


PHYS 2D
DISCUSSION SECTION

2012/4/18

- 
- Email me topics/questions you'd like to discuss
 - Problem section: Ch.2
 - Specific problems?
 - Pick up quiz 1 today & at problem section

Topics for Today



- Relativistic Momentum/Energy
- Black-Body Radiation

Relativistic Momentum

- Definition: relativistic momentum $\vec{p} \equiv \frac{m\vec{u}}{\sqrt{1 - \frac{u^2}{c^2}}} = \gamma m\vec{u}$
- m : mass of object at rest, \mathbf{u} : velocity of object
- Reason:
 1. We want the same laws of physics in all frames
 2. Old form ($\mathbf{p}=m\mathbf{u}$) is not conserved when we Lorentz transform \mathbf{u} (change of inertial frames)
 3. $\mathbf{p}=\gamma m\mathbf{u}$ gives energy & momentum conservation under Lorentz transformation
 4. Reverts back to classical form when $u \ll c$

Force & Kinetic Energy

Finding relativistic energy from relativistic p:

- Newton's 2nd Law: $\vec{F} = d\vec{p}/dt = ma(1 - u^2/c^2)^{-3/2}$
- Kinetic Energy K:

Work done in accelerating from velocity 0 to u

$$\begin{aligned} \square \text{ Work} &= \int_{x_1}^{x_2} \vec{F} \cdot d\vec{x} = K = \int_{x_1}^{x_2} \frac{m \frac{du}{dt}}{\left(1 - \frac{u^2}{c^2}\right)^{3/2}} dx \quad dx = u dt \\ &= \int_{t_1}^{t_2} \frac{m \frac{du}{dt}}{\left(1 - \frac{u^2}{c^2}\right)^{3/2}} u dt \quad du = \frac{du}{dt} dt \\ &= \int_0^u \frac{m u du}{\left(1 - \frac{u^2}{c^2}\right)^{3/2}} \quad y = 1 - \frac{u^2}{c^2}, \quad dy = -2 \frac{u}{c^2} du \\ &= \int_1^{1 - \frac{u^2}{c^2}} \frac{m \left(\frac{c^2}{-2}\right) dy}{y^{3/2}} = \left[y^{-1/2} mc^2 \right]_1^{1 - \frac{u^2}{c^2}} \\ &= \gamma mc^2 - mc^2 \end{aligned}$$

Relativistic Energy

- From $p = \gamma m u$, we get $K = \gamma m c^2 - m c^2$
- Define total relativistic energy $E \equiv K + m c^2 = \gamma m c^2$
- When object is at rest, $\gamma = 1$, $E = m c^2$
- Rest mass m is a measure of the object's energy at rest, which is the total internal energy
Ex. Heating up an object increases its rest mass
Ex. A compressed spring weighs more

Energy Conservation

Q: Why these definitions? $p = \gamma mu$, $E = \gamma mc^2$

A: With these forms relativistic energy & momentum are conserved in all frames, and we get classical results when $u \ll c$

□ $\sum_i E_i^{\text{before}} = \sum_i E_i^{\text{after}}$, $\sum_i P_i^{\text{before}} = \sum_i P_i^{\text{after}}$, i : all objects involved

Mass Energy Interchange

- Energy is converted into rest mass in an inelastic collision



- Classically kinetic energy is lost (turned into heat)
- Relativistic energy accounts for everything and is conserved, so

$$\gamma_1 mc^2 + \gamma_2 mc^2 = \gamma_3 Mc^2, \quad \gamma_1 = \gamma_2 = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad \gamma_3 = 1$$

- Rest mass of composite: $M = 2 \gamma_1 m > 2 m$
- M includes rest mass of 1 & 2 & internal energy (heat)
- Kinetic energy converted into rest mass (as heat)

Black-Body Radiation



- Significance: Led to discovery of quantum mechanics
- Blackbody: Anything that absorbs all incident light
- All blackbodies at the same temperature T give out the same spectrum, regardless of material/surface
- Implies something about light itself, independent of the body

Blackbody

- Blackbody example: the small hole (itself) in a cavity

Why?

1. Light entering through hole almost never comes out
 2. Light keeps bouncing inside cavity, so it will be absorbed completely with enough time
- Incident light is partially absorbed every time it hits the wall, while the wall emits a certain spectrum
 - Light inside cavity keeps interacting with wall and redistribute its frequency (energy) until both reach a state of thermal equilibrium
 - Distribution of energy=spectrum

Black-body Radiation

- Thermal equilibrium is characterized by temperature T
- Light leaking out through the hole is a portion of the light inside, and is by definition black-body radiation
- So spectrum of light inside cavity and exiting through hole is the blackbody spectrum, depending only on T
- Spectrum can be calculated with statistical mechanics

- Classical statistical mechanics gives wrong spectrum
- Planck's quantization of light gives correct spectrum
- Light is quantized
- The start of quantum mechanics



□ Questions?