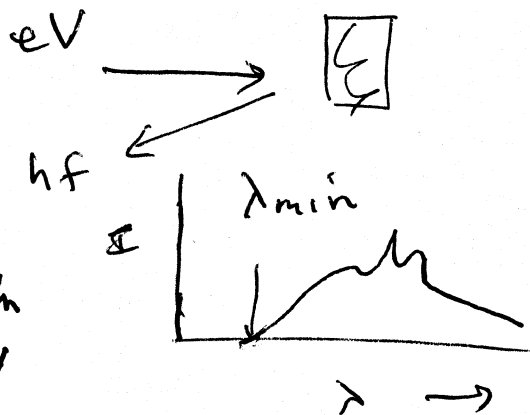


Physics 1C

Spring 2010

Quiz 3 form A



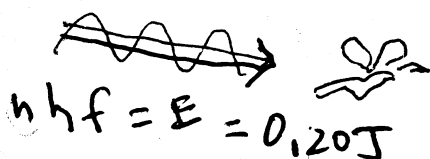
- 1) X ray spectrum -  
shortest wavelength  $\lambda_{\min}$   
all of the electron's energy  
goes to the photon -

$$E = eV = hf = h \frac{c}{\lambda}$$

$$\lambda = \frac{hc}{eV} = \frac{6.6 \times 10^{-34} \text{ J}\cdot\text{s} (3 \times 10^8 \text{ m/s})}{1.6 \times 10^{-19} \text{ J/V} (3000 \text{ V})} = 4.1 \times 10^{-10} \text{ m}$$

$$\lambda = \boxed{0.41 \text{ nm}}$$

2)



$$\lambda = 450 \text{ nm}$$

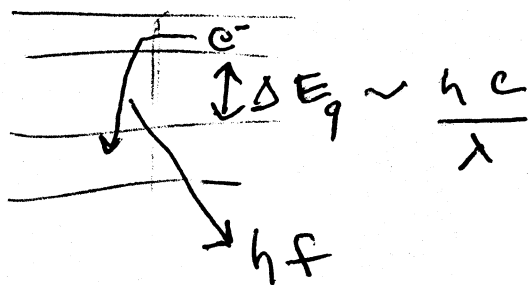
$$E = nhf = nhc \frac{1}{\lambda} = 0.20 \text{ J}$$

$$n = \frac{E\lambda}{hc} = \frac{0.20 \text{ J} (450 \times 10^{-9} \text{ m})}{6.6 \times 10^{-34} \text{ J}\cdot\text{s} (3 \times 10^8 \text{ m/s})}$$

$$n = 4.5 \times 10^{17} \sim \boxed{5 \times 10^{17}}$$

3)

$E \uparrow$

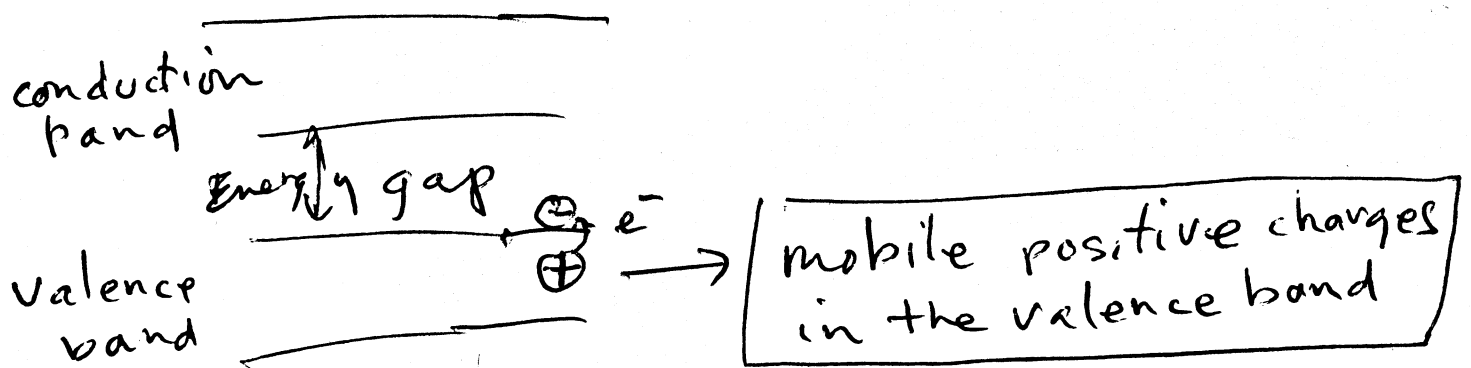


The photon energy  
increases with  $\Delta E_g$  -  
A larger  
(band gap) - gives a  
shorter wavelength.

4) For an electron in the  $n=3$  state.

				no of states
$n=3$	$l=0$	$m_l=0$	$m_s=\pm\frac{1}{2}$	2
	$l=1$	$m_l=-1, 0, 1$	$m_s=\pm\frac{1}{2}$	6
	$l=2$	$m_l=-2, -1, 0, 1, 2$	$m_s=\pm\frac{1}{2}$	10
			total	18

5) p-type semiconductor



6)  $V$  |  $e^-$   $\rightarrow$   $\lambda_{dB} = 0.03 \text{ nm}$

$\lambda = \frac{h}{p}$  de Broglie wavelength

$E = eV = \frac{1}{2} m v^2 = \frac{1}{2} \frac{m^2 v^2}{m} = \frac{1}{2} \frac{p^2}{m}$

$p = \sqrt{2meV}$

$\lambda = \frac{h}{\sqrt{2meV}}$

6)

$$\lambda^2 = \frac{h^2}{2meV}$$

$$V = \frac{h^2}{\lambda^2 (2me)}$$

$$V = \frac{(6.6 \times 10^{-34} \text{ J}\cdot\text{s})^2}{(0.03 \times 10^{-9} \text{ m})^2 (2)(9.1 \times 10^{-31} \text{ kg})(1.6 \times 10^{-19} \frac{\text{J}}{\text{V}})}$$

$$V = \boxed{1.66 \times 10^3 \text{ V}}$$

7) Hydrogen Emission spectrum -

Rydberg equation

n = ∞



n = 3

n = 2

n = 1

transition from n = ∞ to n = 3

$$\frac{1}{\lambda} = R \left( \frac{1}{n_{\text{final}}^2} - \frac{1}{n_{\text{initial}}^2} \right)$$

$$= R \left( \frac{1}{3^2} - \frac{1}{\infty^2} \right) = \frac{R}{9}$$

$$\lambda = \frac{9}{R} = \frac{9}{1.097 \times 10^7 \text{ m}^{-1}} = 8.2 \times 10^{-7} \text{ m}$$

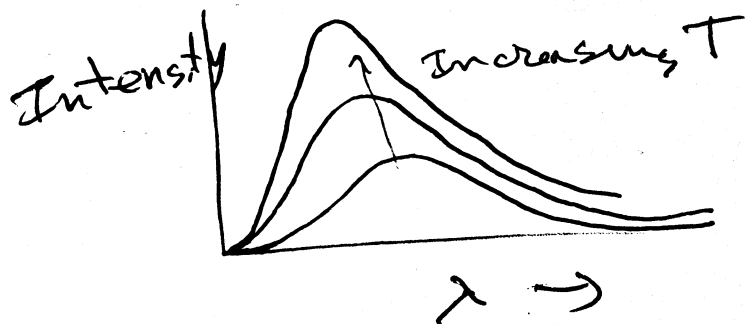
$$\lambda = \boxed{820 \text{ nm}}$$

8) Visible light is in the region from 400-700 nm  
 $\sim 500$  nm is typical.

$$E = h \frac{c}{\lambda} = \frac{6.6 \times 10^{-34} \text{ J}\cdot\text{s} \cdot 3 \times 10^8 \text{ m/s}}{500 \times 10^{-9} \text{ m} (1.6 \times 10^{-19} \text{ J/eV})}$$

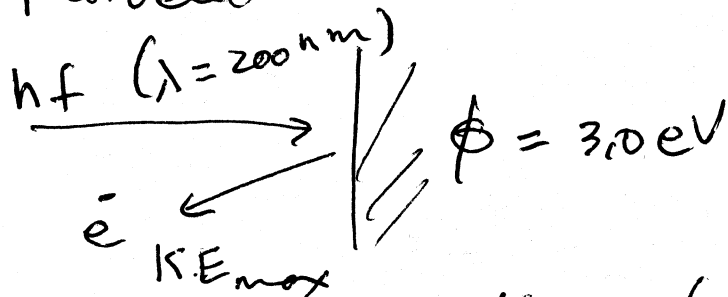
$$\approx 2.5 \text{ eV (closest to } \boxed{1 \text{ eV}})$$

9) Black Body Radiation



as temperature increases  $\lambda_{\text{max}}$  decreases and  
total intensity increases

10) Photoelectric Effect -



$$KE_{\text{max}} = hf - \phi \quad (\text{convert J to eV})$$

$$= \frac{h c}{\lambda e} - \phi \quad \text{divide by } e - \frac{1.6 \times 10^{-19} \text{ J}}{\text{eV}}$$

$$KE_{\text{max}} = \frac{6.6 \times 10^{-34} \text{ J}\cdot\text{s} (3 \times 10^8 \text{ m/s})}{200 \times 10^{-9} \text{ m} (1.6 \times 10^{-19} \text{ J/eV})} - 3.0 \text{ eV} = \boxed{3.2 \text{ eV}}$$