{q}, \omega \rightarrow \infty)=-\frac{2}{\hbar V \omega^{2}} \sum_{j}\left|\langle j| \hat{n}_{\boldsymbol{q}}^{\dagger}\right| 0\right\rangle\left.\right|^{2} \omega_{j 0}=-\frac{2}{\hbar \omega^{2}} \int_{-\infty}^{\infty} \frac{d \omega^{\prime}}{2 \pi} \omega^{\prime} S\left(\boldsymbol{q}, \omega^{\prime}\right) \tag{2.145}
\end{equation*}
$$

Invoking the $f$-sum rule, the above integral is $n \hbar \boldsymbol{q}^{2} / 2 m$, hence

$$
\begin{equation*}
\chi(\boldsymbol{q}, \omega \rightarrow \infty)=-\frac{n \boldsymbol{q}^{2}}{m \omega^{2}} \tag{2.146}
\end{equation*}
$$

and

$$
\begin{equation*}
\epsilon^{-1}(\boldsymbol{q}, \omega \rightarrow \infty)=1+\frac{\omega_{\mathrm{p}}^{2}}{\omega^{2}}, \tag{2.147}
\end{equation*}
$$

where

$$
\begin{equation*}
\omega_{\mathrm{p}} \equiv \sqrt{\frac{4 \pi n e^{2}}{m}} \tag{2.148}
\end{equation*}
$$

is the plasma frequency.

### 2.6.4 Random Phase Approximation (RPA)

The electron charge appears nowhere in the free electron gas response function $\chi^{0}(\boldsymbol{q}, \omega)$. An interacting electron gas certainly does know about electron charge, since the Coulomb repulsion between electrons is part of the Hamiltonian. The idea behind the RPA is to obtain an approximation to the interacting $\chi(\boldsymbol{q}, \omega)$ from the noninteracting $\chi^{0}(\boldsymbol{q}, \omega)$ by self-consistently adjusting the charge so that the perturbing charge density is not $\rho_{\text {ext }}(\boldsymbol{r})$, but rather $\rho_{\text {ext }}(\boldsymbol{r}, t)-e\langle\delta n(\boldsymbol{r}, t)\rangle$. Thus, we write

$$
\begin{align*}
e\langle\delta \hat{n}(\boldsymbol{q}, \omega)\rangle & =\frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \chi^{\mathrm{RPA}}(\boldsymbol{q}, \omega) \hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, \omega)  \tag{2.149}\\
& =\frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \chi^{0}(\boldsymbol{q}, \omega)\left\{\hat{\rho}_{\mathrm{ext}}(\boldsymbol{q}, \omega)-e\langle\delta \hat{n}(\boldsymbol{q}, \omega)\rangle\right\}
\end{align*}
$$

which gives

$$
\begin{equation*}
\chi^{\mathrm{RPA}}(\boldsymbol{q}, \omega)=\frac{\chi^{0}(\boldsymbol{q}, \omega)}{1+\frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \chi^{0}(\boldsymbol{q}, \omega)} \tag{2.150}
\end{equation*}
$$

Several comments are in order.

1. If the electron-electron interaction were instead given by a general $\hat{v}(\boldsymbol{q})$ rather than the specific Coulomb form $\hat{v}(\boldsymbol{q})=4 \pi e^{2} / \boldsymbol{q}^{2}$, we would obtain

$$
\begin{equation*}
\chi^{\mathrm{RPA}}(\boldsymbol{q}, \omega)=\frac{\chi^{0}(\boldsymbol{q}, \omega)}{1+\hat{v}(\boldsymbol{q}) \chi^{0}(\boldsymbol{q}, \omega)} \tag{2.151}
\end{equation*}
$$

2. Within the RPA, there is perfect screening:

$$
\begin{equation*}
\lim _{q \rightarrow 0} \frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \chi^{\mathrm{RPA}}(\boldsymbol{q}, \omega)=1 \tag{2.152}
\end{equation*}
$$

3. The RPA expression may be expanded in an infinite series,

$$
\begin{equation*}
\chi^{\mathrm{RPA}}=\chi^{0}-\chi^{0} \hat{v} \chi^{0}+\chi^{0} \hat{v} \chi^{0} \hat{v} \chi^{0}-\ldots \tag{2.153}
\end{equation*}
$$

which has a diagrammatic interpretation, depicted in Fig. 2.5. The perturbative expansion in the interaction $\hat{v}$ may be resummed to yield the RPA result.
4. The RPA dielectric function takes the simple form

$$
\begin{equation*}
\epsilon^{\mathrm{RPA}}(\boldsymbol{q}, \omega)=1+\frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \chi^{0}(\boldsymbol{q}, \omega) \tag{2.154}
\end{equation*}
$$

5. Explicitly,

$$
\begin{align*}
\operatorname{Re} \epsilon^{\mathrm{RPA}}(\boldsymbol{q}, \omega)=1 & +\frac{q_{\mathrm{TF}}^{2}}{q^{2}}\left\{\frac{1}{2}+\frac{k_{\mathrm{F}}}{4 q}\left[\left(1-\frac{\left(\omega-\hbar q^{2} / 2 m\right)^{2}}{\left(v_{\mathrm{F}} q\right)^{2}}\right) \ln \left|\frac{\omega-v_{\mathrm{F}} q-\hbar q^{2} / 2 m}{\omega+v_{\mathrm{F}} q-\hbar q^{2} / 2 m}\right|\right.\right. \\
& \left.\left.+\left(1-\frac{\left(\omega-\hbar q^{2} / 2 m\right)^{2}}{\left(v_{\mathrm{F}} q\right)^{2}}\right) \ln \left|\frac{\omega-v_{\mathrm{F}} q+\hbar q^{2} / 2 m}{\omega+v_{\mathrm{F}} q+\hbar q^{2} / 2 m}\right|\right]\right\} \tag{2.155}
\end{align*}
$$

and

$$
\operatorname{Im} \epsilon^{\mathrm{RPA}}(\boldsymbol{q}, \omega)= \begin{cases}\frac{\pi \omega}{2 v_{\mathrm{F}} q} \cdot \frac{q_{\mathrm{TF}}^{2}}{q^{2}} & \text { if } 0 \leq \omega \leq v_{\mathrm{F}} q-\hbar q^{2} / 2 m  \tag{2.156}\\ \frac{\pi k_{\mathrm{F}}}{4 q}\left(1-\frac{\left(\omega-\hbar q^{2} / 2 m\right)^{2}}{\left(v_{\mathrm{F}} q\right)^{2}}\right) \frac{q_{\mathrm{TF}}^{2}}{q^{2}} & \text { if } v_{\mathrm{F}} q-\hbar q^{2} / 2 m \leq \omega \leq v_{\mathrm{F}} q+\hbar q^{2} / 2 m \\ 0 & \text { if } \omega>v_{\mathrm{F}} q+\hbar q^{2} / 2 m\end{cases}
$$

6. Note that

$$
\begin{equation*}
\epsilon^{\mathrm{RPA}}(\boldsymbol{q}, \omega \rightarrow \infty)=1-\frac{\omega_{\mathrm{p}}^{2}}{\omega^{2}}, \tag{2.157}
\end{equation*}
$$

in agreement with the $f$-sum rule, and

$$
\begin{equation*}
\epsilon^{\mathrm{RPA}}(\boldsymbol{q} \rightarrow 0, \omega=0)=1+\frac{q_{\mathrm{TF}}^{2}}{q^{2}}, \tag{2.158}
\end{equation*}
$$

in agreement with Thomas-Fermi theory.
7. At $\omega=0$ we have

$$
\begin{equation*}
\epsilon^{\mathrm{RPA}}(\boldsymbol{q}, 0)=1+\frac{q_{\mathrm{TF}}^{2}}{q^{2}}\left\{\frac{1}{2}+\frac{k_{\mathrm{F}}}{2 q}\left(1-\frac{q^{2}}{4 k_{\mathrm{F}}^{2}}\right) \ln \left|\frac{q+2 k_{\mathrm{F}}}{2-2 k_{\mathrm{F}}}\right|\right\}, \tag{2.159}
\end{equation*}
$$

which is real and which has a singularity at $q=2 k_{\mathrm{F}}$. This means that the long-distance behavior of $\langle\delta n(\boldsymbol{r})\rangle$ must oscillate. For a local charge perturbation, $\rho_{\text {ext }}(\boldsymbol{r})=Z e \delta(\boldsymbol{r})$, we have

$$
\begin{equation*}
\langle\delta n(\boldsymbol{r})\rangle=\frac{Z}{2 \pi^{2} r} \int_{0}^{\infty} d q q \sin (q r)\left\{1-\frac{1}{\epsilon(\boldsymbol{q}, 0)}\right\} \tag{2.160}
\end{equation*}
$$

and within the RPA one finds for long distances

$$
\begin{equation*}
\langle\delta n(\boldsymbol{r})\rangle \sim \frac{Z \cos \left(2 k_{\mathrm{F}} r\right)}{r^{3}}, \tag{2.161}
\end{equation*}
$$

rather than the Yukawa form familiar from Thomas-Fermi theory.

### 2.6.5 Plasmons

The RPA response function diverges when $\hat{v}(\boldsymbol{q}) \chi^{0}(\boldsymbol{q}, \omega)=-1$. For a given value of $\boldsymbol{q}$, this occurs for a specific value (or for a discrete set of values) of $\omega$, i.e. it defines a dispersion relation $\omega=\Omega(\boldsymbol{q})$. The poles of $\chi^{\mathrm{RPA}}$ and are identified with elementary excitations of the electron gas known as plasmons.
To find the plasmon dispersion, we first derive a result for $\chi^{0}(\boldsymbol{q}, \omega)$, starting with

$$
\begin{align*}
\chi^{0}(\boldsymbol{q}, t) & =\frac{i}{\hbar V}\langle[\hat{n}(\boldsymbol{q}, t), \hat{n}(-\boldsymbol{q}, 0)]\rangle \\
& =\frac{i}{\hbar V}\left\langle\left[\sum_{\boldsymbol{k} \sigma} \psi_{\boldsymbol{k}, \sigma}^{\dagger} \psi_{\boldsymbol{k}+\boldsymbol{q}, \sigma}, \sum_{\boldsymbol{k}^{\prime}, \sigma^{\prime}} \psi_{\boldsymbol{k}^{\prime}, \sigma^{\prime}}^{\dagger} \psi_{\boldsymbol{k}^{\prime}-\boldsymbol{q}, \sigma^{\prime}}\right]\right\rangle e^{i(\varepsilon(\boldsymbol{k})-\varepsilon(\boldsymbol{k}+\boldsymbol{q})) t / \hbar} \Theta(t), \tag{2.162}
\end{align*}
$$

where $\varepsilon(\boldsymbol{k})$ is the noninteracting electron dispersion. For a free electron gas, $\varepsilon(\boldsymbol{k})=\hbar^{2} \boldsymbol{k}^{2} / 2 m$. Next, using

$$
\begin{equation*}
[A B, C D]=A\{B, C\} D-\{A, C\} B D+C A\{B, D\}-C\{A, D\} B \tag{2.163}
\end{equation*}
$$

we obtain

$$
\begin{equation*}
\chi^{0}(\boldsymbol{q}, t)=\frac{i}{\hbar V} \sum_{\boldsymbol{k} \sigma}\left(f_{\boldsymbol{k}}-f_{\boldsymbol{k}+\boldsymbol{q}}\right) e^{i(\varepsilon(\boldsymbol{k})-\varepsilon(\boldsymbol{k}+\boldsymbol{q})) t / \hbar} \Theta(t) \tag{2.164}
\end{equation*}
$$

and therefore

$$
\begin{equation*}
\chi^{0}(\boldsymbol{q}, \omega)=2 \int \frac{d^{3} k}{(2 \pi)^{3}} \frac{f_{\boldsymbol{k}+\boldsymbol{q}}-f_{k}}{\hbar \omega-\varepsilon(\boldsymbol{k}+\boldsymbol{q})+\varepsilon(\boldsymbol{k})+i \epsilon} \tag{2.165}
\end{equation*}
$$

Here,

$$
\begin{equation*}
f_{k}=\frac{1}{e^{(\varepsilon(k)-\mu) / k_{\mathrm{B}} T}+1} \tag{2.166}
\end{equation*}
$$

is the Fermi distribution. At $T=0, f_{k}=\Theta\left(k_{\mathrm{F}}-k\right)$, and for $\omega \gg v_{\mathrm{F}} q$ we can expand $\chi^{0}(\boldsymbol{q}, \omega)$ in powers of $\omega^{-2}$, yielding

$$
\begin{equation*}
\chi^{0}(\boldsymbol{q}, \omega)=-\frac{k_{\mathrm{F}}^{3}}{3 \pi^{2}} \cdot \frac{q^{2}}{m \omega^{2}}\left\{1+\frac{3}{5}\left(\frac{\hbar k_{\mathrm{F}} q}{m \omega}\right)^{2}+\ldots\right\} \tag{2.167}
\end{equation*}
$$

so the resonance condition becomes

$$
\begin{align*}
0 & =1+\frac{4 \pi e^{2}}{q^{2}} \chi^{0}(\boldsymbol{q}, \omega) \\
& =1-\frac{\omega_{\mathrm{p}}^{2}}{\omega^{2}} \cdot\left\{1+\frac{3}{5}\left(\frac{v_{\mathrm{F}} q}{\omega}\right)^{2}+\ldots\right\} . \tag{2.168}
\end{align*}
$$

This gives the dispersion

$$
\begin{equation*}
\omega=\omega_{\mathrm{p}}\left\{1+\frac{3}{10}\left(\frac{v_{\mathrm{F}} q}{\omega_{\mathrm{p}}}\right)^{2}+\ldots\right\} . \tag{2.169}
\end{equation*}
$$

For the noninteracting electron gas, the energy of the particle-hole continuum is bounded by $\omega_{\min }(q)$ and $\omega_{\max }(q)$, which are given below in Eqs. 2.97 and 2.96. Eventually the plasmon penetrates the particle-hole continuum, at which point it becomes heavily damped, since it can decay into particle-hole excitations.

### 2.6.6 Jellium

Finally, consider an electron gas in the presence of a neutralizing ionic background. We assume one species of ion with mass $M_{\mathrm{i}}$ and charge $+Z_{\mathrm{i}} e$, and we smear the ionic charge into a continuum as an approximation. This nonexistent material is known in the business of many-body physics as jellium. Let the ion number density be $n_{\mathrm{i}}$, and the electron number density be $n_{\mathrm{e}}$. Then Laplace's equation says

$$
\begin{equation*}
\nabla^{2} \phi=-4 \pi \rho_{\text {charge }}=-4 \pi e\left(n_{\mathrm{i}}-n_{\mathrm{e}}+n_{\text {ext }}\right), \tag{2.170}
\end{equation*}
$$

where $n_{\text {ext }}=\rho_{\text {ext }} / e$, where $\rho_{\text {ext }}$ is the test charge density, regarded as a perturbation to the jellium. The ions move according to

$$
\begin{equation*}
M_{\mathrm{i}} \frac{d \boldsymbol{v}}{d t}=Z_{\mathrm{i}} e \boldsymbol{E}=-Z_{\mathrm{i}} e \boldsymbol{\nabla} \phi . \tag{2.171}
\end{equation*}
$$

They also satisfy continuity, which to lowest order in small quantities is governed by the equation

$$
\begin{equation*}
n_{\mathrm{i}}^{0} \boldsymbol{\nabla} \cdot \boldsymbol{v}+\frac{\partial n_{\mathrm{i}}}{\partial t}=0 \tag{2.172}
\end{equation*}
$$

where $n_{i}^{0}$ is the average ionic number density. Taking the time derivative of the above equation, and invoking Newton's law for the ion's as well as Laplace, we find

$$
\begin{equation*}
-\frac{\partial^{2} n_{\mathrm{i}}(\boldsymbol{x}, t)}{\partial t^{2}}=\frac{4 \pi n_{\mathrm{i}}^{0} Z_{\mathrm{i}} e^{2}}{M_{\mathrm{i}}}\left(n_{\mathrm{i}}(\boldsymbol{x}, t)+n_{\mathrm{ext}}(\boldsymbol{x}, t)-n_{\mathrm{e}}(\boldsymbol{x}, t)\right) \tag{2.173}
\end{equation*}
$$

In Fourier space,

$$
\begin{equation*}
\omega^{2} \hat{n}_{\mathrm{i}}(\boldsymbol{q}, \omega)=\Omega_{\mathrm{p}, \mathrm{i}}^{2}\left(\hat{n}_{\mathrm{i}}(\boldsymbol{q}, \omega)+\hat{n}_{\mathrm{ext}}(\boldsymbol{q}, \omega)-\hat{n}_{\mathrm{e}}(\boldsymbol{q}, \omega)\right) \tag{2.174}
\end{equation*}
$$

where

$$
\begin{equation*}
\Omega_{\mathrm{p}, \mathrm{i}}=\sqrt{\frac{4 \pi n_{\mathrm{i}}^{0} Z_{\mathrm{i}} e^{2}}{M_{\mathrm{i}}}} \tag{2.175}
\end{equation*}
$$

is the ionic plasma frequency. Typically $\Omega_{\mathrm{p}, \mathrm{i}} \approx 10^{13} \mathrm{~s}^{-1}$.
Since the ionic mass $M_{\mathrm{i}}$ is much greater than the electron mass, the ionic plasma frequency is much greater than the electron plasma frequency. We assume that the ions may be regarded as 'slow' and that the electrons respond according to Eqn. 2.120, viz.

$$
\begin{equation*}
\hat{n}_{\mathrm{e}}(\boldsymbol{q}, \omega)=\frac{4 \pi e}{\boldsymbol{q}^{2}} \chi_{\mathrm{e}}(\boldsymbol{q}, \omega)\left(\hat{n}_{\mathrm{i}}(\boldsymbol{q}, \omega)+\hat{n}_{\mathrm{ext}}(\boldsymbol{q}, \omega)\right) \tag{2.176}
\end{equation*}
$$

We then have

$$
\begin{equation*}
\frac{\omega^{2}}{\Omega_{\mathrm{p}, \mathrm{i}}^{2}} \hat{n}_{\mathrm{i}}(\boldsymbol{q}, \omega)=\frac{\hat{n}_{\mathrm{i}}(\boldsymbol{q}, \omega)+\hat{n}_{\mathrm{ext}}(\boldsymbol{q}, \omega)}{\epsilon_{\mathrm{e}}(\boldsymbol{q}, \omega)} \tag{2.177}
\end{equation*}
$$

From this equation, we obtain $\hat{n}_{\mathrm{i}}(\boldsymbol{q}, \omega)$ and then $n_{\text {tot }} \equiv n_{\mathrm{i}}-n_{\mathrm{e}}+n_{\text {ext }}$. We thereby obtain

$$
\begin{equation*}
\hat{n}_{\mathrm{tot}}(\boldsymbol{q}, \omega)=\frac{\hat{n}_{\mathrm{ext}}(\boldsymbol{q}, \omega)}{\epsilon_{\mathrm{e}}(\boldsymbol{q}, \omega)-\frac{\Omega_{\mathrm{p}, \mathrm{i}}^{2}}{\omega^{2}}} \tag{2.178}
\end{equation*}
$$

Finally, the dielectric function of the jellium system is given by

$$
\begin{align*}
\epsilon(\boldsymbol{q}, \omega) & =\frac{\hat{n}_{\mathrm{ext}}(\boldsymbol{q}, \omega)}{\hat{n}_{\mathrm{tot}}(\boldsymbol{q}, \omega)} \\
& =\epsilon_{\mathrm{e}}(\boldsymbol{q}, \omega)-\frac{\omega^{2}}{\Omega_{\mathrm{p}, \mathrm{i}}^{2}} \tag{2.179}
\end{align*}
$$

At frequencies low compared to the electron plasma frequency, we approximate $\epsilon_{\mathrm{e}}(\boldsymbol{q}, \omega)$ by the Thomas-Fermi form, $\epsilon_{\mathrm{e}}(\boldsymbol{q}, \omega) \approx\left(q^{2}+q_{\mathrm{TF}}^{2}\right) / q^{2}$. Then

$$
\begin{equation*}
\epsilon(\boldsymbol{q}, \omega) \approx 1+\frac{q_{\mathrm{TF}}^{2}}{q^{2}}-\frac{\Omega_{\mathrm{p}, \mathrm{i}}^{2}}{\omega^{2}} \tag{2.180}
\end{equation*}
$$

The zeros of this function, given by $\epsilon\left(\boldsymbol{q}, \omega_{\boldsymbol{q}}\right)=0$, occur for

$$
\begin{equation*}
\omega_{\boldsymbol{q}}=\frac{\Omega_{\mathrm{p}, \mathrm{i}} q}{\sqrt{q^{2}+q_{\mathrm{TF}}^{2}}} \tag{2.181}
\end{equation*}
$$

This allows us to write

$$
\begin{equation*}
\frac{4 \pi e^{2}}{\boldsymbol{q}^{2}} \frac{1}{\epsilon(\boldsymbol{q}, \omega)}=\frac{4 \pi e^{2}}{\boldsymbol{q}^{2}+q_{\mathrm{TF}}^{2}} \cdot \frac{\omega^{2}}{\omega^{2}-\omega_{\boldsymbol{q}}^{2}} \tag{2.182}
\end{equation*}
$$

This is interpreted as the effective interaction between charges in the jellium model, arising from both electronic and ionic screening. Note that the interaction is negative, i.e. attractive, for $\omega^{2}<\omega_{\boldsymbol{q}}^{2}$. At frequencies high compared to $\omega_{q}$, but low compared to the electronic plasma frequency, the effective potential is of the Yukawa form. Only the electrons then participate in screening, because the phonons are too slow.

### 2.7 Electromagnetic Response

Consider an interacting system consisting of electrons of charge $-e$ in the presence of a time-varying electromagnetic field. The electromagnetic field is given in terms of the 4-potential $A^{\mu}=\left(A^{0}, \boldsymbol{A}\right)$ :

$$
\begin{align*}
& \boldsymbol{E}=-\boldsymbol{\nabla} A^{0}-\frac{1}{c} \frac{\partial \boldsymbol{A}}{\partial t}  \tag{2.183}\\
& \boldsymbol{B}=\boldsymbol{\nabla} \times \boldsymbol{A} .
\end{align*}
$$

The Hamiltonian for an $N$-particle system is

$$
\begin{align*}
\hat{H}\left(A^{\mu}\right) & =\sum_{i=1}^{N}\left\{\frac{1}{2 m}\left(\boldsymbol{p}_{i}+\frac{e}{c} \boldsymbol{A}\left(\boldsymbol{x}_{i}, t\right)\right)^{2}-e A^{0}\left(\boldsymbol{x}_{i}, t\right)+U\left(\boldsymbol{x}_{i}\right)\right\}+\sum_{i<j} v\left(\boldsymbol{x}_{i}-\boldsymbol{x}_{j}\right)  \tag{2.184}\\
& =\hat{H}(0)-\frac{1}{c} \int d^{3} x j_{\mu}^{\mathrm{p}}(\boldsymbol{x}) A^{\mu}(\boldsymbol{x}, t)+\frac{e^{2}}{2 m c^{2}} \int d^{3} x n(\boldsymbol{x}) \boldsymbol{A}^{2}(\boldsymbol{x}, t)
\end{align*}
$$

where we have defined

$$
\begin{align*}
n(\boldsymbol{x}) & \equiv \sum_{i=1}^{N} \delta\left(\boldsymbol{x}-\boldsymbol{x}_{i}\right)  \tag{2.185}\\
\boldsymbol{j}^{\mathrm{p}}(\boldsymbol{x}) & \equiv-\frac{e}{2 m} \sum_{i=1}^{N}\left\{\boldsymbol{p}_{i} \delta\left(\boldsymbol{x}-\boldsymbol{x}_{i}\right)+\delta\left(\boldsymbol{x}-\boldsymbol{x}_{i}\right) \boldsymbol{p}_{i}\right\}  \tag{2.186}\\
j_{0}^{\mathrm{p}}(\boldsymbol{x}) & \equiv n(\boldsymbol{x}) e c \tag{2.187}
\end{align*}
$$

Throughout this discussion we invoke covariant/contravariant notation, using the metric

$$
g_{\mu \nu}=g^{\mu \nu}=\left(\begin{array}{cccc}
-1 & 0 & 0 & 0  \tag{2.188}\\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
0 & 0 & 0 & 1
\end{array}\right)
$$

so that if $j^{\mu}=\left(j^{0}, j^{1}, j^{2}, j^{3}\right) \equiv\left(j^{0}, \boldsymbol{j}\right)$, then

$$
\begin{align*}
j_{\mu} & =g_{\mu \nu} j^{\nu}=\left(-j^{0}, j^{1}, j^{2}, j^{3}\right)  \tag{2.189}\\
j_{\mu} A^{\mu} & =j^{\mu} g_{\mu \nu} A^{\nu}=-j^{0} A^{0}+j \cdot \boldsymbol{A} \equiv j \cdot A
\end{align*}
$$

The quantity $j_{\mu}^{\mathrm{p}}(\boldsymbol{x})$ is known as the paramagnetic current density. The physical current density $j_{\mu}(\boldsymbol{x})$ also contains a diamagnetic contribution:

$$
\begin{align*}
j_{\mu}(\boldsymbol{x}) & =-c \frac{\delta \hat{H}}{\delta A^{\mu}(\boldsymbol{x})}=j_{\mu}^{\mathrm{p}}(\boldsymbol{x})+j_{\mu}^{\mathrm{d}}(\boldsymbol{x})  \tag{2.190}\\
j^{\mathrm{d}}(\boldsymbol{x}) & =-\frac{e^{2}}{m c} n(\boldsymbol{x}) \boldsymbol{A}(\boldsymbol{x})=-\frac{e}{m c^{2}} j_{0}^{\mathrm{p}}(\boldsymbol{x}) \boldsymbol{A}(\boldsymbol{x})
\end{align*}
$$

The temporal component of the diamagnetic current is zero: $j_{0}^{\mathrm{d}}(\boldsymbol{x})=0$. The electromagnetic response tensor $K_{\mu \nu}$ is defined via

$$
\begin{equation*}
\left\langle j_{\mu}(\boldsymbol{x}, t)\right\rangle=-\frac{c}{4 \pi} \int d^{3} x^{\prime} \int d t K_{\mu \nu}\left(\boldsymbol{x} t ; \boldsymbol{x}^{\prime} t^{\prime}\right) A^{\nu}\left(\boldsymbol{x}^{\prime}, t^{\prime}\right), \tag{2.191}
\end{equation*}
$$

valid to first order in the external 4-potential $A^{\mu}$. From

$$
\begin{align*}
\left\langle j_{\mu}^{\mathrm{p}}(\boldsymbol{x}, t)\right\rangle & =\frac{i}{\hbar c} \int d^{3} x^{\prime} \int d t^{\prime}\left\langle\left[j_{\mu}^{\mathrm{p}}(\boldsymbol{x}, t), j_{\nu}^{\mathrm{p}}\left(\boldsymbol{x}^{\prime}, t^{\prime}\right)\right]\right\rangle \Theta\left(t-t^{\prime}\right) A^{\nu}\left(\boldsymbol{x}^{\prime}, t^{\prime}\right)  \tag{2.192}\\
\left\langle j_{\mu}^{\mathrm{d}}(\boldsymbol{x}, t)\right\rangle & =-\frac{e}{m c^{2}}\left\langle j_{0}^{\mathrm{p}}(\boldsymbol{x}, t)\right\rangle A^{\mu}(\boldsymbol{x}, t)\left(1-\delta_{\mu 0}\right)
\end{align*}
$$

we conclude

$$
\begin{align*}
K_{\mu \nu}\left(\boldsymbol{x} t ; \boldsymbol{x}^{\prime} t^{\prime}\right)= & \frac{4 \pi}{i \hbar c^{2}}\left\langle\left[j_{\mu}^{\mathrm{p}}(\boldsymbol{x}, t), j_{\nu}^{\mathrm{p}}\left(\boldsymbol{x}^{\prime}, t^{\prime}\right)\right]\right\rangle \Theta\left(t-t^{\prime}\right)  \tag{2.193}\\
& +\frac{4 \pi e}{m c^{3}}\left\langle j_{0}^{\mathrm{p}}(\boldsymbol{x}, t)\right\rangle \delta\left(\boldsymbol{x}-\boldsymbol{x}^{\prime}\right) \delta\left(t-t^{\prime}\right) \delta_{\mu \nu}\left(1-\delta_{\mu 0}\right)
\end{align*}
$$

The first term is sometimes known as the paramagnetic response kernel, $K_{\mu \nu}^{\mathrm{p}}\left(x ; x^{\prime}\right)=\left(4 \pi i / i \hbar c^{2}\right)\left\langle\left[j_{\mu}^{\mathrm{p}}(x), j_{\nu}^{\mathrm{p}}\left(x^{\prime}\right)\right]\right\rangle \Theta(t-$ $t^{\prime}$ ) is not directly calculable by perturbation theory. Rather, one obtains the time-ordered response function $K_{\mu \nu}^{\mathrm{p}, \mathrm{T}}\left(x ; x^{\prime}\right)=\left(4 \pi / i \hbar c^{2}\right)\left\langle\mathcal{T} j_{\mu}^{\mathrm{p}}(x) j_{\nu}^{\mathrm{p}}\left(x^{\prime}\right)\right\rangle$, where $x^{\mu} \equiv(c t, \boldsymbol{x})$.

## Second Quantized Notation

In the presence of an electromagnetic field described by the 4-potential $A^{\mu}=(c \phi, \boldsymbol{A})$, the Hamiltonian of an interacting electron system takes the form

$$
\begin{align*}
& \hat{H}=\sum_{\sigma} \int d^{3} x \psi_{\sigma}^{\dagger}(\boldsymbol{x})\left\{\frac{1}{2 m}\left(\frac{\hbar}{i} \boldsymbol{\nabla}+\frac{e}{c} \boldsymbol{A}\right)^{2}-e A^{0}(\boldsymbol{x})+U(\boldsymbol{x})\right\} \psi_{\sigma}(\boldsymbol{x}) \\
&+ \frac{1}{2} \sum_{\sigma, \sigma^{\prime}} \int d^{3} x \int d^{3} x^{\prime} \psi_{\sigma}^{\dagger}(\boldsymbol{x}) \psi_{\sigma^{\prime}}^{\dagger}\left(\boldsymbol{x}^{\prime}\right) v\left(\boldsymbol{x}-\boldsymbol{x}^{\prime}\right) \psi_{\sigma^{\prime}}\left(\boldsymbol{x}^{\prime}\right) \psi_{\sigma}(\boldsymbol{x}) \tag{2.194}
\end{align*}
$$

where $v\left(\boldsymbol{x}-\boldsymbol{x}^{\prime}\right)$ is a two-body interaction, e.g. $e^{2} /\left|\boldsymbol{x}-\boldsymbol{x}^{\prime}\right|$, and $U(\boldsymbol{x})$ is the external scalar potential. Expanding in powers of $A^{\mu}$,

$$
\begin{equation*}
\hat{H}\left(A^{\mu}\right)=\hat{H}(0)-\frac{1}{c} \int d^{3} x j_{\mu}^{\mathrm{p}}(\boldsymbol{x}) A^{\mu}(\boldsymbol{x})+\frac{e^{2}}{2 m c^{2}} \sum_{\sigma} \int d^{3} x \psi_{\sigma}^{\dagger}(\boldsymbol{x}) \psi_{\sigma}(\boldsymbol{x}) \boldsymbol{A}^{2}(\boldsymbol{x}) \tag{2.195}
\end{equation*}
$$

where the paramagnetic current density $j_{\mu}^{\mathrm{p}}(\boldsymbol{x})$ is defined by

$$
\begin{align*}
j_{0}^{\mathrm{p}}(\boldsymbol{x}) & =c e \sum_{\sigma} \psi_{\sigma}^{\dagger}(\boldsymbol{x}) \psi_{\sigma}(\boldsymbol{x}) \\
\boldsymbol{j}^{\mathrm{p}}(\boldsymbol{x}) & =\frac{i e \hbar}{2 m} \sum_{\sigma}\left\{\psi_{\sigma}^{\dagger}(\boldsymbol{x}) \boldsymbol{\nabla} \psi_{\sigma}(\boldsymbol{x})-\left(\boldsymbol{\nabla} \psi_{\sigma}^{\dagger}(\boldsymbol{x})\right) \psi_{\sigma}(\boldsymbol{x})\right\} . \tag{2.196}
\end{align*}
$$

### 2.7.1 Gauge Invariance and Charge Conservation

In Fourier space, with $q^{\mu}=(\omega / c, \boldsymbol{q})$, we have, for homogeneous systems,

$$
\begin{equation*}
\left\langle j_{\mu}(q)\right\rangle=-\frac{c}{4 \pi} K_{\mu \nu}(q) A^{\nu}(q) \tag{2.197}
\end{equation*}
$$

Note our convention on Fourier transforms:

$$
\begin{align*}
& H(x)=\int \frac{d^{4} k}{(2 \pi)^{4}} \hat{H}(k) e^{+i k \cdot x}  \tag{2.198}\\
& \hat{H}(k)=\int d^{4} x H(x) e^{-i k \cdot x}
\end{align*}
$$

where $k \cdot x \equiv k_{\mu} x^{\mu}=\boldsymbol{k} \cdot \boldsymbol{x}-\omega t$. Under a gauge transformation, $A^{\mu} \rightarrow A^{\mu}+\partial^{\mu} \Lambda$, i.e.

$$
\begin{equation*}
A^{\mu}(q) \rightarrow A^{\mu}(q)+i \Lambda(q) q^{\mu} \tag{2.199}
\end{equation*}
$$

where $\Lambda$ is an arbitrary scalar function. Since the physical current must be unchanged by a gauge transformation, we conclude that $K_{\mu \nu}(q) q^{\nu}=0$. We also have the continuity equation, $\partial^{\mu} j_{\mu}=0$, the Fourier space version of which says $q^{\mu} j_{\mu}(q)=0$, which in turn requires $q^{\mu} K_{\mu \nu}(q)=0$. Therefore,

$$
\begin{equation*}
\sum_{\mu} q^{\mu} K_{\mu \nu}(q)=\sum_{\nu} K_{\mu \nu}(q) q^{\nu}=0 \tag{2.200}
\end{equation*}
$$

In fact, the above conditions are identical owing to the reciprocity relations,

$$
\begin{align*}
& \operatorname{Re} K_{\mu \nu}(q)=+\operatorname{Re} K_{\nu \mu}(-q)  \tag{2.201}\\
& \operatorname{Im} K_{\mu \nu}(q)=-\operatorname{Im} K_{\nu \mu}(-q)
\end{align*}
$$

which follow from the spectral representation of $K_{\mu \nu}(q)$. Thus,

$$
\begin{equation*}
\text { gauge invariance } \Longleftrightarrow \text { charge conservation } \tag{2.202}
\end{equation*}
$$

### 2.7.2 A Sum Rule

If we work in a gauge where $A^{0}=0$, then $\boldsymbol{E}=-c^{-1} \dot{\boldsymbol{A}}$, hence $\boldsymbol{E}(q)=i q^{0} \boldsymbol{A}(q)$, and

$$
\begin{align*}
\left\langle j_{i}(q)\right\rangle & =-\frac{c}{4 \pi} K_{i j}(q) A^{j}(q) \\
& =-\frac{c}{4 \pi} K_{i j}(q) \frac{c}{i \omega} E^{j}(q) \\
& \equiv \sigma_{i j}(q) E^{j}(q) \tag{2.203}
\end{align*}
$$

Thus, the conductivity tensor is given by

$$
\begin{equation*}
\sigma_{i j}(\boldsymbol{q}, \omega)=\frac{i c^{2}}{4 \pi \omega} K_{i j}(\boldsymbol{q}, \omega) \tag{2.204}
\end{equation*}
$$

If, in the $\omega \rightarrow 0$ limit, the conductivity is to remain finite, then we must have

$$
\begin{equation*}
\int d^{3} x \int_{0}^{\infty} d t\left\langle\left[j_{i}^{\mathrm{p}}(\boldsymbol{x}, t), j_{j}^{\mathrm{p}}(0,0)\right]\right\rangle e^{+i \omega t}=\frac{i e^{2} n}{m} \delta_{i j} \tag{2.205}
\end{equation*}
$$

where $n$ is the electron number density. This relation is spontaneously violated in a superconductor, where $\sigma(\omega) \propto$ $\omega^{-1}$ as $\omega \rightarrow 0$.

### 2.7.3 Longitudinal and Transverse Response

In an isotropic system, the spatial components of $K_{\mu \nu}$ may be resolved into longitudinal and transverse components, since the only preferred spatial vector is $\boldsymbol{q}$ itself. Thus, we may write

$$
\begin{equation*}
K_{i j}(\boldsymbol{q}, \omega)=K_{\|}(\boldsymbol{q}, \omega) \hat{q}_{i} \hat{q}_{j}+K_{\perp}(\boldsymbol{q}, \omega)\left(\delta_{i j}-\hat{q}_{i} \hat{q}_{j}\right), \tag{2.206}
\end{equation*}
$$

where $\hat{q}_{i} \equiv q_{i} /|\boldsymbol{q}|$. We now invoke current conservation, which says $q^{\mu} K_{\mu \nu}(q)=0$. When $\nu=j$ is a spatial index,

$$
\begin{equation*}
q^{0} K_{0 j}+q^{i} K_{i j}=\frac{\omega}{c} K_{0 j}+K_{\|} q_{j}, \tag{2.207}
\end{equation*}
$$

which yields

$$
\begin{equation*}
K_{0 j}(\boldsymbol{q}, \omega)=-\frac{c}{\omega} q^{j} K_{\|}(\boldsymbol{q}, \omega)=K_{j 0}(\boldsymbol{q}, \omega) \tag{2.208}
\end{equation*}
$$

In other words, the three components of $K_{0 j}(q)$ are in fact completely determined by $K_{\|}(q)$ and $q$ itself. When $\nu=0$,

$$
\begin{equation*}
0=q^{0} K_{00}+q^{i} K_{i 0}=\frac{\omega}{c} K_{00}-\frac{c}{\omega} \boldsymbol{q}^{2} K_{\|} \tag{2.209}
\end{equation*}
$$

which says

$$
\begin{equation*}
K_{00}(\boldsymbol{q}, \omega)=\frac{c^{2}}{\omega^{2}} \boldsymbol{q}^{2} K_{\|}(\boldsymbol{q}, \omega) \tag{2.210}
\end{equation*}
$$

Thus, of the 10 freedoms of the symmetric $4 \times 4$ tensor $K_{\mu \nu}(q)$, there are only two independent ones - the functions $K_{\|}(q)$ and $K_{\perp}(q)$.

### 2.7.4 Neutral Systems

In neutral systems, we define the number density and number current density as

$$
\begin{align*}
n(\boldsymbol{x}) & =\sum_{i=1}^{N} \delta\left(\boldsymbol{x}-\boldsymbol{x}_{i}\right)  \tag{2.211}\\
\boldsymbol{j}(\boldsymbol{x}) & =\frac{1}{2 m} \sum_{i=1}^{N}\left\{\boldsymbol{p}_{i} \delta\left(\boldsymbol{x}-\boldsymbol{x}_{i}\right)+\delta\left(\boldsymbol{x}-\boldsymbol{x}_{i}\right) \boldsymbol{p}_{i}\right\}
\end{align*}
$$

The charge and current susceptibilities are then given by

$$
\begin{align*}
\chi(\boldsymbol{x}, t) & =\frac{i}{\hbar}\langle[n(\boldsymbol{x}, t), n(0,0)]\rangle \Theta(t) \\
\chi_{i j}(\boldsymbol{x}, t) & =\frac{i}{\hbar}\left\langle\left[j_{i}(\boldsymbol{x}, t), j_{j}(0,0)\right]\right\rangle \Theta(t) . \tag{2.212}
\end{align*}
$$

We define the longitudinal and transverse susceptibilities for homogeneous systems according to

$$
\begin{equation*}
\chi_{i j}(\boldsymbol{q}, \omega)=\chi_{\|}(\boldsymbol{q}, \omega) \hat{q}_{i} \hat{q}_{j}+\chi_{\perp}(\boldsymbol{q}, \omega)\left(\delta_{i j}-\hat{q}_{i} \hat{q}_{j}\right) . \tag{2.213}
\end{equation*}
$$

From the continuity equation,

$$
\begin{equation*}
\boldsymbol{\nabla} \cdot \boldsymbol{j}+\frac{\partial n}{\partial t}=0 \tag{2.214}
\end{equation*}
$$

follows the relation

$$
\begin{equation*}
\chi_{\|}(\boldsymbol{q}, \omega)=\frac{n}{m}+\frac{\omega^{2}}{\boldsymbol{q}^{2}} \chi(\boldsymbol{q}, \omega) . \tag{2.215}
\end{equation*}
$$

EXERCISE: Derive eqn. (2.215).

The relation between $K_{\mu \nu}(q)$ and the neutral susceptibilities defined above is then

$$
\begin{align*}
K_{00}(\boldsymbol{x}, t) & =-4 \pi e^{2} \chi(\boldsymbol{x}, t) \\
K_{i j}(\boldsymbol{x}, t) & =\frac{4 \pi e^{2}}{c^{2}}\left\{\frac{n}{m} \delta(\boldsymbol{x}) \delta(t)-\chi_{i j}(\boldsymbol{x}, t)\right\} \tag{2.216}
\end{align*}
$$

and therefore

$$
\begin{align*}
K_{\|}(\boldsymbol{q}, \omega) & =\frac{4 \pi e^{2}}{c^{2}}\left\{\frac{n}{m}-\chi_{\|}(\boldsymbol{q}, \omega)\right\} \\
K_{\perp}(\boldsymbol{q}, \omega) & =\frac{4 \pi e^{2}}{c^{2}}\left\{\frac{n}{m}-\chi_{\perp}(\boldsymbol{q}, \omega)\right\} \tag{2.217}
\end{align*}
$$

### 2.7.5 The Meissner Effect and Superfluid Density

Suppose we apply an electromagnetic field $\boldsymbol{E}$. We adopt a gauge in which $A^{0}=0$, in which case $\boldsymbol{E}=-c^{-1} \dot{\boldsymbol{A}}$. As always, $\boldsymbol{B}=\boldsymbol{\nabla} \times \boldsymbol{A}$. To satisfy Maxwell's equations, we have $\boldsymbol{q} \cdot \boldsymbol{A}(\boldsymbol{q}, \omega)=0$, i.e. $\boldsymbol{A}(\boldsymbol{q}, \omega)$ is purely transverse. But then

$$
\begin{equation*}
\langle\boldsymbol{j}(\boldsymbol{q}, \omega)\rangle=-\frac{c}{4 \pi} K_{\perp}(\boldsymbol{q}, \omega) \boldsymbol{A}(\boldsymbol{q}, \omega) . \tag{2.218}
\end{equation*}
$$

This leads directly to the Meissner effect whenever $\lim _{\boldsymbol{q} \rightarrow 0} K_{\perp}(\boldsymbol{q}, 0)$ is finite. To see this, we write

$$
\begin{align*}
\boldsymbol{\nabla} \times \boldsymbol{B}=\boldsymbol{\nabla}(\boldsymbol{\nabla} \cdot \boldsymbol{A})-\boldsymbol{\nabla}^{2} \boldsymbol{A} & =\frac{4 \pi}{c} \boldsymbol{j}+\frac{1}{c} \frac{\partial \boldsymbol{E}}{\partial t} \\
& =\frac{4 \pi}{c}\left(-\frac{c}{4 \pi}\right) K_{\perp}\left(-i \boldsymbol{\nabla}, i \partial_{t}\right) \boldsymbol{A}-\frac{1}{c^{2}} \frac{\partial^{2} \boldsymbol{A}}{\partial t^{2}} \tag{2.219}
\end{align*}
$$

which yields

$$
\begin{equation*}
\left(\boldsymbol{\nabla}^{2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{2}}\right) \boldsymbol{A}=K_{\perp}\left(-i \boldsymbol{\nabla}, i \partial_{t}\right) \boldsymbol{A} \tag{2.220}
\end{equation*}
$$

In the static limit, $\boldsymbol{\nabla}^{2} \boldsymbol{A}=K_{\perp}(i \boldsymbol{\nabla}, 0) \boldsymbol{A}$, and we define

$$
\begin{equation*}
\frac{1}{\lambda_{\mathrm{L}}^{2}} \equiv \lim _{\boldsymbol{q} \rightarrow 0} K_{\perp}(\boldsymbol{q}, 0) \tag{2.221}
\end{equation*}
$$

$\lambda_{\mathrm{L}}$ is the London penetration depth, which is related to the superfluid density $n_{\mathrm{s}}$ by

$$
\begin{align*}
n_{\mathrm{s}} & \equiv \frac{m c^{2}}{4 \pi e^{2} \lambda_{\mathrm{L}}^{2}}  \tag{2.222}\\
& =n-m \lim _{\boldsymbol{q} \rightarrow 0} \chi_{\perp}(\boldsymbol{q}, 0) .
\end{align*}
$$

## Ideal Bose Gas

We start from the susceptibility,

$$
\begin{equation*}
\chi_{i j}(\boldsymbol{q}, t)=\frac{i}{\hbar V}\left\langle\left[j_{i}(\boldsymbol{q}, t), j_{j}(-\boldsymbol{q}, 0)\right]\right\rangle \Theta(t) \tag{2.223}
\end{equation*}
$$

where the current operator is given by

$$
\begin{equation*}
j_{i}(\boldsymbol{q})=\frac{\hbar}{2 m} \sum_{\boldsymbol{k}}\left(2 k_{i}+q_{i}\right) \psi_{\boldsymbol{k}}^{\dagger} \psi_{\boldsymbol{k}+\boldsymbol{q}} . \tag{2.224}
\end{equation*}
$$

For the free Bose gas, with dispersion $\omega_{\boldsymbol{k}}=\hbar \boldsymbol{k}^{2} / 2 m$,

$$
\begin{equation*}
j_{i}(\boldsymbol{q}, t)=\left(2 k_{i}+q_{i}\right) e^{i\left(\omega_{\boldsymbol{k}}-\omega_{\boldsymbol{k}+\boldsymbol{q}}\right) t} \psi_{\boldsymbol{k}}^{\dagger} \psi_{\boldsymbol{k}+\boldsymbol{q}}, \tag{2.225}
\end{equation*}
$$

hence

$$
\begin{align*}
{\left[j_{i}(\boldsymbol{q}, t), j_{j}(-\boldsymbol{q}, 0)\right]=\frac{\hbar^{2}}{4 m^{2}} \sum_{\boldsymbol{k}, \boldsymbol{k}^{\prime}}\left(2 k_{i}+q_{i}\right)\left(2 k_{j}^{\prime}-q_{j}\right) e^{i\left(\omega_{\boldsymbol{k}}-\omega_{\boldsymbol{k}+\boldsymbol{q}}\right) t} }  \tag{2.226}\\
\times\left[\psi_{\boldsymbol{k}}^{\dagger} \psi_{\boldsymbol{k}+\boldsymbol{q}}, \psi_{\boldsymbol{k}^{\prime}}^{\dagger} \psi_{\boldsymbol{k}^{\prime}-\boldsymbol{q}}\right]
\end{align*}
$$

Using

$$
\begin{equation*}
[A B, C D]=A[B, C] D+A C[B, D]+C[A, D] B+[A, C] D B, \tag{2.227}
\end{equation*}
$$

we obtain

$$
\begin{equation*}
\left[j_{i}(\boldsymbol{q}, t), j_{j}(-\boldsymbol{q}, 0)\right]=\frac{\hbar^{2}}{4 m^{2}} \sum_{k}\left(2 k_{i}+q_{i}\right)\left(2 k_{j}+q_{j}\right) e^{i\left(\omega_{\boldsymbol{k}}-\omega_{\boldsymbol{k}+\boldsymbol{q}}\right) t}\left\{n^{0}\left(\omega_{\boldsymbol{k}}\right)-n^{0}\left(\omega_{\boldsymbol{k}+\boldsymbol{q}}\right)\right\}, \tag{2.228}
\end{equation*}
$$

where $n^{0}(\omega)$ is the equilibrium Bose distribution, ${ }^{5}$

$$
\begin{equation*}
n^{0}(\omega)=\frac{1}{e^{\beta \hbar \omega} e^{-\beta \mu}-1} . \tag{2.229}
\end{equation*}
$$

Thus,

$$
\begin{align*}
\chi_{i j}(\boldsymbol{q}, \omega)= & \frac{\hbar}{4 m^{2} V} \sum_{k}\left(2 k_{i}+q_{i}\right)\left(2 k_{j}+q_{j}\right) \frac{n^{0}\left(\omega_{\boldsymbol{k}+\boldsymbol{q}}\right)-n^{0}\left(\omega_{\boldsymbol{k}}\right)}{\omega+\omega_{\boldsymbol{k}}-\omega_{\boldsymbol{k}+\boldsymbol{q}}+i \epsilon} \\
= & \frac{\hbar n_{0}}{4 m^{2}}\left\{\frac{1}{\omega+\omega_{\boldsymbol{q}}+i \epsilon}-\frac{1}{\omega-\omega_{\boldsymbol{q}}+i \epsilon}\right\} q_{i} q_{j}  \tag{2.230}\\
& +\frac{\hbar}{m^{2}} \int \frac{d^{3} k}{(2 \pi)^{3}} \frac{n^{0}\left(\omega_{\boldsymbol{k}+\boldsymbol{q} / 2}\right)-n^{0}\left(\omega_{\boldsymbol{k}-\boldsymbol{q} / 2}\right)}{\omega+\omega_{\boldsymbol{k}-\boldsymbol{q} / 2}-\omega_{\boldsymbol{k}+\boldsymbol{q} / 2}+i \epsilon} k_{i} k_{j},
\end{align*}
$$

where $n_{0}=N_{0} / V$ is the condensate number density. Taking the $\omega=0$ and $\boldsymbol{q} \rightarrow 0$ limit yields

$$
\begin{equation*}
\chi_{i j}(\boldsymbol{q} \rightarrow 0,0)=\frac{n_{0}}{m} \hat{q}_{i} \hat{q}_{j}+\frac{n^{\prime}}{m} \delta_{i j}, \tag{2.231}
\end{equation*}
$$

where $n^{\prime}$ is the density of uncondensed bosons. From this we read off

$$
\begin{equation*}
\chi_{\|}(\boldsymbol{q} \rightarrow 0,0)=\frac{n}{m} \quad, \quad \chi_{\perp}(\boldsymbol{q} \rightarrow 0,0)=\frac{n^{\prime}}{m}, \tag{2.232}
\end{equation*}
$$

where $n=n_{0}+n^{\prime}$ is the total boson number density. The superfluid density, according to (2.222), is $n_{\mathrm{s}}=n_{0}(T)$. In fact, the ideal Bose gas is not a superfluid. Its excitation spectrum is too 'soft' - any superflow is unstable toward decay into single particle excitations.

[^3]
### 2.8 Electron-phonon Hamiltonian

Let $\boldsymbol{R}_{i}=\boldsymbol{R}_{i}^{0}+\delta \boldsymbol{R}_{i}$ denote the position of the $i^{\text {th }}$ ion, and let $U(\boldsymbol{r})=-Z e^{2} \exp \left(-r / \lambda_{\mathrm{TF}}\right) / r$ be the electron-ion interaction. Expanding in terms of the ionic displacements $\delta \boldsymbol{R}_{i}$,

$$
\begin{equation*}
\hat{H}_{\mathrm{el}-\mathrm{ion}}=\sum_{i} U\left(\boldsymbol{r}-\boldsymbol{R}_{i}^{0}\right)-\sum_{i} \delta \boldsymbol{R}_{i} \cdot \nabla U\left(\boldsymbol{r}-\boldsymbol{R}_{i}^{0}\right), \tag{2.233}
\end{equation*}
$$

where $i$ runs from 1 to $N_{\text {ion }}$, the number of ions ${ }^{6}$. The deviation $\delta \boldsymbol{R}_{i}$ may be expanded in terms of the vibrational normal modes of the lattice, i.e. the phonons, as

$$
\begin{equation*}
\delta R_{i}^{\alpha}=\frac{1}{\sqrt{N_{\mathrm{ion}}}} \sum_{\boldsymbol{q} \lambda}\left(\frac{\hbar}{2 \omega_{\lambda}(\boldsymbol{q})}\right)^{1 / 2} \hat{\mathrm{e}}_{\lambda}^{\alpha}(\boldsymbol{q}) e^{i \boldsymbol{q} \cdot \boldsymbol{R}_{i}^{0}}\left(a_{\boldsymbol{q} \lambda}+a_{-\boldsymbol{q} \lambda}^{\dagger}\right) \tag{2.234}
\end{equation*}
$$

The phonon polarization vectors satisfy $\hat{\boldsymbol{e}}_{\lambda}(\boldsymbol{q})=\hat{\boldsymbol{e}}_{\lambda}^{*}(-\boldsymbol{q})$ as well as the generalized orthonormality relations

$$
\begin{align*}
& \sum_{\alpha} \hat{\mathrm{e}}_{\lambda}^{\alpha}(\boldsymbol{q}) \hat{\mathrm{e}}_{\lambda^{\prime}}^{\alpha}(-\boldsymbol{q})=M^{-1} \delta_{\lambda \lambda^{\prime}}  \tag{2.235}\\
& \sum_{\lambda} \hat{e}_{\lambda}^{\alpha}(\boldsymbol{q}) \hat{e}_{\lambda}^{\beta}(-\boldsymbol{q})=M^{-1} \delta_{\alpha \beta}
\end{align*}
$$

where $M$ is the ionic mass. The number of unit cells in the crystal is $N_{\text {ion }}=V / \Omega$, where $\Omega$ is the Wigner-Seitz cell volume. Again, we approximate Bloch states by plane waves $\psi_{\boldsymbol{k}}(\boldsymbol{r})=\exp (i \boldsymbol{k} \cdot \boldsymbol{r}) / \sqrt{V}$, in which case

$$
\begin{equation*}
\left\langle\boldsymbol{k}^{\prime}\right| \nabla U\left(\boldsymbol{r}-\boldsymbol{R}_{i}^{0}\right)|\boldsymbol{k}\rangle=-\frac{i}{V} e^{i\left(\boldsymbol{k}-\boldsymbol{k}^{\prime}\right) \cdot \boldsymbol{R}_{i}^{0}} \frac{4 \pi Z e^{2}\left(\boldsymbol{k}-\boldsymbol{k}^{\prime}\right)}{\left(\boldsymbol{k}-\boldsymbol{k}^{\prime}\right)^{2}+\lambda_{\mathrm{TF}}^{-2}} . \tag{2.236}
\end{equation*}
$$

The sum over lattice sites gives

$$
\begin{equation*}
\sum_{i=1}^{N_{\text {ion }}} e^{i\left(\boldsymbol{k}-\boldsymbol{k}^{\prime}+\boldsymbol{q}\right) \cdot \boldsymbol{R}_{i}^{0}}=N_{\mathrm{ion}} \delta_{\boldsymbol{k}^{\prime}, \boldsymbol{k}+\boldsymbol{q} \bmod \boldsymbol{G}} \tag{2.237}
\end{equation*}
$$

so that

$$
\begin{equation*}
\hat{H}_{\mathrm{el}-\mathrm{ph}}=\frac{1}{\sqrt{V}} \sum_{\substack{k k^{\prime} \sigma \\ \boldsymbol{q} \lambda \boldsymbol{G}}} g_{\lambda}\left(\boldsymbol{k}, \boldsymbol{k}^{\prime}\right)\left(a_{\boldsymbol{q} \lambda}^{\dagger}+a_{-\boldsymbol{q} \lambda}\right) \psi_{\boldsymbol{k} \sigma}^{\dagger} \psi_{\boldsymbol{k}^{\prime} \sigma} \delta_{\boldsymbol{k}^{\prime}, \boldsymbol{k}+\boldsymbol{q}+\boldsymbol{G}} \tag{2.238}
\end{equation*}
$$

with

$$
\begin{equation*}
g_{\lambda}(\boldsymbol{k}, \boldsymbol{k}+\boldsymbol{q}+\boldsymbol{G})=-i\left(\frac{\hbar}{2 \Omega \omega_{\lambda}(\boldsymbol{q})}\right)^{1 / 2} \frac{4 \pi Z e^{2}}{(\boldsymbol{q}+\boldsymbol{G})^{2}+\lambda_{\mathrm{TF}}^{-2}}(\boldsymbol{q}+\boldsymbol{G}) \cdot \hat{\boldsymbol{e}}_{\lambda}^{*}(\boldsymbol{q}) . \tag{2.239}
\end{equation*}
$$

In an isotropic solid ${ }^{7}$ ('jellium'), the phonon polarization at wavevector $\boldsymbol{q}$ either is parallel to $\boldsymbol{q}$ (longitudinal waves), or perpendicular to $\boldsymbol{q}$ (transverse waves). We see that only longitudinal waves couple to the electrons. This is because transverse waves do not result in any local accumulation of charge density, and it is to the charge density that electrons couple, via the Coulomb interaction.

Restricting our attention to the longitudinal phonon, we have $\hat{\boldsymbol{e}}_{\mathrm{L}}(\boldsymbol{q})=\hat{\boldsymbol{q}} / \sqrt{M}$ and hence, for small $\boldsymbol{q}=\boldsymbol{k}^{\prime}-\boldsymbol{k}$,

$$
\begin{equation*}
g_{\mathrm{L}}(\boldsymbol{k}, \boldsymbol{k}+\boldsymbol{q})=-i\left(\frac{\hbar}{2 M \Omega}\right)^{1 / 2} \frac{4 \pi Z e^{2}}{q^{2}+\lambda_{\mathrm{TF}}^{-2}} c_{\mathrm{L}}^{-1 / 2} q^{1 / 2}, \tag{2.240}
\end{equation*}
$$

[^4]| Metal | $\Theta_{\mathrm{s}}$ | $\Theta_{\mathrm{D}}$ | $\lambda_{\text {el-ph }}$ | Metal | $\Theta_{\mathrm{s}}$ | $\Theta_{\mathrm{D}}$ | $\lambda_{\text {el-ph }}$ |
| :--- | ---: | ---: | :---: | :--- | ---: | ---: | :---: |
| Na | 220 | 150 | 0.47 | Au | 310 | 170 | 0.08 |
| K | 150 | 100 | 0.25 | Be | 1940 | 1000 | 0.59 |
| Cu | 490 | 315 | 0.16 | Al | 910 | 394 | 0.90 |
| Ag | 340 | 215 | 0.12 | In | 300 | 129 | 1.05 |

Table 2.1: Electron-phonon interaction parameters for some metals. Temperatures are in Kelvins.
where $c_{\mathrm{L}}$ is the longitudinal phonon velocity. Thus, for small $q$ we that the electron-longitudinal phonon coupling $g_{\mathrm{L}}(\boldsymbol{k}, \boldsymbol{k}+\boldsymbol{q}) \equiv g_{\boldsymbol{q}}$ satisfies

$$
\begin{equation*}
\left|g_{\boldsymbol{q}}\right|^{2}=\lambda_{\mathrm{el}-\mathrm{ph}} \cdot \frac{\hbar c_{\mathrm{L}} q}{g\left(\varepsilon_{\mathrm{F}}\right)}, \tag{2.241}
\end{equation*}
$$

where $g\left(\varepsilon_{\mathrm{F}}\right)$ is the electronic density of states, and where the dimensionless electron-phonon coupling constant is

$$
\begin{equation*}
\lambda_{\mathrm{el}-\mathrm{ph}}=\frac{Z^{2}}{2 M c_{\mathrm{L}}^{2} \Omega g\left(\varepsilon_{\mathrm{F}}\right)}=\frac{2 Z}{3} \frac{m^{*}}{M}\left(\frac{\varepsilon_{\mathrm{F}}}{k_{\mathrm{B}} \Theta_{\mathrm{s}}}\right)^{2}, \tag{2.242}
\end{equation*}
$$

with $\Theta_{\mathrm{S}} \equiv \hbar c_{\mathrm{L}} k_{\mathrm{F}} / k_{\mathrm{B}}$. Table 2.1 lists $\Theta_{\mathrm{S}}$, the Debye temperature $\Theta_{\mathrm{D}}$, and the electron-phonon coupling $\lambda_{\mathrm{el}-\mathrm{ph}}$ for various metals.


[^0]:    ${ }^{1}$ Note that $f(t)$ has dimensions of acceleration.

[^1]:    ${ }^{2}$ If not, define $\delta Q_{i} \equiv Q_{i}-\left\langle Q_{i}\right\rangle_{0}$ and consider $\left\langle\delta Q_{i}(t)\right\rangle$.

[^2]:    ${ }^{3}$ The spectral representation is sometimes known as the Lehmann representation.
    ${ }^{4}$ We assume all the $Q_{i}$ are Hermitian, i.e. $Q_{i}=Q_{i}^{\dagger}$.

[^3]:    ${ }^{5}$ Recall that $\mu=0$ in the condensed phase.

[^4]:    ${ }^{6}$ We assume a Bravais lattice, for simplicity.
    ${ }^{7}$ The jellium model ignores $\boldsymbol{G} \neq 0$ Umklapp processes.

