

141/241 Midterm Project: The “Mice” Collision



Units Chosen

$$L = 16.5 \text{ kpc}$$

$$M = 10^{11} M_{\text{Sun}}$$

$$t = 10^8 \text{ yrs}$$

$$dt = 10^7 \text{ yrs}$$

$$V = 16.5 \text{ kpc} / 10^8 \text{ yrs}$$

$$G = x L V^2 / M$$

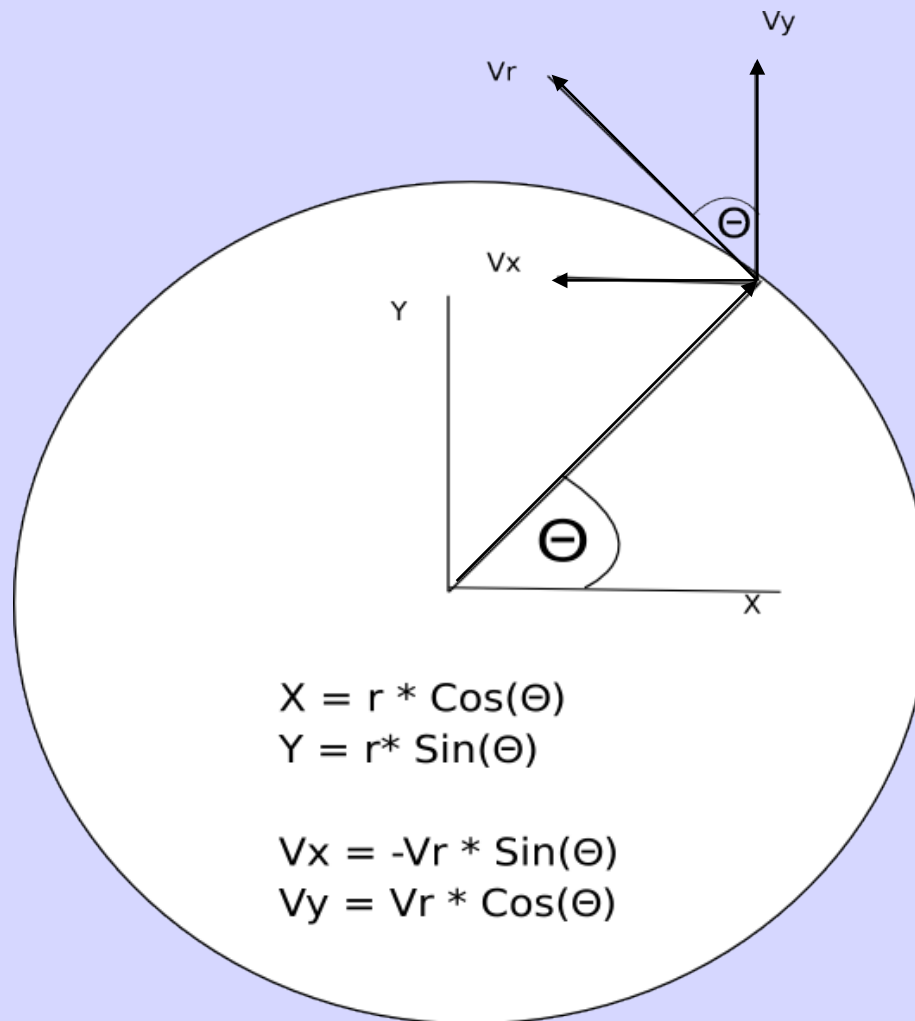
$$\rightarrow x = 1$$

$$G = L^3 / (t^2 M)$$

Galaxy Setup

- $R_{\min} = 25\text{kpc}$
- $10^{11} M_{\text{Sun}}$ at origin
- 11 concentric rings from $0.2 R_{\min} \rightarrow 0.7 R_{\min}$
- $dR_{\min} = 0.05 R_{\min}$
- nParticles from innermost to outermost ring =
 $\{12,15,18,21,24,27,30,33,36,39,42\}$
- Total Particles = 297
- vParticles = $\text{Sqrt}(GM/r)$ ← w.r.t. galaxy center

Coordinate Transformation



The Angles of Inclination

- Node-to-peri angles:

$$w_A = w_B = -90^\circ$$

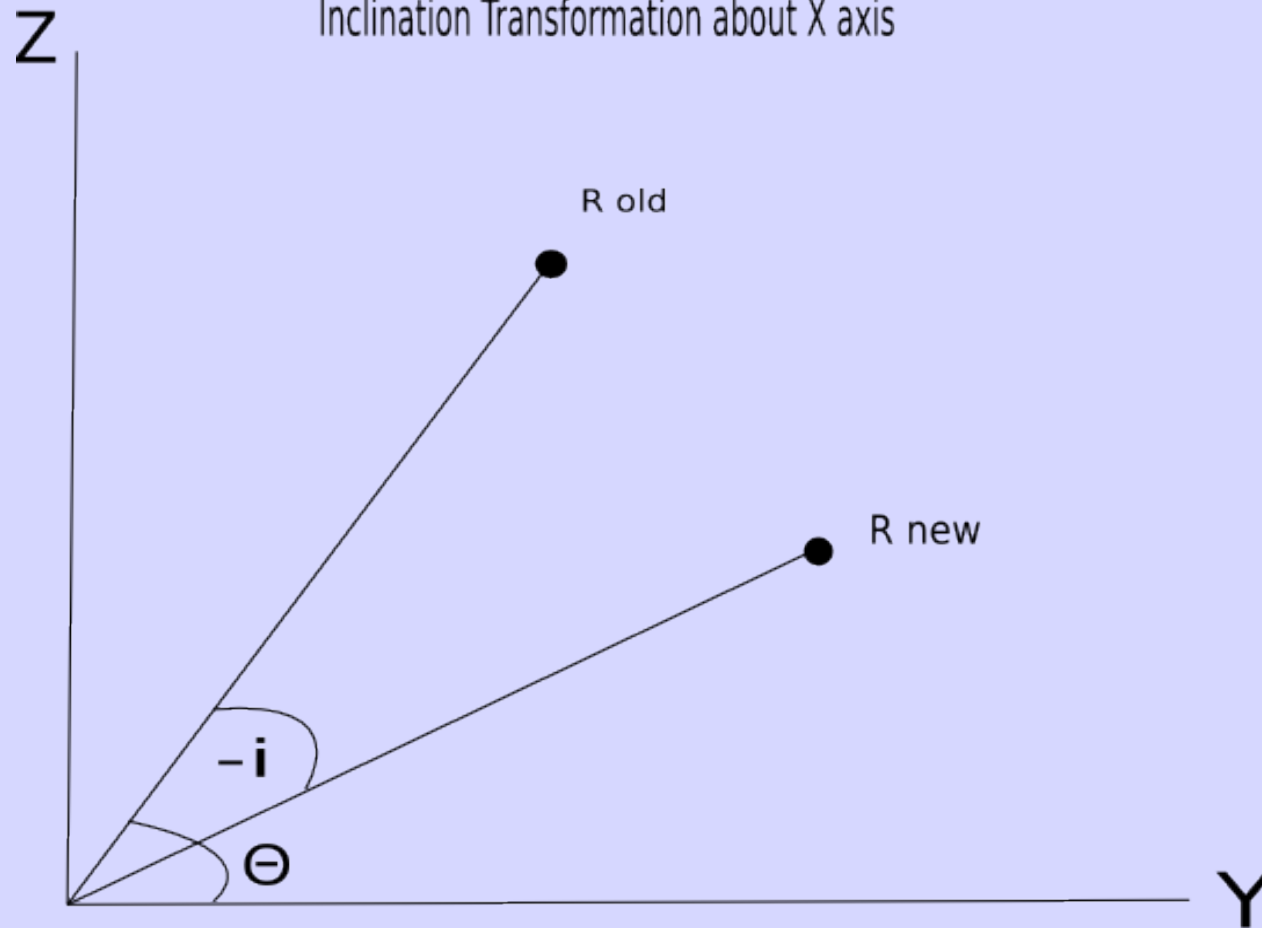
- Inclination angles:

$$i_A = 15^\circ$$

$$i_B = 60^\circ$$

Galaxy Orientation

Inclination Transformation about X axis



$$Z_{new} = Z_{old} - r * \sin(\Theta - i)$$

Since Θ is Zero initially, as well as Z_{old} , reduces to:

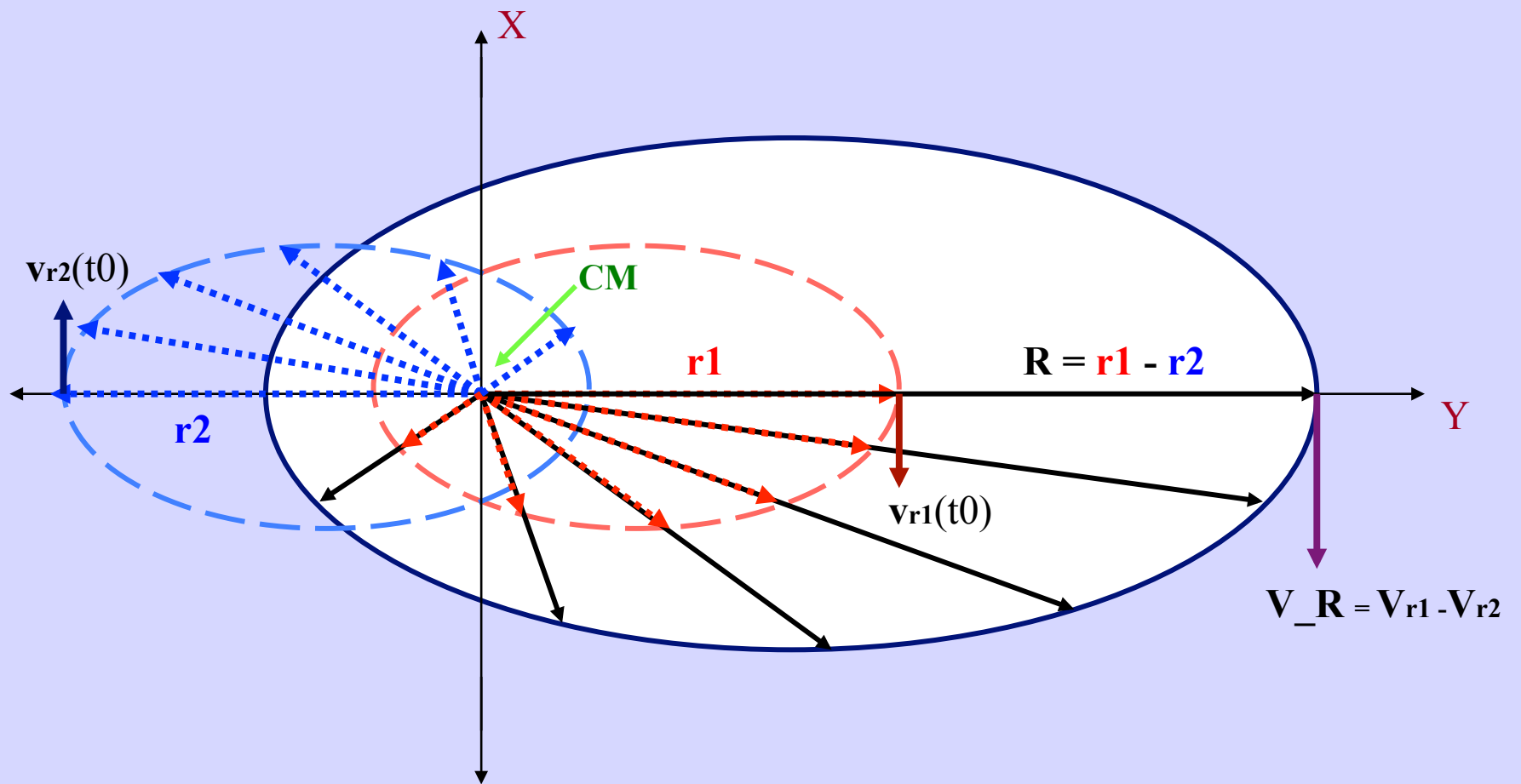
$$Z_{new} = -Y_{old} * \sin(-i)$$

$$Y_{new} = Y_{old} * \cos(i)$$

Equations for the C.M. Ellipse

- $k = G m_1 m_2$
- $a = L^2 / (m k) = 2 L^2$
- $m = m_1 m_2 / (m_1 + m_2) = M/2$
- $R = a / (1 + e \cos(q))$
- $L = m R^2 (dq/dt)$
- $dq/dt = L / R^2$
- $v_R = R^* (dq/dt) = 2 L / R$

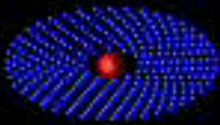
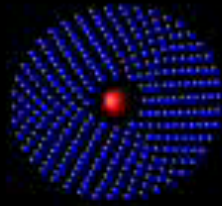
An Elliptic Encounter



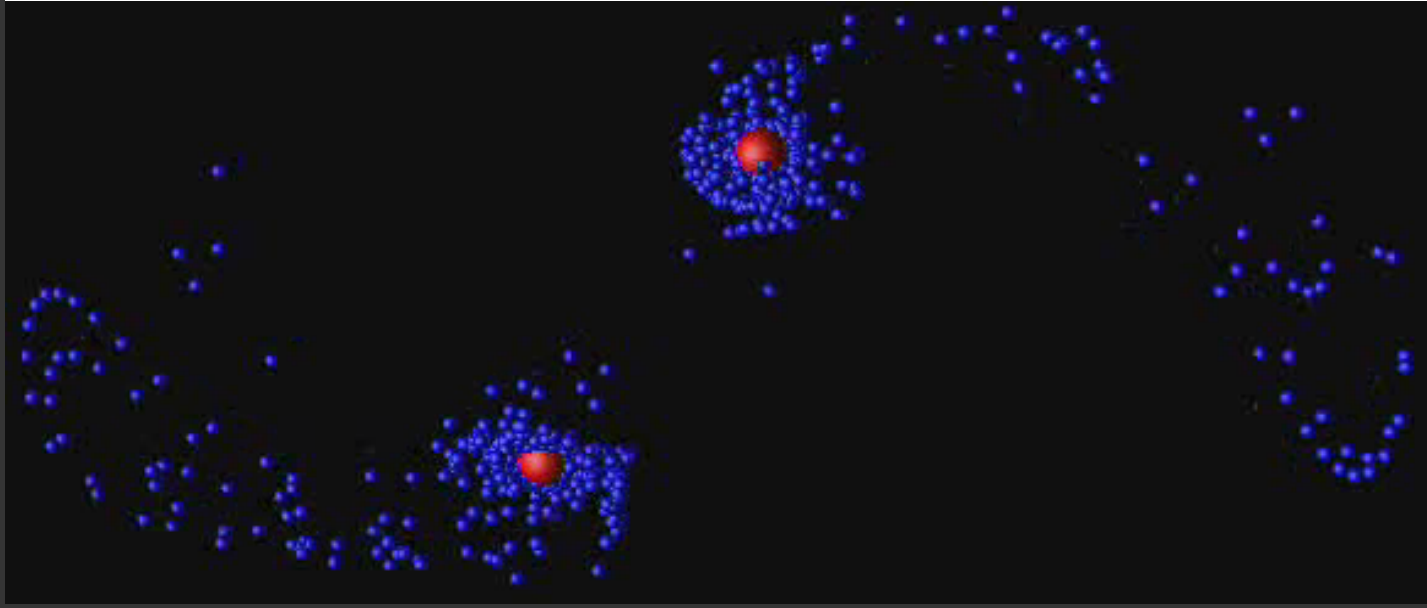
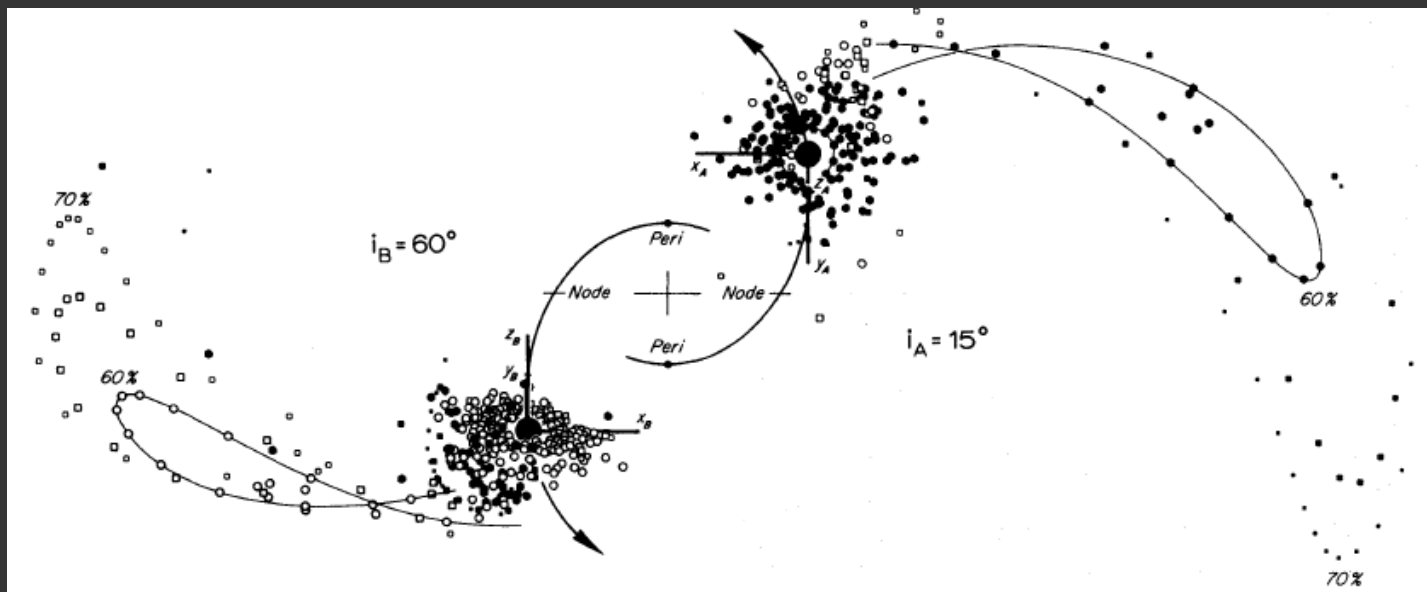
Equivalent 2-body Problem

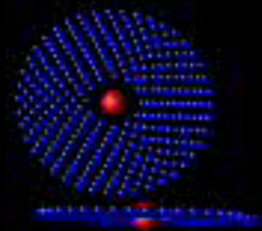
Since the masses of the
two galaxies are equal:

- $R = r_1 - r_2$
- $v_R = v_{r1} - v_{r2}$

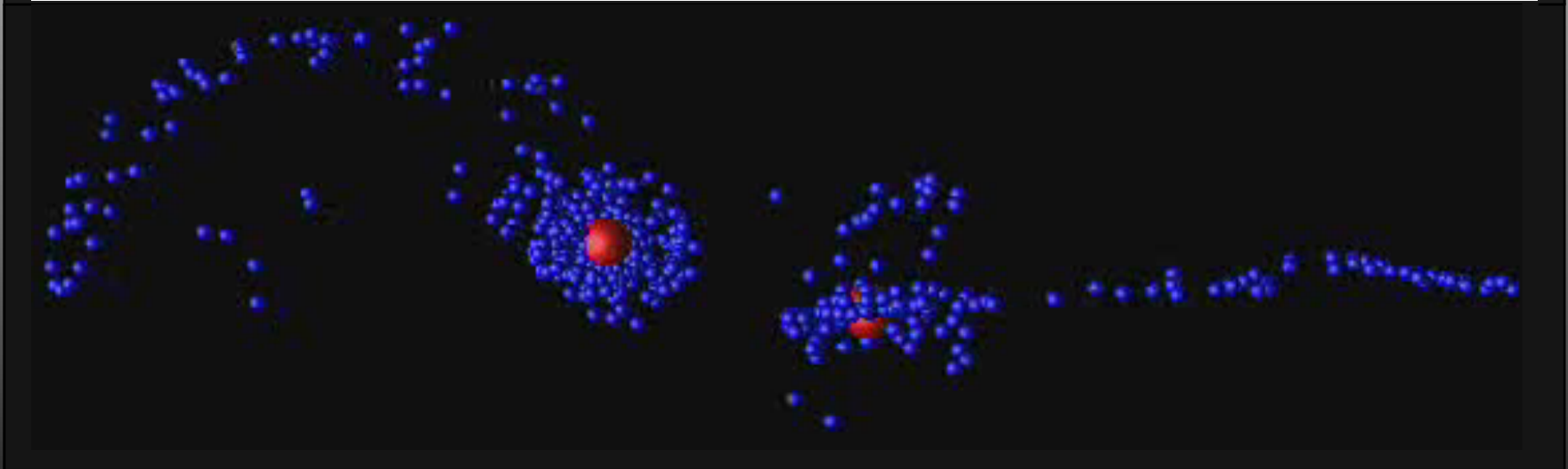
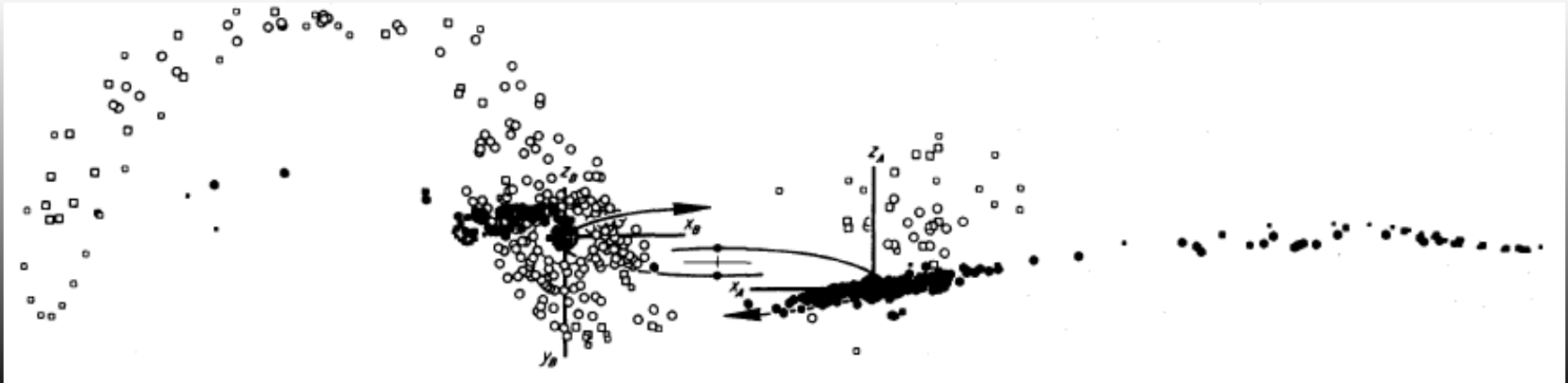


Top-View Comparison to Toomre & Toomre





Side-View Comparison to Toomre & Toomre



Explanation of Tidal Tails

- When the galaxies approach the pericenter of their encounter the massive centers are whipped around the C.M. more quickly than the stars on the outside of the collision path.

