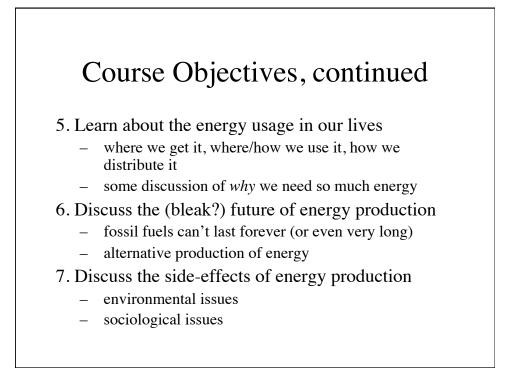


## Course Objectives, continued

- 3. Address Energy Conservation
  - loss-less exchange of energy between forms
  - never created or destroyed—just converted/exchanged
  - expose my ignorance/limitations of PowerPoint
- 4. Learn to calculate energy content/conversion
  - most quantitative part of course
  - forms foundation for all that follows
  - applications to familiar everyday systems
  - math isn't hard, but units can be a pain



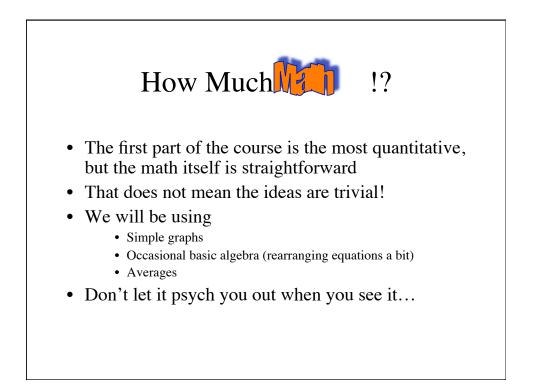
## **Course Structure**

- Class meeting times:
  - Lectures in York 2622: 2-2:50pm MWF
    - Homework due on Mondays
    - ABCD Card to promote interaction: bring to class every day; print from web site
    - Midterm: Mon May 9 (in class);
    - Final: Friday June 10, 3pm-6pm
  - Discussion section meets weekly (current day (Wed) may change to Friday
    - Opportunity for discussions on course material, exam prep, etc.
    - Math background and exercises
    - Work out example problems and questions

•	Your Fellow Students!
	<ul> <li>Encouraged to work together on homework, exercises. However, copying anything is cheating and I take cheating seriously.</li> </ul>
•	Professor : Kim Griest
	Office in SERF building, Room 337, office hours M 11-12 PM, or by appointment, 858-534-8914, or drop in any time (call first to see if I am there!)
•	Teaching Assistant: Joe Salamon: jsalamon@ucsd.edu Office hours TBA
•	Web: physics.ucsd.edu and follow links to course web page
•	Text:
	- Energy and the Environment, 2 <sup>nd</sup> edition Ristinen and Kraushaar

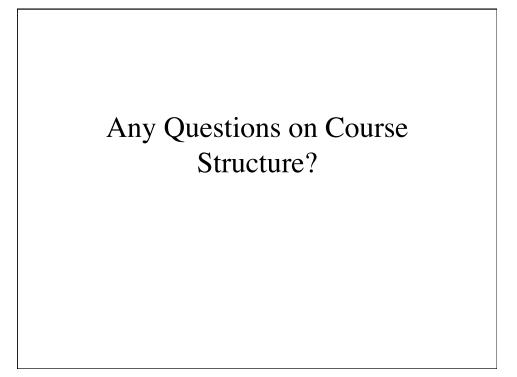
Grading						
Weight	Format	Due	Comments			
20%	Homework	Mon, in class	Graded 0-4, mostly on effort (& spot check)			
up to 7%	Class Particip.	every lecture	via hand in questions			
<30%*	Midterm*	May 9 (Mon)				
<50%*	Final*	June 10 (Fri)				

\* Midterm and Final may count for as little as 23% or 43%, respectively, given extent of classroom participation and worst exam performance. Example: if you have 90% of the participation credit and bomb the final, the midterm still counts 30%, but the final will Counts 6.3% less, or 43.7% instead of 50%, the rest made up by participation credit

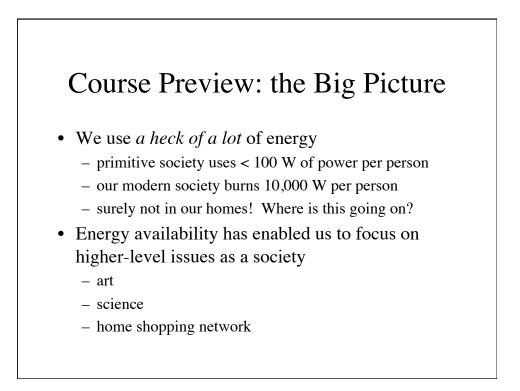


## Expectations

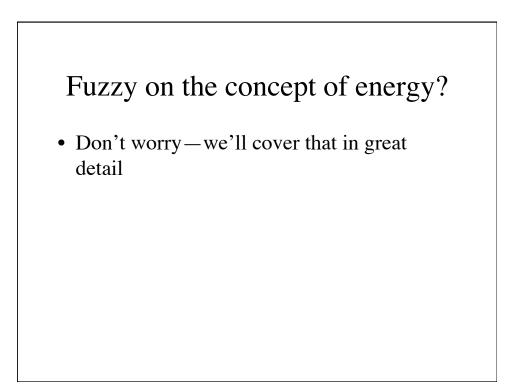
- Attend lectures and discussion section
- Participate!
  - If it doesn't make sense, **ask**! Everyone learns that way.
  - Don't be bashful about answering questions posed.
  - In-class voting system should make this fun
- Do the work:
  - It's the only way this stuff will really sink in
  - exams become easy
- Explore, think, ask, speculate, admire, enjoy!
  - Bring interesting topical ideas to class

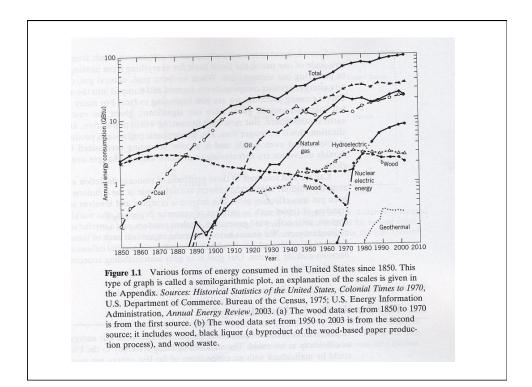


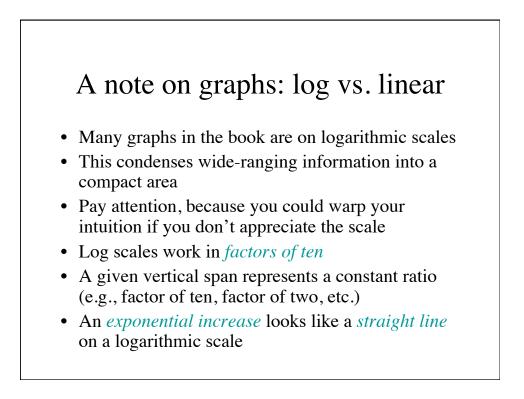


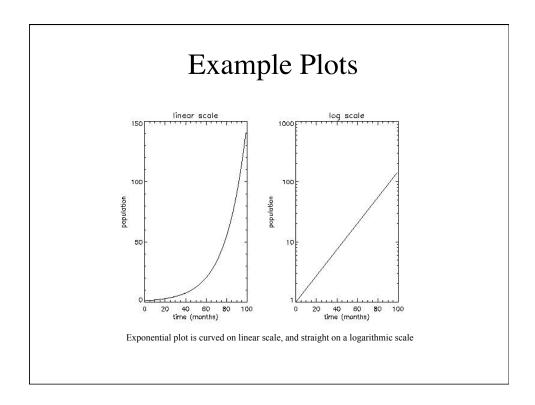


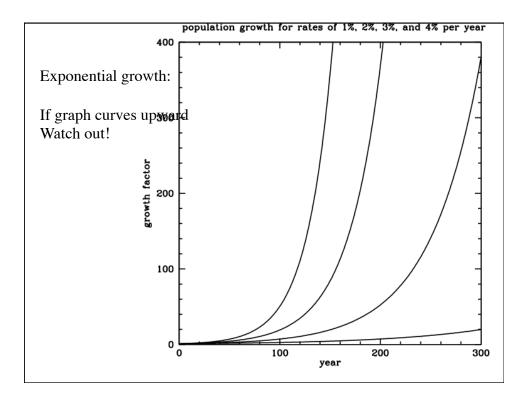
- Long ago, almost all of our energy came from food (delivering muscle power), and almost all our energy went into securing food for ourselves
- Enter the work animal, supplementing our muscle power and enabling larger-scale agriculture
- Next burn wood to run boilers, trains
- 150 years ago, muscular effort and firewood provided *most* of our energy—and today this is less than 1% of the story
- Today, much more energy *goes into* growing/ harvesting food than *comes out of* food!
- Today in US 86% of our energy comes from fossil fuels (oil,natural gas, coal)

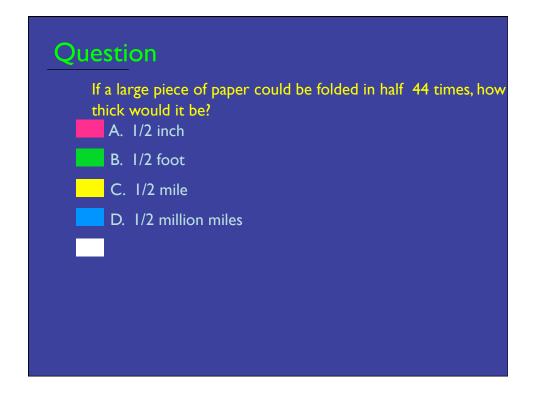


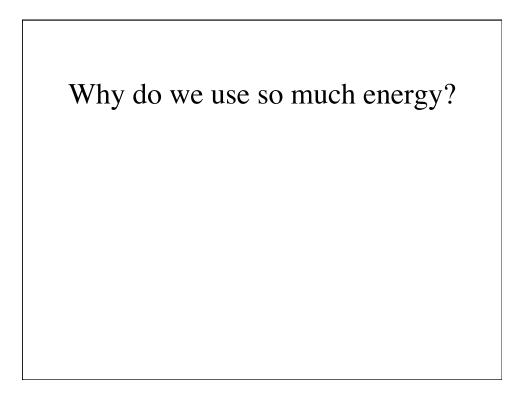


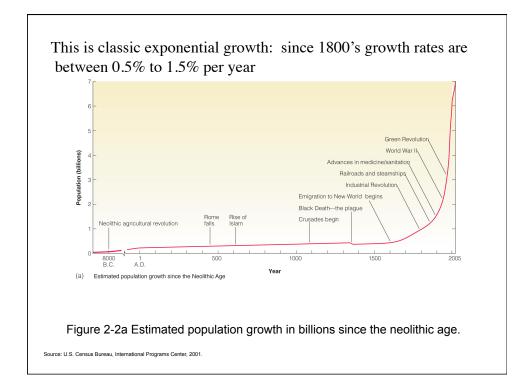


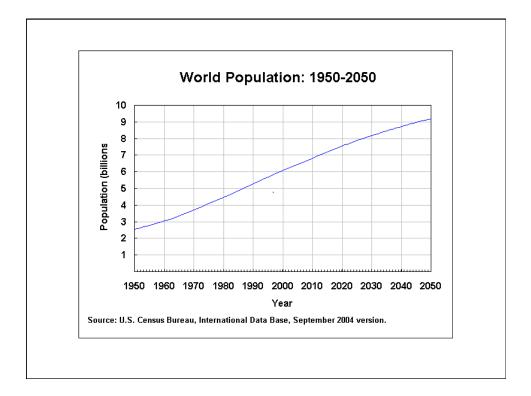












<ul> <li>Population History</li> <li>The total human population on the planet was small and increased slowly through most of human history. Maybe 1 million humans 125,000 years ago, growing only to 5-10 million by 10,000 years ago</li> <li>The population started to grow more rapidly due to human inventions: first agriculture, then industrial technology, and finally fertilizers and</li> </ul>						
mechanized agricultu - 1000 BCE - 0 CE - 1000 CE - 1500 CE - 1800 CE - 1930 CE - 1960 CE - 1974 CE - 1987 CE - 1987 CE - 2009 CE - 2010 CE	solution       50 million (0.05 billion)         0.15 billion       0.25 billion         0.25 billion (r < 0.1%: 8000BCE to 1650 CE: $t_{1/2}$ =1000yr)         1 billion (r = 0.46%: 1650CE to 1800 CE: $t_{1/2}$ =150yr)         2 billion (r = 0.54%: 1800CE to 1930 CE: $t_{1/2}$ =130yr)         3 billion (r = 2%: 1960's and 70's: $t_{1/2}$ =35yr)         4 billion         5 billion         6 billion         6.79 billion         6.9 billion					

Population increase depends upon: number born – number dying.						
Number born is related to TFR (total fertility rate) = average of number of children women bear:						
TFR=2.1 will keep population constant. TFR>2.1 $\Rightarrow$						
exponential increase						
To place this in perspective currently Roughly equivalent to population of:						
360,000 people born every day	Anaheim					
150,000 people die every day	Oceanside					
210,000 people added to the heap every day	Irvine					
AND, if population continued to double every $\sim$ 40 years (1%/year growth rate)						
<ul> <li>2050– 10 billion</li> <li>2100– 21 billion</li> <li>2200– 46 billion</li> <li>2300– 124 billion</li> </ul>						

