

**1** Expansion laws

Draw a  $\sim 1''$  radius disk on a piece of paper. Imagine that the disk is composed of a very large number of points. Consider two different versions of Hubble's Law which hold for a time,  $t$ :

- Version 1: Constant velocity. All particles separate from each other at the same rate, i.e.,  $v = a = \text{constant}$
- Version 2: Hubble's law. Particles move according to  $v = \dot{a}r$ , where  $r$  is the separation between them.

Sketch the initial distribution of particles and the final distribution after time  $t$ , from the point of view of an observer at the center (using the cosmological principle this is as good a location as any).

**2** Relativistic Redshift

The velocity-redshift law discussed in class is not correct at velocities approaching the speed of light,  $c$ . The full relativistic formula is:

$$1 + z = \sqrt{\frac{1 + v/c}{1 - v/c}}$$

Expand this equation in powers of  $v/c$  to obtain the non-relativistic formula used in class.

**3** Planck function

(a) Write down and sketch both the Planck function for the occupation number (number of photons per spatial mode) and the energy density in a frequency interval  $df$  as a function of blackbody temperature  $T$  and frequency  $f$ .

(b) Is there a peak in the occupation number? In one sentence, why does the energy density peak?

(c) Expand the formula for the energy density when  $h\nu/k_bT \ll 1$ .

**4** Ryden HW Problem 2.2**5** Ryden HW Problem 2.3**6** Ryden HW Problem 2.4