## PHYSICS 110A : CLASSICAL MECHANICS <br> FALL 2007 FINAL EXAM SOLUTIONS

[1] Two masses and two springs are configured linearly and externally driven to rotate with angular velocity $\omega$ about a fixed point on a horizontal surface, as shown in fig. 1. The unstretched length of each spring is $a$.


Figure 1: Two masses and two springs rotate with angular velocity $\omega$.
(a) Choose as generalized coordinates the radial distances $r_{1,2}$ from the origin. Find the Lagrangian $L\left(r_{1}, r_{2}, \dot{r}_{1}, \dot{r}_{2}, t\right)$.
[5 points]
The Lagrangian is

$$
\begin{equation*}
L=\frac{1}{2} m\left(\dot{r}_{1}^{2}+\dot{r}_{2}^{2}+\omega^{2} r_{1}^{2}+\omega^{2} r_{2}^{2}\right)-\frac{1}{2} k\left(r_{1}-a\right)^{2}-\frac{1}{2} k\left(r_{2}-r_{1}-a\right)^{2} . \tag{1}
\end{equation*}
$$

(b) Derive expressions for all conserved quantities.
[5 points]
The Hamiltonian is conserved. Since the kinetic energy is not homogeneous of degree 2 in the generalized velocities, $H \neq T+U$. Rather,

$$
\begin{align*}
H & =\sum_{\sigma} p_{\sigma} \dot{q}_{\sigma}-L  \tag{2}\\
& =\frac{1}{2} m\left(\dot{r}_{1}^{2}+\dot{r}_{2}^{2}\right)-\frac{1}{2} m \omega^{2}\left(r_{1}^{2}+r_{2}^{2}\right)+\frac{1}{2} k\left(r_{1}-a\right)^{2}+\frac{1}{2} k\left(r_{2}-r_{1}-a\right)^{2} . \tag{3}
\end{align*}
$$

We could define an effective potential

$$
\begin{equation*}
U_{\mathrm{eff}}\left(r_{1}, r_{2}\right)=-\frac{1}{2} m \omega^{2}\left(r_{1}^{2}+r_{2}^{2}\right)+\frac{1}{2} k\left(r_{1}-a\right)^{2}+\frac{1}{2} k\left(r_{2}-r_{1}-a\right)^{2} . \tag{4}
\end{equation*}
$$

Note the first term, which comes from the kinetic energy, has an interpretation of a fictitious potential which generates a centrifugal force.
(c) What equations determine the equilibrium radii $r_{1}^{0}$ and $r_{2}^{0}$ ? (You do not have to solve these equations.)
[5 points]
The equations of equilibrium are $F_{\sigma}=0$. Thus,

$$
\begin{align*}
& 0=F_{1}=\frac{\partial L}{\partial r_{1}}=m \omega^{2} r_{1}-k\left(r_{1}-a\right)+k\left(r_{2}-r_{1}-a\right)  \tag{5}\\
& 0=F_{2}=\frac{\partial L}{\partial r_{2}}=m \omega^{2} r_{2}-k\left(r_{2}-r_{1}-a\right) . \tag{6}
\end{align*}
$$

(d) Suppose now that the system is not externally driven, and that the angular coordinate $\phi$ is a dynamical variable like $r_{1}$ and $r_{2}$. Find the Lagrangian $L\left(r_{1}, r_{2}, \phi, \dot{r}_{1}, \dot{r}_{2}, \dot{\phi}, t\right)$. [5 points]

Now we have

$$
\begin{equation*}
L=\frac{1}{2} m\left(\dot{r}_{1}^{2}+\dot{r}_{2}^{2}+r_{1}^{2} \dot{\phi}^{2}+r_{2}^{2} \dot{\phi}^{2}\right)-\frac{1}{2} k\left(r_{1}-a\right)^{2}-\frac{1}{2} k\left(r_{2}-r_{1}-a\right)^{2} . \tag{7}
\end{equation*}
$$

(e) For the system described in part (d), find expressions for all conserved quantities.
[5 points]
There are two conserved quantities. One is $p_{\phi}$, owing to the fact the $\phi$ is cyclic in the Lagrangian. I.e. $\phi \rightarrow \phi+\zeta$ is a continuous one-parameter coordinate transformation which leaves $L$ invariant. We have

$$
\begin{equation*}
p_{\phi}=\frac{\partial L}{\partial \dot{\phi}}=m\left(r_{1}^{2}+r_{2}^{2}\right) \dot{\phi} . \tag{8}
\end{equation*}
$$

The second conserved quantity is the Hamiltonian, which is now $H=T+U$, since $T$ is homogeneous of degree 2 in the generalized velocities. Using conservation of momentum, we can write

$$
\begin{equation*}
H=\frac{1}{2} m\left(\dot{r}_{1}^{2}+\dot{r}_{2}^{2}\right)+\frac{p_{\phi}^{2}}{2 m\left(r_{1}^{2}+r_{2}^{2}\right)}+\frac{1}{2} k\left(r_{1}-a\right)^{2}+\frac{1}{2} k\left(r_{2}-r_{1}-a\right)^{2} . \tag{9}
\end{equation*}
$$

Once again, we can define an effective potential,

$$
\begin{equation*}
U_{\mathrm{eff}}\left(r_{1}, r_{2}\right)=\frac{p_{\phi}^{2}}{2 m\left(r_{1}^{2}+r_{2}^{2}\right)}+\frac{1}{2} k\left(r_{1}-a\right)^{2}+\frac{1}{2} k\left(r_{2}-r_{1}-a\right)^{2} \tag{10}
\end{equation*}
$$

which is different than the effective potential from part (b). However in both this case and in part (b), we have that the radial coordinates obey the equations of motion

$$
\begin{equation*}
m \ddot{r}_{j}=-\frac{\partial U_{\mathrm{eff}}}{\partial r_{j}} \tag{11}
\end{equation*}
$$

for $j=1,2$. Note that this equation of motion follows directly from $\dot{H}=0$.
[2] A point mass $m$ slides inside a hoop of radius $R$ and mass $M$, which itself rolls without slipping on a horizontal surface, as depicted in fig. 2.


Figure 2: A mass point $m$ rolls inside a hoop of mass $M$ and radius $R$ which rolls without slipping on a horizontal surface.

Choose as general coordinates $(X, \phi, r)$, where $X$ is the horizontal location of the center of the hoop, $\phi$ is the angle the mass $m$ makes with respect to the vertical ( $\phi=0$ at the bottom of the hoop), and $r$ is the distance of the mass $m$ from the center of the hoop. Since the mass $m$ slides inside the hoop, there is a constraint:

$$
G(X, \phi, r)=r-R=0 .
$$

Nota bene: The kinetic energy of the moving hoop, including translational and rotational components (but not including the mass $m$ ), is $T_{\text {hoop }}=M \dot{X}^{2}$ (i.e. twice the translational contribution alone).
(a) Find the Lagrangian $L(X, \phi, r, \dot{X}, \dot{\phi}, \dot{r}, t)$.
[5 points]
The Cartesian coordinates and velocities of the mass $m$ are

$$
\begin{array}{ll}
x=X+r \sin \phi & \dot{x}=\dot{X}+\dot{r} \sin \phi+r \dot{\phi} \cos \phi \\
y=R-r \cos \phi & \dot{y}=-\dot{r} \cos \phi+r \dot{\phi} \sin \phi \tag{13}
\end{array}
$$

The Lagrangian is then

$$
\begin{equation*}
L=\overbrace{\left(M+\frac{1}{2} m\right) \dot{X}^{2}+\frac{1}{2} m\left(\dot{r}^{2}+r^{2} \dot{\phi}^{2}\right)+m \dot{X}(\dot{r} \sin \phi+r \dot{\phi} \cos \phi)}^{T}-\overbrace{m g(R-r \cos \phi)}^{U} \tag{14}
\end{equation*}
$$

Note that we are not allowed to substitute $r=R$ and hence $\dot{r}=0$ in the Lagrangian prior to obtaining the equations of motion. Only after the generalized momenta and forces are computed are we allowed to do so.
(b) Find all the generalized momenta $p_{\sigma}$, the generalized forces $F_{\sigma}$, and the forces of constraint $Q_{\sigma}$.

```
[10 points]
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The generalized momenta are

$$
\begin{align*}
p_{r} & =\frac{\partial L}{\partial \dot{r}}=m \dot{r}+m \dot{X} \sin \phi  \tag{15}\\
p_{X} & =\frac{\partial L}{\partial \dot{X}}=(2 M+m) \dot{X}+m \dot{r} \sin \phi+m r \dot{\phi} \cos \phi  \tag{16}\\
p_{\phi} & =\frac{\partial L}{\partial \dot{\phi}}=m r^{2} \dot{\phi}+m r \dot{X} \cos \phi \tag{17}
\end{align*}
$$

The generalized forces and the forces of constraint are

$$
\begin{array}{ll}
F_{r}=\frac{\partial L}{\partial r}=m r \dot{\phi}^{2}+m \dot{X} \dot{\phi} \cos \phi+m g \cos \phi & Q_{r}=\lambda \frac{\partial G}{\partial r}=\lambda \\
F_{X}=\frac{\partial L}{\partial X}=0 & Q_{X}=\lambda \frac{\partial G}{\partial X}=0 \\
F_{\phi}=\frac{\partial L}{\partial \phi}=m \dot{X} \dot{r} \cos \phi-m \dot{X} \dot{\phi} \sin \phi-m g r \sin \phi & Q_{\phi}=\lambda \frac{\partial G}{\partial \phi}=0 \tag{20}
\end{array}
$$

The equations of motion are

$$
\begin{equation*}
\dot{p}_{\sigma}=F_{\sigma}+Q_{\sigma} . \tag{21}
\end{equation*}
$$

At this point, we can legitimately invoke the constraint $r=R$ and set $\dot{r}=0$ in all the $p_{\sigma}$ and $F_{\sigma}$.
(c) Derive expressions for all conserved quantities.
[5 points]
There are two conserved quantities, which each derive from continuous invariances of the Lagrangian which respect the constraint. The first is the total momentum $p_{X}$ :

$$
\begin{equation*}
F_{X}=0 \quad \Longrightarrow \quad P \equiv p_{X}=\text { constant } \tag{22}
\end{equation*}
$$

The second conserved quantity is the Hamiltonian, which in this problem turns out to be the total energy $E=T+U$. Incidentally, we can use conservation of $P$ to write the energy in terms of the variable $\phi$ alone. From

$$
\begin{equation*}
\dot{X}=\frac{P}{2 M+m}-\frac{m R \cos \phi}{2 M+m} \dot{\phi} \tag{23}
\end{equation*}
$$

we obtain

$$
\begin{align*}
E & =\frac{1}{2}(2 M+m) \dot{X}^{2}+\frac{1}{2} m R^{2} \dot{\phi}^{2}+m R \dot{X} \dot{\phi} \cos \phi+m g R(1-\cos \phi) \\
& =\frac{\alpha P^{2}}{2 m(1+\alpha)}+\frac{1}{2} m R^{2}\left(\frac{1+\alpha \sin ^{2} \phi}{1+\alpha}\right) \dot{\phi}^{2}+m g R(1-\cos \phi), \tag{24}
\end{align*}
$$

where we've defined the dimensionless ratio $\alpha \equiv m / 2 M$. It is convenient to define the quantity

$$
\begin{equation*}
\Omega^{2} \equiv\left(\frac{1+\alpha \sin ^{2} \phi}{1+\alpha}\right) \dot{\phi}^{2}+2 \omega_{0}^{2}(1-\cos \phi), \tag{25}
\end{equation*}
$$

with $\omega_{0} \equiv \sqrt{g / R}$. Clearly $\Omega^{2}$ is conserved, as it is linearly related to the energy $E$ :

$$
\begin{equation*}
E=\frac{\alpha P^{2}}{2 m(1+\alpha)}+\frac{1}{2} m R^{2} \Omega^{2} \tag{26}
\end{equation*}
$$

(d) Derive a differential equation of motion involving the coordinate $\phi(t)$ alone. I.e. your equation should not involve $r, X$, or the Lagrange multiplier $\lambda$.
[5 points]
From conservation of energy,

$$
\begin{equation*}
\frac{d\left(\Omega^{2}\right)}{d t}=0 \quad \Longrightarrow \quad\left(\frac{1+\alpha \sin ^{2} \phi}{1+\alpha}\right) \ddot{\phi}+\left(\frac{\alpha \sin \phi \cos \phi}{1+\alpha}\right) \dot{\phi}^{2}+\omega_{0}^{2} \sin \phi=0, \tag{27}
\end{equation*}
$$

again with $\alpha=m / 2 M$. Incidentally, one can use these results in eqns. 25 and 27 to eliminate $\dot{\phi}$ and $\ddot{\phi}$ in the expression for the constraint force, $Q_{r}=\lambda=\dot{p}_{r}-F_{r}$. One finds

$$
\begin{align*}
\lambda & =-m R \frac{\dot{\phi}^{2}+\omega_{0}^{2} \cos \phi}{1+\alpha \sin ^{2} \phi} \\
& =-\frac{m R \omega_{0}^{2}}{\left(1+\alpha \sin ^{2} \phi\right)^{2}}\left\{(1+\alpha)\left(\frac{\Omega^{2}}{\omega_{0}^{2}}-4 \sin ^{2}\left(\frac{1}{2} \phi\right)\right)+\left(1+\alpha \sin ^{2} \phi\right) \cos \phi\right\} . \tag{28}
\end{align*}
$$

This last equation can be used to determine the angle of detachment, where $\lambda$ vanishes and the mass $m$ falls off the inside of the hoop. This is because the hoop can only supply a repulsive normal force to the mass $m$. This was worked out in detail in my lecture notes on constrained systems.
[3] Two objects of masses $m_{1}$ and $m_{2}$ move under the influence of a central potential $U=k\left|\boldsymbol{r}_{1}-\boldsymbol{r}_{2}\right|^{1 / 4}$.
(a) Sketch the effective potential $U_{\text {eff }}(r)$ and the phase curves for the radial motion. Identify for which energies the motion is bounded.
[5 points]


Figure 3: The effective $U_{\text {eff }}(r)=E_{0} \mathcal{U}_{\text {eff }}\left(r / r_{0}\right)$, where $r_{0}$ and $E_{0}$ are the radius and energy of the circular orbit.

The effective potential is

$$
\begin{equation*}
U_{\mathrm{eff}}(r)=\frac{\ell^{2}}{2 \mu r^{2}}+k r^{n} \tag{29}
\end{equation*}
$$

with $n=\frac{1}{4}$. In sketching the effective potential, I have rendered it in dimensionless form,

$$
\begin{equation*}
U_{\mathrm{eff}}(r)=E_{0} \mathcal{U}_{\mathrm{eff}}\left(r / r_{0}\right), \tag{30}
\end{equation*}
$$

where $r_{0}=\left(\ell^{2} / n k \mu\right)^{(n+2)^{-1}}$ and $E_{0}=\left(\frac{1}{2}+\frac{1}{n}\right) \ell^{2} / \mu r_{0}^{2}$, which are obtained from the results of part (b). One then finds

$$
\begin{equation*}
\mathcal{U}_{\mathrm{eff}}(x)=\frac{n x^{-2}+2 x^{n}}{n+2} \tag{31}
\end{equation*}
$$

Although it is not obvious from the detailed sketch in fig. 3, the effective potential does diverge, albeit slowly, for $r \rightarrow \infty$. Clearly it also diverges for $r \rightarrow 0$. Thus, the relative coordinate motion is bounded for all energies; the allowed energies are $E \geq E_{0}$.
(b) What is the radius $r_{0}$ of the circular orbit? Is it stable or unstable? Why?
[5 points]
For the general power law potential $U(r)=k r^{n}$, with $n k>0$ (attractive force), setting $U_{\text {eff }}^{\prime}\left(r_{0}\right)=0$ yields

$$
\begin{equation*}
-\frac{\ell^{2}}{\mu r_{0}^{3}}+n k r_{0}^{n-1}=0 \tag{32}
\end{equation*}
$$

Thus,

$$
\begin{equation*}
r_{0}=\left(\frac{\ell^{2}}{n k \mu}\right)^{\frac{1}{n+2}}=\left(\frac{4 \ell^{2}}{k \mu}\right)^{\frac{4}{9}} \tag{33}
\end{equation*}
$$

The orbit $r(t)=r_{0}$ is stable because the effective potential has a local minimum at $r=r_{0}$, i.e. $U_{\text {eff }}^{\prime \prime}\left(r_{0}\right)>0$. This is obvious from inspection of the graph of $U_{\text {eff }}(r)$ but can also be computed explicitly:

$$
\begin{align*}
U_{\mathrm{eff}}^{\prime \prime}\left(r_{0}\right) & =\frac{3 \ell^{2}}{\mu r_{0}^{4}}+n(n-1) k r_{0}^{n} \\
& =(n+2) \frac{\ell^{2}}{\mu r_{0}^{4}} \tag{34}
\end{align*}
$$

Thus, provided $n>-2$ we have $U_{\text {eff }}^{\prime \prime}\left(r_{0}\right)>0$.
(c) For small perturbations about a circular orbit, the radial coordinate oscillates between two values. Suppose we compare two systems, with $\ell^{\prime} / \ell=2$, but $\mu^{\prime}=\mu$ and $k^{\prime}=k$. What is the ratio $\omega^{\prime} / \omega$ of their frequencies of small radial oscillations?
[5 points]
From the radial coordinate equation $\mu \ddot{r}=-U_{\text {eff }}^{\prime}(r)$, we expand $r=r_{0}+\eta$ and find

$$
\begin{equation*}
\mu \ddot{\eta}=-U_{\mathrm{eff}}^{\prime \prime}\left(r_{0}\right) \eta+\mathcal{O}\left(\eta^{2}\right) . \tag{35}
\end{equation*}
$$

The radial oscillation frequency is then

$$
\begin{equation*}
\omega=(n+2)^{1 / 2} \frac{\ell}{\mu r_{0}^{2}}=(n+2)^{1 / 2} n^{\frac{2}{n+2}} k^{\frac{2}{n+2}} \mu^{-\frac{n}{n+2}} \ell^{\frac{n-2}{n+2}} . \tag{36}
\end{equation*}
$$

The $\ell$ dependence is what is key here. Clearly

$$
\begin{equation*}
\frac{\omega^{\prime}}{\omega}=\left(\frac{\ell^{\prime}}{\ell}\right)^{\frac{n-2}{n+2}} \tag{37}
\end{equation*}
$$

In our case, with $n=\frac{1}{4}$, we have $\omega \propto \ell^{-7 / 9}$ and thus

$$
\begin{equation*}
\frac{\omega^{\prime}}{\omega}=2^{-7 / 9} \tag{38}
\end{equation*}
$$

(d) Find the equation of the shape of the slightly perturbed circular orbit: $r(\phi)=r_{0}+\eta(\phi)$. That is, find $\eta(\phi)$. Sketch the shape of the orbit.
[5 points]
We have that $\eta(\phi)=\eta_{0} \cos \left(\beta \phi+\delta_{0}\right)$, with

$$
\begin{equation*}
\beta=\frac{\omega}{\dot{\phi}}=\frac{\mu r_{0}^{2}}{\ell} \cdot \omega=\sqrt{n+2} . \tag{39}
\end{equation*}
$$

With $n=\frac{1}{4}$, we have $\beta=\frac{3}{2}$. Thus, the radial coordinate makes three oscillations for every two rotations. The situation is depicted in fig. 4.


Figure 4: Radial oscillations with $\beta=\frac{3}{2}$.
(e) What value of $n$ would result in a perturbed orbit shaped like that in fig. 5?
[5 points]


Figure 5: Closed precession in a central potential $U(r)=k r^{n}$.
Clearly $\beta=\sqrt{n+2}=4$, in order that $\eta(\phi)=\eta_{0} \cos \left(\beta \phi+\delta_{0}\right)$ executes four complete periods over the interval $\phi \in[0,2 \pi]$. This means $n=14$.
[4] Two masses and three springs are arranged as shown in fig. 6. You may assume that in equilibrium the springs are all unstretched with length $a$. The masses and spring constants are simple multiples of fundamental values, viz.

$$
\begin{equation*}
m_{1}=m \quad, \quad m_{2}=4 m \quad, \quad k_{1}=k \quad, \quad k_{2}=4 k \quad, \quad k_{3}=28 k \tag{40}
\end{equation*}
$$



Figure 6: Coupled masses and springs.
(a) Find the Lagrangian.
[5 points]
Choosing displacements relative to equilibrium as our generalized coordinates, we have

$$
\begin{equation*}
T=\frac{1}{2} m \dot{\eta}_{1}^{2}+2 m \dot{\eta}_{2}^{2} \tag{41}
\end{equation*}
$$

and

$$
\begin{equation*}
U=\frac{1}{2} k \eta_{1}^{2}+2 k\left(\eta_{2}-\eta_{1}\right)^{2}+14 k \eta_{2}^{2} . \tag{42}
\end{equation*}
$$

Thus,

$$
\begin{equation*}
L=T-U=\frac{1}{2} m \dot{\eta}_{1}^{2}+2 m \dot{\eta}_{2}^{2}-\frac{1}{2} k \eta_{1}^{2}-2 k\left(\eta_{2}-\eta_{1}\right)^{2}-14 k \eta_{2}^{2} . \tag{43}
\end{equation*}
$$

You are not required to find the equilibrium values of $x_{1}$ and $x_{2}$. However, suppose all the unstretched spring lengths are $a$ and the total distance between the walls is $L$. Then, with $x_{1,2}$ being the location of the masses relative to the left wall, we have

$$
\begin{equation*}
U=\frac{1}{2} k_{1}\left(x_{1}-a\right)^{2}+\frac{1}{2} k_{2}\left(x_{2}-x_{1}-a\right)^{2}+\frac{1}{2} k_{3}\left(L-x_{2}-a\right)^{2} . \tag{44}
\end{equation*}
$$

Differentiating with respect to $x_{1,2}$ then yields

$$
\begin{align*}
& \frac{\partial U}{\partial x_{1}}=k_{1}\left(x_{1}-a\right)-k_{2}\left(x_{2}-x_{1}-a\right)  \tag{45}\\
& \frac{\partial U}{\partial x_{2}}=k_{2}\left(x_{2}-x_{1}-a\right)-k_{3}\left(L-x_{2}-a\right) \tag{46}
\end{align*}
$$

Setting these both to zero, we obtain

$$
\begin{align*}
\left(k_{1}+k_{2}\right) x_{1}-k_{2} x_{2} & =\left(k_{1}-k_{2}\right) a  \tag{47}\\
-k_{2} x_{1}+\left(k_{2}+k_{3}\right) x_{2} & =\left(k_{2}-k_{3}\right) a+k_{3} L . \tag{48}
\end{align*}
$$

Solving these two inhomogeneous coupled linear equations for $x_{1,2}$ then yields the equilibrium positions. However, we don't need to do this to solve the problem.
(b) Find the T and V matrices.
[5 points]
We have

$$
\mathrm{T}_{\sigma \sigma^{\prime}}=\frac{\partial^{2} T}{\partial \dot{\eta}_{\sigma} \partial \dot{\eta}_{\sigma^{\prime}}}=\left(\begin{array}{cc}
m & 0  \tag{49}\\
0 & 4 m
\end{array}\right)
$$

and

$$
\mathrm{V}_{\sigma \sigma^{\prime}}=\frac{\partial^{2} U}{\partial \eta_{\sigma} \partial \eta_{\sigma^{\prime}}}=\left(\begin{array}{cc}
5 k & -4 k  \tag{50}\\
-4 k & 32 k
\end{array}\right) .
$$

(c) Find the eigenfrequencies $\omega_{1}$ and $\omega_{2}$.
[5 points]
We have

$$
\begin{align*}
\mathrm{Q}(\omega) \equiv \omega^{2} \mathrm{~T}-\mathrm{V} & =\left(\begin{array}{cc}
m \omega^{2}-5 k & 4 k \\
4 k & 4 m \omega^{2}-32 k
\end{array}\right) \\
& =k\left(\begin{array}{cc}
\lambda-5 & 4 \\
4 & 4 \lambda-32
\end{array}\right), \tag{51}
\end{align*}
$$

where $\lambda=\omega^{2} / \omega_{0}^{2}$, with $\omega_{0}=\sqrt{k / m}$. Setting $\operatorname{det} Q(\omega)=0$ then yields

$$
\begin{equation*}
\lambda^{2}-13 \lambda+36=0, \tag{52}
\end{equation*}
$$

the roots of which are $\lambda_{-}=4$ and $\lambda_{+}=9$. Thus, the eigenfrequencies are

$$
\begin{equation*}
\omega_{-}=2 \omega_{0} \quad, \quad \omega_{+}=3 \omega_{0} \tag{53}
\end{equation*}
$$

(d) Find the modal matrix $\mathrm{A}_{\sigma i}$.
[5 points]
To find the normal modes, we set

$$
\left(\begin{array}{cc}
\lambda_{ \pm}-5 & 4  \tag{54}\\
4 & 4 \lambda_{ \pm}-32
\end{array}\right)\binom{\psi_{1}^{( \pm)}}{\psi_{2}^{( \pm)}}=0 .
$$

This yields two linearly dependent equations, from which we can determine only the ratios $\psi_{2}^{( \pm)} / \psi_{1}^{( \pm)}$. Plugging in for $\lambda_{ \pm}$, we find

$$
\begin{equation*}
\binom{\psi_{1}^{(-)}}{\psi_{2}^{(-)}}=\mathcal{C}_{-}\binom{4}{1} \quad, \quad\binom{\psi_{1}^{(+)}}{\psi_{2}^{(+)}}=\mathcal{C}_{+}\binom{1}{-1} . \tag{55}
\end{equation*}
$$

We then normalize by demanding $\psi_{\sigma}^{(i)} \mathrm{T}_{\sigma \sigma^{\prime}} \psi_{\sigma^{\prime}}^{(j)}=\delta_{i j}$. We can practically solve this by inspection:

$$
\begin{equation*}
20 m\left|\mathcal{C}_{-}\right|^{2}=1 \quad, \quad 5 m\left|\mathcal{C}_{+}\right|^{2}=1 \tag{56}
\end{equation*}
$$

We may now write the modal matrix,

$$
A=\frac{1}{\sqrt{5 m}}\left(\begin{array}{cc}
2 & 1  \tag{57}\\
\frac{1}{2} & -1
\end{array}\right)
$$

(e) Write down the most general solution for the motion of the system.
[5 points]
The most general solution is

$$
\begin{equation*}
\binom{\eta_{1}(t)}{\eta_{2}(t)}=B_{-}\binom{4}{1} \cos \left(2 \omega_{0} t+\varphi_{-}\right)+B_{+}\binom{1}{-1} \cos \left(3 \omega_{0} t+\varphi_{+}\right) . \tag{58}
\end{equation*}
$$

Note that there are four constants of integration: $B_{ \pm}$and $\varphi_{ \pm}$.

