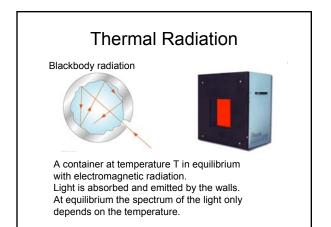
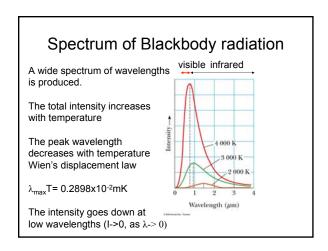
6.1 Quantum Physics. Particle Nature of Light

 Particle nature of Light Blackbody Radiation Photoelectric Effect
Properties of photons

- lonizing radiation Radiation damage
- x-rays Compton effect X-ray diffraction

Photons When light exchanges energy with atoms it behaves as a particle - called the photon The energy of a photon is proportional to the frequency f of light $E_{photon} = hf$ Planck's Constant h=6.626x10⁻³⁴ J·s

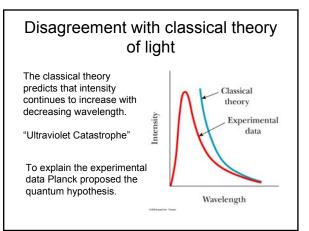




Demonstration of blackbody radiation

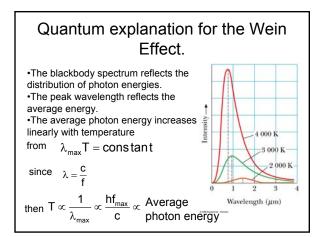
A tungsten filament light bulb is approximately a black body radiator.

http://www.physics.ucla.edu/demoweb/demomanual/astronom y/quantum_mechanics/blackbody_radiation.html

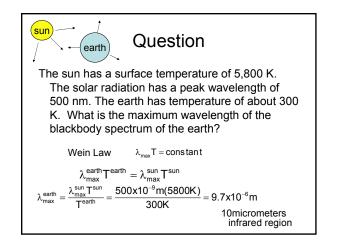


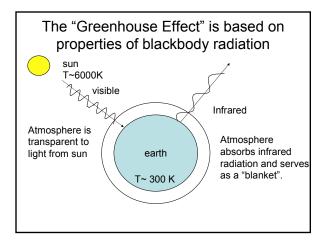
Planck's constant Classical and Planck picture Suppose we have a box that contains light waves Planck proposed that light with different wavelengths. could only have certain energies The energy is contained in energy states containing E=hf particles with different energies Then the energy of oscillators Classical theory predicts that the number of energy states in the black body could only have increased with decreasing wavelength. "smaller particles certain fixed values are more numerous' Planck proposed that in addition the short wavelength particles are more "energetically expensive"

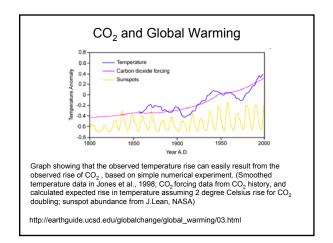
So at short wavelength, they would be hard to produce. This explains the peak in the black body spectrum

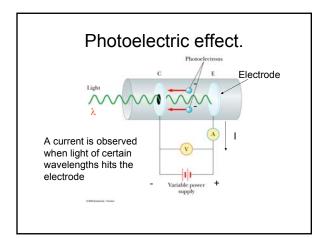


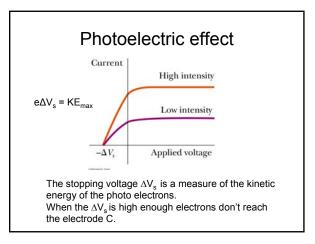
Max Planck

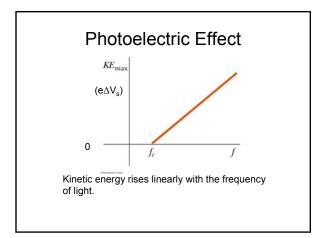


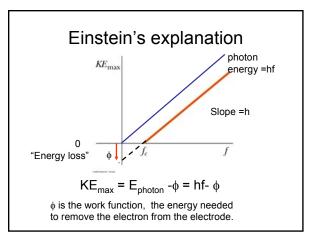


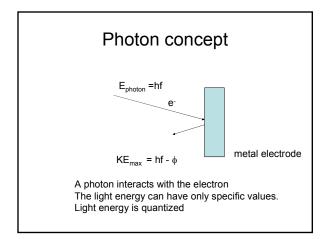


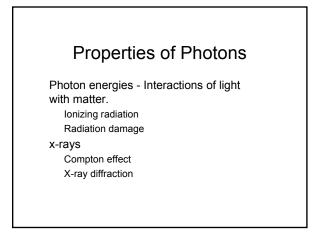




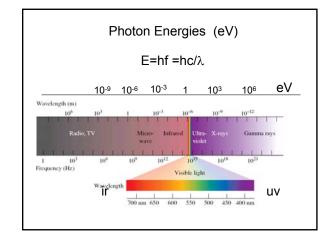


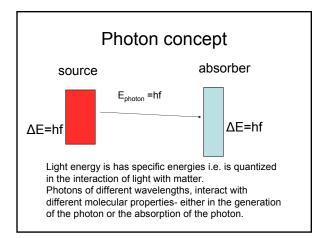






Photon EnergyFind the energy of a photon with a wavelength of500 nm. Use units of electron volts
$$(1eV=1.60x10^{-19} J)$$
 $E = hf = h \frac{c}{\lambda} = 6.63x10^{-34} Js \left(\frac{3x10^8 m/s}{500x10^{-9} m} \right) = 4.0x10^{-19} J$ $E = \frac{4x10^{-19} J}{1.6x10^{-19} J/eV} = 2.5eV$ An electron volt is the energy change in moving
an electron across a potential of 1 volt.A few electron volts is the energy of electrons in
molecules. This is why visible light is absorbed by
molecules (pigment molecules).





		<u>Photon Er</u>	101 9100	
light	Typical Wavelength (m)	Typical Photon energy (eV)	Molecular interactions	applications
radio	10	10-7 eV	nuclear magnetic	NMR imaging
microwaves	10-2	10 ⁻⁴ eV	Molecular rotations	Microwave oven cell phone
Infrared	10-5	10 ⁻¹ eV	Molecular vibrations	Heat lamp
Visible	400-700 nm	2-3 eV	Low energy electrons (pigments)	Vision Photosynthesis Photography
Ultraviolet	200-300 nm	4-5 eV	bonding electrons	Radiation damage Skin cancer
X-rays	1 nm	104 eV	Electrons scattering, tightly bound electrons	X-ray imaging X-ray diffraction

